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THE APPLICATION OF FINITE ELEMENTS AND SPACE-ANGLE SYNTHESIS TO THE ANISOTROPIC STEADY STATE BOLTZMANN (TRANSPORT) EQUATION

THESIS

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THE APPLICATION OF FINITE ELEMENTS AND SPACE-ANGLE SYNTHESIS TO THE ANISOTROPIC STEADY STATE BOLTZMANN (TRANSPORT) EQUATION

#### THESIS

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by

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#### Preface

This research project is a part of the ongoing effort here at the Air Force Institute of Technology to develop alternate methods for solving the two-dimensional steady state anisotropic neutral particle transport equation. The thrust of this research effort is towards an accurate and cost-effective solution of the steady state transport of neutrons, gamma rays and high energy x-rays from a low altitude nuclear burst. This problem which is modeled as a point source in a two-dimensional cylindrical (r,z) geometry with the air ground interface included, is of particular interest in the areas of nuclear weapons effects and radiation physics.

Presently the most widely used computational methods for solving the (air-over-ground) problem are Monte Carlo and discrete ordinates. However, these methods have severe limitations and computational problems. My research plan was to formulate and evaluate a solution technique which did not have these disadvantages. A finite element solution method which is based on a space-angle synthesis flux expansion of bicubic splines and spherical harmonics was chosen. The merits of this solution technique were examined and a computer algorithm for the numerical solution of this problem was developed.

I wish to acknowledge and express my appreciation for the assistance and encouragement which I have received

from the staff and students of the Air Force Institute of Technology. Special thanks are due to my advisor, Captain David D. Hardin, without whose direction, encouragement and many hours of discussion and counselling this thesis would not have been possible. I am also grateful for the support, advice and encouragement that was provided by Dr. J. Jones of the Air Force Institute of Technology Mathematics Department.

Finally, I wish to express my appreciation to my wife and daughter for their understanding, patience and constant support throughout this project. To my wife, Cynthia, I must also express a special thanks for her effort in typing this thesis.

Eze E. Wills

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#### Abstract

A finite element space-angle synthesis solution of the steady state anisotropic Boltzmann (transport) equation in a two-dimensional cylindrical geometry has been developed. Starting from a variational principle the Bubnov-Galerkin solution method was applied to the second order even parity form of the Boltzmann equation. A trial function flux expansion in bicubic splines and spherical (surface) harmonics was used. A first scatter (collision) source and an exponentially varying atmosphere was also incorporated into this development.

Finite element space-angle synthesis (FESAS) was developed as an alternate solution approach and an improvement in comparison to the methods of Monte Carlo and discrete ordinates. FESAS does not have the inherent characteristics which have produced the ray effect problem in discrete ordinates. Also, FESAS may result in lower computational costs than those of Monte Carlo and discrete ordinates.

The second order even parity form of the Boltzmann equation was derived and shown to be symmetric, positive definite and self-adjoint. The equivalence of a variational minimization principle and the Bubnov-Galerkin method of weighted residuals was established. The finite element space angle synthesis system of equations was

expanded and a numerical computer solution approach was implemented. A computer program was written to solve for the trial function expansion (mixing) coefficients, and also to compute the particle flux.

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#### I Introduction

#### Background

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The Air-Over-Ground Problem. The transport of neutrons, gamma rays and high-energy x-rays, away from a low altitude nuclear explosion (air-burst), is of special interest in assessing the vulnerability and survivability of military weapon systems and in making radiation exposure and dose predictions. This neutral particle transport problem increases in complexity because of the exponentially varying air density and the air-ground interface. A description of neutral particle transport and, therefore, the air-over-ground problem is given by the Boltzmann transport equation.

Numerical solutions to this problem already exist.

The main solution techniques are Monte Carlo and discrete ordinates. However, discrete ordinates and Monte Carlo have severe difficulties and disadvantages. To perform an accurate the problem valentation requires the use of hours of costly disappear time. Sincrete ordinates uses less computer execution time than Monte Carlo, however, it is subject to a computational difficulty called ray affects (Ref. 1:357).

Ray Effects and Discrete Ordinates. Ray effects are a result of the angular discretization of the particle flux in the discrete ordinate method. It is not a numerical problem, but originates in the derivation of the discrete ordinate  $S_n$  equations. In a physical sense these equations only allow source particles to travel in specific directions. However, in most practical problems these particles move in all directions. An in-depth analysis of the  $S_n$  equations and the nature and reasons for ray effects can be found elsewhere (Ref 1:357).

Ray effects produce non-physical distortions of the angular flux in regions where there are strong absorbers, localized sources, or high energy streaming particles (Ref 2:255-268). These distortions in the numerical formulation of the discrete ordinate method produce solutions which are inaccurate. The degree of these inaccuracies is dependent upon the specific problem and the nature of the absorbing media and sources. In the air-over-ground problem this effect will be significant because it is essentially a streaming particle problem with localized first scatter sources.

A considerable amount of work has already been done in an attempt to eliminate ray effects from the  $\mathbf{S}_n$  equations and the discrete ordinate method. Ray effects can be mitimated by the use of a fine angular mesh in the finite differencing scheme of the  $\mathbf{S}_n$  equations. However, this approach increases the computational time and the already

high computer costs. Other approaches involve a spherical harmonic-like formulation (Ref 2:255-268), piecewise bilinear finite element approximations (Ref 3:205-217), and space-angle synthesis with specially tailored trial functions (Ref 4:322-343).

#### Problem, Scope and Solution Approach

The purpose of this research project was to divelop a finite element solution to the air-over-ground problem by using a space-angle synthesis of bicubic splines in space and spherical (surface) harmonics in angle. Specifically, a solution of the monoenergetic Boltzmann equation in the context of the air-over-ground problem was sought. Working from a variational principle and using a judicious choice of trial functions the problem of ray effects may be eliminated (Ref 3:214). Also, this judicious trial function choice, and a Bubnov-Galerkin solution method may be more efficient and less costly than Monte Carlo or discrete ordinates.

The steady state solution of the Boltzmann equation with first scatter sources, anisotropic scattering and an exponentially changing atmosphere is desired. The problem is formulated from a variational principle and in a two-dimensional cylindrical (r,z) spatial geometry with the air-over-ground interface included. Fluence values as a function of two spatial (r,z) and two angular (µ,x) variables are sought. This is a four dimensional problem. Finally, a numerical solution algorithm and the computer implementation of this problem is required.

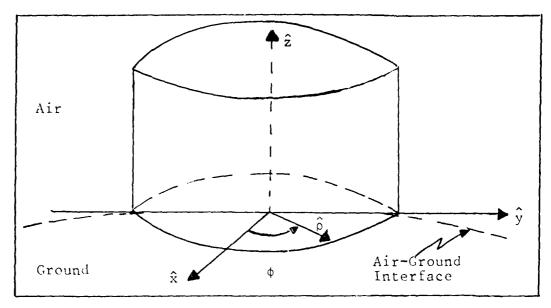


Figure 1. Cylindrical  $(\rho, z, \phi)$  Problem Geometry

#### Assumptions

There are two basic assumptions which are made in the formulation of this problem. A time-independent (steady state) solution and axial symmetry is assumed. Because of the exponentially changing air density the air-over-ground problem is four dimensional with a spatially dependent (r,z) solution. An assumption of axial symmetry is made possible by ignoring the curvature of the earth. Within the problem domain of most practical problems the curvature of the earth is small and can therefore be ignored. Figure 1 shows the spatial cylindrical problem geometry.

The flux from an air burst is non-zero for a fraction of a second (microseconds). Therefore, particle fluence (number/area) and not flux is the more useful quantity.

A steady state formulation of the air-over-ground problem is obtained by integrating the time dependence out of the

Boltzmann equation. This integration which is carried out over time limits when the flux is zero produces a time integrated or fluence equation.

#### <u>Development</u>

In the next chapter of this report the problem equations are presented. A finite element formulation of the air-over-ground problem is developed in Chapter III. The Bubnov-Galerkin method of weighted residuals is incorporated into this development. In Chapter IV a space-angle synthesis of bicubic splines and spherical harmonics is performed. An interpolation of the source terms is also outlined in Chapter IV. A computer implementation of the problem solution is examined in Chapter V. Finally, conclusions and recommendations are presented in Chapter VI.

#### II The Problem Equation

The application of finite elements and a variational principle to the air-over-ground problem and the monoenergetic steady state Boltzmann equation is not a new concept (Ref 5). As in the work by Wheaton (Ref 5) only the monoenergetic problem will be considered. It is assumed that energy dependence can be easily incorporated into this treatment by the use of standard multigroup methods. The air-over-ground problem which is in effect the steady state transport of neutral particles can be described by the Boltzmann (transport) equation and appropriate boundary conditions as follows:

$$\hat{\Omega} \cdot \nabla \phi(\hat{\mathbf{r}}, \hat{\Omega}) + \sigma_{\mathsf{t}}(\hat{\mathbf{r}}) \phi(\hat{\mathbf{r}}, \hat{\Omega}) = \int_{4\pi} \sigma^{\mathsf{s}}(\mathbf{r}, \hat{\Omega} \cdot \hat{\Omega}') \phi(\mathbf{r}, \Omega') d\Omega' + S(\hat{\mathbf{r}}, \hat{\Omega})$$
(1)

This is the one speed monoenergetic Boltzmann equation in general geometry where

 $\hat{r}$  = the spatial position vector,

 $\hat{\Omega}$  = a unit direction or velocity vector,

V = gradient operator,

 $\sigma_t(\hat{r})$  = total macroscopic interaction cross section at spatial position  $\hat{r}$ ,

 $\sigma^{s}(\hat{r}, \hat{\beta}, \hat{\beta})$  = macroscopic scattering cross section. The probability of a particle at position  $\hat{r}$  and direction  $\hat{\beta}$ , scattering into direction  $\hat{\beta}$ . It is a function of the scattering angle  $\hat{\gamma} \cdot \hat{\alpha}$  and not a function of the individual directions (isotropic media),

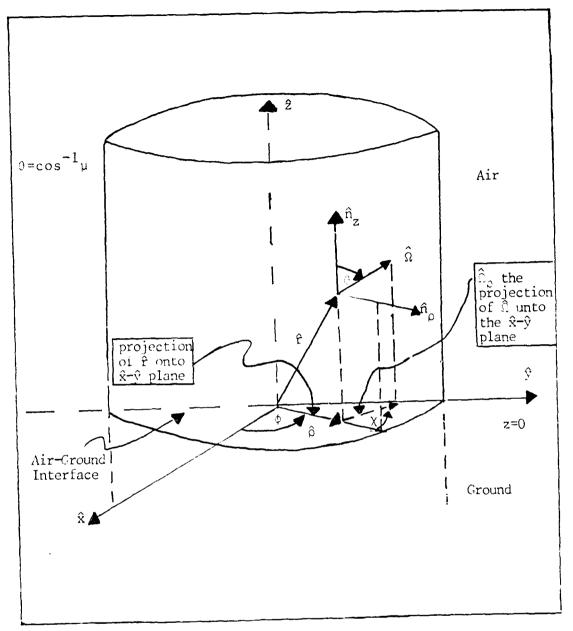


Figure 2. Problem Coometry and Coordinate System.

 $S(r, \cdot)$  = Source a nsity in marricles/unit volume. The integration is carried out over all directions (4% stera-dians).

The air-over-ground problem will be formulated in cylindrical (r,z,) grometry. The coordinate system and directional angles of this content are shown in Figure 2

(Ref 6:57).  $\mu$  is the cosine of the angle formed by the z-axis and the particle velocity vector  $\hat{\Omega}$ .  $\chi$  is the angle between the planes formed by the  $\hat{r}$  vector and the z-axis and that of the  $\hat{\Omega}$  vector and z-axis.

The scattering properties of air show a directional dependence which is highly peaked in the forward direction especially at the high particle energies that exist in the air-over-ground problem. Because of this the exterior boundary condition for this problem will be approximated by a vacuum boundary condition:

$$(\hat{r}_s, \hat{l}) = 0$$
 for  $\hat{r}_s$  on the boundary of the problem domain and  $\hat{l} \cdot \hat{n} < 0$  (2)

where first in outward unit normal to the boundary surface. In physical terms this is a non-reentrant boundary condition. No particles are allowed to reenter the region once they leave it.

In two-dimensional cylindrical (r,z) geometry there is symmetry in the angle . This symmetry can be written as

$$\phi(\hat{r}, \hat{\beta}_1) = \pm (\hat{r}, \hat{\beta}_2) \text{ for } \hat{\beta}_{p2} = \hat{\beta}_{p1} - 2(\hat{n} \cdot \hat{\beta}_{p1})\hat{n}$$
 (3a)

or as

$$z(\hat{x}, y, \zeta) = (\hat{y}, y, -\zeta) \tag{35}$$

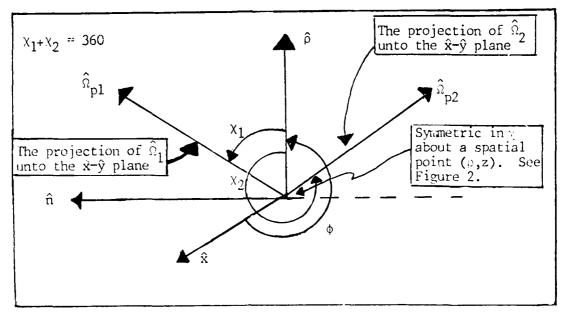


Figure 3. A Schematic of the Angular Symmetry.

This symmetry in  $\chi$  is shown in Figure 3, where the vectors  $\hat{n}$  and  $\hat{n}$  are perpendicular.

Note that because of the exponentially varying air density (in the z direction), only azimuthal symmetry in the angle  $\chi$  is assured. There is no symmetry in  $\mu(\cos\,\theta)$  and therefore  $\chi(\hat{r},\hat{s})$  will not be equal to  $\chi(\hat{r},-\hat{s})$ . The symmetry condition, Eqs (3a) and (3b), implies that

$$\mathfrak{z}(\hat{\mathbf{r}}, \mu, \chi) = \text{even function in } \chi$$
 (3c)

At the air around interface  $\psi(\hat{r}, \hat{\beta})$  is continuous except at z=0 (Ref 6:160), i.e.

$$\gamma(\mathbf{r}, \dot{\gamma}) = \gamma(\mathbf{r}, \dot{\gamma}) \text{ at } z = 0 \text{ and } \mu \neq 0$$

$$\text{air} \qquad \text{around} \qquad (4)$$

Mowever the derivatives of (r. ) are not continuous.

The problem geometry and coordinate systems as shown in Figure 2 implies that when  $\rho$  is equal to zero the angle  $\chi$  must also assume a value of zero. Therefore, along the z axis ( $\rho$ =0) the angle  $\chi$  does not vary between 0 and  $2\pi$ . This means that there is no  $\chi$  variation in  $\varphi(\hat{\mathbf{r}},\hat{\Omega})$  at  $\rho$ =0 and that

$$\frac{\partial}{\partial \chi} \phi(\hat{\mathbf{r}}, \hat{\Omega}) = 0 \text{ for } \hat{\Omega} = 0$$
 (5)

#### III The Finite Element Method

The finite element method is a mathematical and numerical technique for approximating the solution to a large class of problems. Initially it was developed and used to solve problems in stress analysis (Ref 7:9). Later, as the mathematical foundation of the method was established it gained widespread acceptance and use in solving a larger class of problems.

Finite elements are an extension of the Rayleigh-Ritz technique of first recasting the problem in an equivalent variational form and then seeking a solution on the basis of an energy minimization principle (Ref 8:1). In the Rayleigh-Ritz method a solution in the form of a linearly independent set of trial functions is assumed. These trial functions must satisfy the essential boundary conditions. The approximate or "best" solution to the problem is the linear combination of these trial functions which maximizes (or minimizes) the variational principle (functional). If this linear combination of functions is not an extremum (maximum or minimum) of the functional, then the class (or space) of trial functions is expanded by the addition of more functions. This expansion of the trial function space is continued until a linear combination of functions which is an extremum of the functional is obtained.

The finite element solution technique is similar to that of Rayleigh-Ritz. The only difference lies in the choice of trial functions. In finite elements the problem domain is divided into smaller regions (grids) or elements. Each trial function is usually associated with only a few elements. Unlike Rayleigh-Ritz, finite elements uses trial functions which are zero over parts of the solution domain (a local basis). Also, the trial space (number of trial functions) is expanded by using more elements (mesh points) and not by the addition of a new class of functions. Because of these differences the finite element method is more adaptable towards a numerical (computer) solution than Rayleigh-Ritz.

There are two basic approaches to the finite element formulation of a problem. One approach is to find the extremum of the functional which originates from a variational principle and the calculus of variations (Rayleigh-Ritz with a local basis). The other is by the method of weighted residuals. The method of weighted residuals does not include the use of a variational principle or the calculus of variations. In some problems a variational principle has not been developed or may not exist, and therefore, the Rayleigh-Ritz approach cannot be used. However, in such cases, the method of weighted residuals can be used to solve the problem. Therefore, the method of weighted residuals can be extended to a wider class of problems than Rayleigh-Ritz.

The method of weighted residuals is another approach for developing a set of (algebraic) problem equations to which the finite element method can be applied. There are

three basic mathematical "recipes" through which the method of weighted residuals can be developed. These are the methods of least squares, collocation and Galerkin. In some problems where a variational principle (functional) exists it can be shown (Ref 9:735) that the Galerkin method of weighted residuals is equivalent to Rayleigh-Ritz. An identical set of matrix equations and therefore the same solution is achieved by either method.

In the following paragraphs a variational principle for the air-over-ground problem and the even parity form of the Boltzmann equation is examined. A weak form of the variational principle and the boundary conditions are incorporated within this development. Finally, the Galerkin method of weighted residuals is discussed and an equivalence to the variational approach for the air-over-ground problem is established.

#### Even and Odd Parity Second Order Forms

In order that a variational principle may be used the even and odd parity forms of the anisotropic Boltzmann equation and associated boundary conditions will be developed. The starting point of this development is Eqs (1) and (2). Following the derivation of Kaplan and Davis (Ref 10:166) and that of Wheaton (Ref 5) the monomenergetic steady state transport equation can be written in terms of the  $-\hat{\Sigma}$  vector by changing  $\hat{\omega}$  to  $-\hat{\omega}$  in Eq (1). Thus gives

$$-\hat{\theta} \cdot \nabla : (\hat{\mathbf{r}}, -\hat{\theta}) + \exists \mathbf{t}(\hat{\mathbf{r}}) \cdot \mathbf{t}(\hat{\mathbf{r}}, -\hat{\theta}) = \int_{\mathbb{R}^3} \mathbf{t}^{\mathbf{s}}(\hat{\mathbf{r}}, -\hat{\theta} \cdot \hat{\theta}) \cdot (\hat{\mathbf{r}}, \hat{\theta}) \cdot d\hat{\theta}$$

$$+ \exists (\mathbf{r}, -\mathbf{t})$$
(6)

The even and odd parity terms will now be defined as

$$\Psi(\hat{\mathbf{r}}, \hat{\Omega}) = \frac{1}{2} \{ \phi(\hat{\mathbf{r}}, \hat{\Omega}) + \phi(\hat{\mathbf{r}}, -\hat{\Omega}) \}$$
 (7)

$$\chi(\hat{\mathbf{r}}, \hat{\Omega}) = \frac{1}{2} \{ \phi(\hat{\mathbf{r}}, \hat{\Omega}) - \phi(\hat{\mathbf{r}}, -\hat{\Omega}) \}$$
 (8)

$$S^{g}(\hat{r},\hat{\Omega}) = \frac{1}{2} \{S(\hat{r},\hat{\Omega}) + S(\hat{r},-\hat{\Omega})\}$$
(9)

$$S^{U}(\hat{r},\hat{\Omega}) = \frac{1}{2} \{S(\hat{r},\hat{\Omega}) - S(\hat{r},-\hat{\Omega})\}$$
 (10)

$$\sigma^{\mathbf{g}}(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}) = \frac{1}{2} \{ \sigma^{\mathbf{g}}(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}') + \sigma^{\mathbf{g}}(\hat{\mathbf{r}}, -\hat{\Omega} \cdot \hat{\Omega}') \}$$
(11)

$$\sigma^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}') = \frac{1}{2} \{ \sigma^{\mathbf{s}}(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}') - \sigma^{\mathbf{s}}(\hat{\mathbf{r}}, -\hat{\Omega} \cdot \hat{\Omega}') \}$$
 (12)

where

 $\Psi(\hat{r}, \hat{\Omega})$  = even parity fluence

 $\chi(\hat{\mathbf{r}}, \hat{\Omega}) = \text{odd parity fluence}$ 

 $S^g(\hat{r}, \hat{\Omega})$  = even parity source

 $S^{u}(\hat{r}, \hat{\Omega}) = \text{odd parity source}$ 

 $\sigma^{g}(\hat{\mathbf{r}}, \hat{\mathbf{n}} \cdot \hat{\mathbf{n}}) = \text{even parity scattering cross-section}$ 

 $\sigma^{u}(\hat{r}, \hat{\Omega} \cdot \hat{k}) = \text{odd parity scattering cross-section}$ 

Adding Eqs (1) and (6), then dividing throughout by two and using the above definitions gives

$$\hat{\Omega} \cdot \nabla \chi(\hat{r}, \hat{\Omega}) + \sigma_{\mathsf{t}}(\hat{r}) \Psi(\hat{r}, \hat{\Omega}) = \int_{\mathbb{R}^{3}} \sigma^{\mathsf{g}}(\hat{r}, \hat{\Omega} \cdot \hat{\Omega}') \phi(\hat{r}, \hat{\Omega}') d\Omega' + S^{\mathsf{g}}(\hat{r}, \hat{\Omega})$$

$$(13)$$

Using the derivation of Wheaton (Ref 5:8), which is also reproduced in Appendix A, the scattering kernel term in (13) can be written as

$$\int_{4\pi} r^{\lambda}(\mathbf{r}, \mathbf{d} \cdot \hat{\mathbf{d}}) \gamma(\mathbf{r}, \lambda') d\hat{\mathbf{d}} = \int_{4\pi} r^{\lambda}(\hat{\mathbf{r}}, \hat{\mathbf{d}} \cdot \mathbf{d}) \gamma(\hat{\mathbf{r}}, \hat{\mathbf{d}}) d\hat{\mathbf{d}} \qquad (14)$$

where the even properties of the even parity scattering cross-section and the even parity fluence have been used in the derivation of Eq (14). Eq (13) now becomes

$$\hat{\Omega} \cdot \nabla \chi(\hat{\mathbf{r}}, \hat{\Omega}) + \sigma_{\mathbf{t}}(\hat{\mathbf{r}}) \Psi(\hat{\mathbf{r}}, \hat{\Omega}) = \int_{\Psi_{\mathbf{f}}} \sigma^{\mathbf{g}}(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}) \Psi(\hat{\mathbf{r}}, \hat{\Omega}') d\hat{\Omega}' + S^{\mathbf{g}}(\hat{\mathbf{r}}, \hat{\Omega})$$

$$(15)$$

Similarly, by substracting Eqs (1) and (6) and rearranging the scattering kernel (See Appendix A) gives

$$\hat{\mathbb{G}} \cdot \nabla \mathbb{P}(\hat{\mathbf{r}}, \hat{\mathbb{G}}) + \sigma_{\mathbf{t}}(\hat{\mathbf{r}}) \chi(\hat{\mathbf{r}}, \hat{\mathbb{G}}) = \int_{4\pi} \sigma^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\mathbb{G}} \cdot \hat{\mathbb{G}}') \chi(\hat{\mathbf{r}}, \hat{\mathbb{G}}') d\hat{\mathbb{G}}' + S^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\mathbb{G}})$$

$$+ S^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\mathbb{G}})$$

$$(16)$$

Eqs (15) and (16) are referred to by Kaplan and Davis (Ref 10) as canonical forms. The natural boundary condition Eq (2) can also be rewritten as

$$f(\hat{\mathbf{r}}_s, \hat{\mathbf{n}}) + \chi(\hat{\mathbf{r}}_s, \hat{\mathbf{n}}) = 0 \text{ for } \hat{\mathbf{n}} < 0$$
 (17)

and

$$\Psi(\hat{\mathbf{r}}_{s},\hat{\hat{\mathbf{c}}}) = \chi(\hat{\mathbf{r}}_{s},\hat{\hat{\mathbf{c}}}) = 0 \text{ for } \hat{\hat{\mathbf{c}}} \cdot \hat{\mathbf{n}} > 0$$
 (18)

The symmetric conditions mas (3a) and (3c) now become

$$\mathbb{E}(\hat{\mathbf{r}}, \hat{\mathbf{r}}_{1}) + \chi(\hat{\mathbf{r}}, \hat{\mathbf{r}}_{1}) = \mathbb{E}(\hat{\mathbf{r}}, \hat{\mathbf{r}}_{2}) + \chi(\hat{\mathbf{r}}, \hat{\mathbf{r}}_{2})$$

$$\text{for } \hat{\mathbf{r}}_{p2} = \hat{\mathbf{r}}_{p1} - 2(\hat{\mathbf{r}}_{p1} \cdot \hat{\mathbf{n}})\hat{\mathbf{n}}$$
(19)

where  $\gamma(r,1)$  and  $\gamma(r,1)$  must be even functions in the azimuthal angle  $\gamma$  (see Fig. 3). Therefore it follows that for  $\frac{1}{\sqrt{p_1^2-2}}(\frac{1}{\sqrt{p_1^2-2}}(\frac{1}{\sqrt{p_1^2-2}}))^{\frac{n}{n}}$ 

$$z(\hat{x},\hat{y}) \approx z(\hat{x},\hat{y}) \tag{20}$$

and

$$\chi(\hat{\mathbf{r}}, \hat{\Omega}_1) = \chi(\hat{\mathbf{r}}, \hat{\Omega}_2) \tag{21}$$

Also  $\chi(\hat{r}, \hat{\Omega})$  and  $\Psi(\hat{r}, \hat{\Omega})$  are continuous at the air ground interface (when  $\mu \neq 0$ ) but their derivatives are discontinuous.

A further simplification is now introduced into this development by defining the even and odd operators,  $\boldsymbol{G}^g$  and  $\boldsymbol{G}^u$  as

$$G^{S}(\hat{\mathbf{r}})f(\hat{\mathbf{r}},\hat{\Omega}) = \sigma_{t}(\hat{\mathbf{r}})f(\hat{\mathbf{r}},\hat{\Omega}) - \int_{4\pi} \sigma^{g}(\hat{\mathbf{r}},\hat{\Omega}\cdot\hat{\Omega})f(\hat{\mathbf{r}},\hat{\Omega})d\hat{\Omega}$$
 (22)

$$G^{\mathrm{u}}(\hat{\mathbf{r}}) f(\hat{\mathbf{r}}, \hat{\Omega}) = \sigma_{\mathrm{t}}(\hat{\mathbf{r}}) f(\hat{\mathbf{r}}, \hat{\Omega}) - \int_{4\pi} \sigma^{\mathrm{u}}(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}) f(\hat{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}) d\hat{\Omega}^{\mathrm{r}}$$
(23)

where the scattering cross sections can be expanded in spherical (surface) harmonics (see Appendix A) to give

$$G^{g}(\hat{\mathbf{r}})f(\hat{\mathbf{r}},\hat{\Omega}) = \sigma_{t}(\hat{\mathbf{r}})f(\hat{\mathbf{r}},\hat{\Omega})$$
$$-\Sigma_{g}\Sigma_{m}\sigma_{x}^{g}(\hat{\mathbf{r}})Y_{m}(\hat{\Omega})\int_{4\pi}\mathring{Y}_{m}^{f}(\hat{\Omega}')f(\hat{\mathbf{r}},\hat{\Omega}')d\hat{\Omega}' \qquad (24)$$

where the even parity Legendre expansion cross-section  $\sigma_k^g$  is defined as

$$\sigma_{\lambda}^{S}(\hat{\mathbf{r}}) = \begin{cases} \sigma_{\lambda}^{S}(\hat{\mathbf{r}}) & \text{for } \ell \text{ even} \\ 0 & \text{for } \ell \text{ odd} \end{cases}$$
 (25)

and

 $\sigma_{\chi}^{s}(\hat{r})$  = Legendre expansion scattering cross-section coefficients (Ref 5:29)

The derivation of  $\mathcal{J}(\hat{z}) = \mathcal{J}(r)$  can be found in Appendix A.

The odd parity scattering cross section can also be expanded to give

$$G^{\mathbf{u}}(\hat{\mathbf{r}})f(\hat{\mathbf{r}},\hat{\Omega}) = \sigma_{\mathbf{t}}(\hat{\mathbf{r}})f(\hat{\mathbf{r}},\hat{\Omega}) - \sum_{\mathbf{k}} \sum_{m} \sigma_{\mathbf{k}}^{\mathbf{u}}(\hat{\mathbf{r}})Y_{\mathbf{k}m}(\hat{\Omega}) \int_{4\pi} Y_{\mathbf{k}m}^{*}(\hat{\Omega})f(\hat{\mathbf{r}},\hat{\Omega})d\hat{\Omega}$$
 (26)

where only odd expansions in  $\ell$  are used or  $\sigma_0^{\mathsf{u}}$  is defined as

$$\sigma_{k}^{\mathbf{u}}(\hat{\mathbf{r}}) = \begin{cases} \sigma_{k}^{\mathbf{s}}(\mathbf{r}) & \text{for } k \text{ odd} \\ 0 & \text{for } k \text{ even} \end{cases}$$
 (27)

The  $G^g$  and  $G^u$  operators are self adjoint positive definite (Ref 10:174). Inserting these operators into Eqs (15) and (16) they become

$$G^{g}(\hat{r})\Psi(\hat{r},\hat{\Omega}) = S^{g}(\hat{r},\hat{\Omega}) - \hat{\Omega}\cdot\nabla\chi(\hat{r},\hat{\Omega})$$
(28)

and

$$G^{u}(\hat{r})\chi(\hat{r},\hat{\Omega}) = S^{u}(\hat{r},\hat{\Omega}) - \hat{\Omega} \cdot \nabla \Psi(\hat{r},\hat{\Omega})$$
(29)

Solving for  $\Psi(\hat{r}, \hat{\Omega})$  and  $\chi(\hat{r}, \hat{\Omega})$  from Eqs (28) and (29) produces

$$\Psi(\hat{\mathbf{r}}, \hat{\Omega}) = \left[ \mathbf{G}^{g}(\hat{\mathbf{r}}) \right]^{-1} \left\{ \mathbf{S}^{g}(\hat{\mathbf{r}}, \hat{\Omega}) - \hat{\Omega} \cdot \nabla_{X}(\hat{\mathbf{r}}, \hat{\Omega}) \right\}$$
(30)

$$\chi(\hat{\mathbf{r}}, \hat{\hat{\mathbf{x}}}) = \left[ \mathbf{G}^{\mathbf{u}}(\hat{\mathbf{r}}) \right]^{-1} \left\{ \mathbf{S}^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\hat{\mathbf{x}}}) - \hat{\mathbf{u}} \cdot \mathbf{VY}(\hat{\mathbf{r}}, \hat{\hat{\mathbf{x}}}) \right\}$$
(31)

where using the notation of Kaplan and Davis

$$K^{u}(\hat{r}) = \{G^{u}(\hat{r})\}^{-1} = \text{inverse of the operator}$$

$$G^{u}(\hat{r}) \qquad (32)$$

$$K^{\mathfrak{S}}(\hat{\mathbf{r}}) = \{G^{\mathfrak{S}}(\hat{\mathbf{r}})\}^{-1} = \text{inverse of the operator}$$

$$G^{\mathfrak{S}}(\hat{\mathbf{r}}) \qquad (33)$$

A detailed mathematical derivation of these inverse operators is presented in Appendix A. They are defined as (Ref 11:481)

$$K^{g}(\hat{\mathbf{r}}) = \sigma_{t}^{-1}(\hat{\mathbf{r}}) \left[ 1 + \sum_{\ell} \mathbb{E}_{m} \{ \sigma_{\ell}^{g}(\hat{\mathbf{r}}) / (\sigma_{t}(\hat{\mathbf{r}}) - \sigma_{\ell}^{g}(\hat{\mathbf{r}})) \} Y_{\ell m}(\hat{\Omega}) \int_{4\pi} Y_{\ell m}^{k}(\hat{\Omega}) d\hat{\Omega}^{r} \right]$$

$$(34)$$

where  $\phi_{i}^{g}(\hat{r})$  is defined by Eq (25).

$$K^{\mathbf{u}}(\hat{\mathbf{r}}) = \sigma_{\mathbf{t}}^{-1}(\mathbf{r}) \left[ 1 + \sum_{\hat{\boldsymbol{z}}} \sum_{\hat{\boldsymbol{z}}} \{ \sigma_{\hat{\boldsymbol{z}}}^{\mathbf{u}}(\mathbf{r}) / (\sigma_{\mathbf{t}}(\hat{\mathbf{r}}) - \sigma_{\hat{\boldsymbol{z}}}^{\mathbf{u}}(\hat{\mathbf{r}})) \} Y_{\hat{\boldsymbol{z}}m}(\hat{\boldsymbol{\Omega}}) \int_{\mathcal{A}_{m}} Y_{\hat{\boldsymbol{z}}m}^{\hat{\boldsymbol{z}}}(\hat{\boldsymbol{\Omega}}) d\hat{\boldsymbol{\Omega}} \right]$$
(35)

Both  $K^{u}$  and  $K^{g}$  must be self adjoint positive definite since they are the inverses of  $G^{u}$  and  $G^{g}$  which are both positive definite and self adjoint.

The functions  $Y_{\ell m}(\hat{\Omega})$  and  $Y_{\ell m}^{*}(\hat{\Omega})$  are the well known normalized spherical (surface) harmonics (Ref 6:609) which are defined as

$$Y_{\ell m}(\hat{\Omega}) = Y_{\ell m}(\mu, \chi) = \sqrt{\frac{2\ell + 1}{4\pi} \cdot \frac{(\ell - m)!}{(\ell + m)!}} P_{\ell}^{m}(\mu) e^{im\chi}$$
(36)

and  $Y_{\mathbb{Q}\mathfrak{m}}$  ( $\hat{\Omega}$ ) being the complex conjugate of  $Y_{\mathbb{Q}\mathfrak{m}}$  ( $\hat{\Omega}$ ) is defined as

$$Y_{\ell m}^{*}(\hat{\Omega}) = Y_{\ell m}^{*}(\mu, \chi) = \sqrt{\frac{2\ell + 1}{4\pi} \cdot \frac{(\ell - m)!}{(\ell + m)!}} P_{\ell}^{m}(\mu) e^{-im\chi}$$
(37)

where  $\mu = \cos\theta$  and  $P_{\ell}^{m}(\mu)$  is the associated Legendre functions.

Eqs (30 and (31) can now be written as

$$\Psi(\hat{\mathbf{r}}, \hat{\Omega}) = K^{g}(\hat{\mathbf{r}}) \{ S^{g}(\hat{\mathbf{r}}, \hat{\Omega}) - \hat{\Omega} \cdot \nabla \chi(\hat{\mathbf{r}}, \hat{\Omega}) \}$$
(38)

and

$$\chi(\hat{\mathbf{r}}, \hat{\Omega}) = K^{\mathbf{u}}(\hat{\mathbf{r}}) \{ S^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\Omega}) - \hat{\Omega} \cdot \nabla \Psi(\hat{\mathbf{r}}, \hat{\Omega}) \}$$
(39)

Substituting Eq (39) into Eq (28) produces the second order even parity form of the Boltzmann equation

$$-\hat{\Omega} \cdot \nabla K^{\mathbf{u}}(\hat{\mathbf{r}}) \hat{\Omega} \cdot \nabla \Psi(\hat{\mathbf{r}}, \hat{\Omega}) + G^{\mathbf{g}}(\hat{\mathbf{r}}) \Psi(\hat{\mathbf{r}}, \hat{\Omega}) = S^{\mathbf{g}}(\hat{\mathbf{r}}, \hat{\Omega})$$
$$- \hat{\Omega} \cdot \nabla K^{\mathbf{u}}(\hat{\mathbf{r}}) S^{\mathbf{u}}(\hat{\mathbf{r}}, \hat{\Omega}) \tag{40}$$

and inserting Eq (39) into Eqs (17) and (18) gives the appropriate surface boundary conditions for Eq (40) as

$$\Psi(\hat{\mathbf{r}}_{s}, \hat{\Omega}) + K^{\mathbf{u}}(\hat{\mathbf{r}}_{s}) \{ S^{\mathbf{u}}(\hat{\mathbf{r}}_{s}, \hat{\Omega}) - \hat{\Omega} \cdot \mathbf{V} (\hat{\mathbf{r}}_{s}, \hat{\Omega}) \}$$

$$= 0 \text{ for } \hat{\Omega} \cdot \hat{\mathbf{n}} < 0$$
(41)

and

$$\Psi(\hat{\mathbf{r}}_{s}, \hat{\Omega}) - K^{u}(\hat{\mathbf{r}}_{s}) \{ S^{u}(\hat{\mathbf{r}}_{s}, \hat{\Omega}) - \hat{\Omega} \cdot \nabla \Psi(\hat{\mathbf{r}}_{s}, \hat{\Omega}) \}$$

$$= 0 \text{ for } \hat{\Omega} \cdot \hat{\mathbf{n}} > 0$$

$$(42)$$

Eqs (38) and (39) give the second order odd parity form of the Boltzmann equation which is

$$-\hat{\Omega} \cdot \nabla K^{g}(\hat{r}) \hat{\Omega} \cdot \nabla \chi(\hat{r}, \hat{\Omega}) + G^{u}(\hat{r}) \chi(\hat{r}, \hat{\Omega}) = S^{u}(\hat{r}, \hat{\Omega})$$
$$-\hat{\Omega} \cdot \nabla K^{g}(\hat{r}) S^{g}(\hat{r}, \hat{\Omega})$$
(43)

Inserting Eq (38) into Eqs (17) and (18) the surface boundary conditions for Eq (43) are

$$\chi(\hat{r}_{s}, \hat{z}) + K^{g}(\hat{r})\{S^{g}(\hat{r}_{s}, \hat{z}) - \hat{z} \cdot \nabla \chi(\hat{r}_{s}, \hat{z})\}$$

$$= 0 \text{ for } \hat{z} \cdot \hat{n} < 0$$
(44)

$$\{(\hat{\mathbf{r}}_{s}, \hat{\mathbb{Q}}) = K^{g}(\hat{\mathbf{r}}) \{S^{g}(\hat{\mathbf{r}}_{s}, \hat{\mathbb{Q}}) = \hat{\mathbb{Q}} \cdot \forall \chi(\hat{\mathbf{r}}_{s}, \hat{\mathbb{Q}})\}$$

$$= 0 \text{ for } \hat{\mathbb{Q}} \cdot \mathbf{n} > 0$$
 (45)

#### The Variational Principle

Variational principles for the monoenergetic transport equation have been found. These principles are related by a series of transformations (Ref 10:166). The functional whose Euler equations are the even and odd parity second order Boltzmann equations will be used in this work. Primarily the even parity component will be used because it is always positive, self-adjoint and can be integrated to give the scalar flux or fluence (Ref 12:148). The odd-parity flux which can be negative integrates (over all directions) into the net current (Ref 13:12). Another "nice" feature of this even and odd parity formulation is that it produces a solution matrix which is positive definite and symmetric.

Defining the inner product of two functions as

$$\langle \mathbf{f}, \mathbf{g} \rangle = \int_{4\pi} \mathbf{f}'(\hat{\Omega}) \mathbf{g}(\hat{\Omega}) d\hat{\Omega}$$
 (46)

where \* means the complex conjugate, the functional which corresponds to Eqs (40) (41) and (42) is given as (Ref 10:169)

$$F(\mathbf{u}) = \int_{\mathbb{R}} \{\langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} (\hat{\Omega} \cdot \nabla \mathbf{u}) \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle - 2 \langle \hat{\Omega} \cdot \nabla \mathbf{u}, \mathbf{K}^{\mathbf{u}} \mathbf{S}^{\mathbf{u}} \rangle + \langle \mathbf{u}, \mathbf{G}^{\mathbf{g}} \mathbf{u} \rangle$$

where  $\oint_{s}$  represents a surface integral.

It can be shown Ref (10:169) that the Euler equation (stationary point) of this functional is indeed the even parity second order form of the Boltzmann equation and that Eqs (41) and (42) are the natural boundary conditions. A similar functional for the odd parity equation has also been found.

In Appendix B the stationary point of Eq (47) is shown to be a minimum and the weak or Galerkin form is obtained. This weak form is

$$\int_{\mathbb{R}^{d}} \{(\hat{\Omega} \cdot \nabla \eta, K^{\mathbf{u}}(\hat{\Omega} \cdot \nabla \Psi) > + \langle \eta, \mathcal{G}^{g} \Psi \rangle \} d\hat{\mathbf{r}} + \oint_{S} \int_{4\pi} |\hat{\Omega} \cdot \hat{\mathbf{n}}| \, \eta \Psi d\hat{\Omega} d\hat{\mathbf{s}}$$

$$= \int_{\mathbb{R}^{d}} \{\langle \hat{\Omega} \cdot \nabla \eta, K^{\mathbf{u}} S^{\mathbf{u}} \rangle + \langle \eta, S^{g} \rangle \} d\hat{\mathbf{r}} \qquad (48)$$

where

 $\eta$  = test or weight function and

 $\Psi$  = trial function

The natural boundary conditions Eqs (41) and (42) are incorporated into the weak form of equation (48). However, the boundary condition at the ground interface is an essential condition.

A solution to the air-over-ground problem can be obtained from a solution to Eq (48) and this essential boundary condition.

The Galerkin or weak form of Eq (48) produces a solution matrix which is positive definite and symmetric. It is symmetric because the test and trial functions are the same in the Galerkin solution method. The matrix is positive definite because  $K^u$  and  $G^g$  are positive definite and it is obvious that for the Galerkin solution the term  $\langle \hat{\Omega} \cdot \nabla r, K^u(\hat{\Omega} \cdot \nabla r) \rangle$  is also positive definite.

#### The Method of Weighted Residuals

The method of weighted residuals is a straightforward and simple prescription for solving a wide class of problems.

Unlike Rayleigh-Ritz it does not depend on a variational principle or the calculus of variations. However, in solving certain types of problems the use of a variational principle or the Galerkin method of weighted residuals is equivalent and they produce solutions which are identical. For the air-over-ground problem and the second order even parity form of the Boltzmann equation it will be shown that the Galerkin method and the weak form, which is given by Eq (48), are equivalent formulations of the same problem.

In the method of weighted residuals an approximate solution which is a linear combination of trial functions is assumed to exist. These trial functions are required to satisfy the necessary boundary and continuity conditions. The approximate solution, when inserted into the problem equation, is then required to be an exact solution of the problem with respect to several weight functions (Ref 7:106). The choice of weight functions determines whether the method of weighted residuals is one of collocation, least squares or Galerkin. In the Galerkin method the weight functions are chosen to be the same as the trial functions.

Applying the Galerkin method of weighted residuals to the second order form of the Boltzmann equation is equivalent to using a variational principle. In Appendix C the weak form, Eq (48), is derived from a Galerkin formulation of this problem. The natural boundary condition is incorporated into this development and an equation identical to Eq (48)

is produced. Therefore, solutions to the second order form of the Boltzmann equation, by using a variational principle or the Galerkin method of weighted residuals are equivalent. This equivalency exists because the second order even parity operator of Eq (40) is positive definite and self-adjoint.

# IV Space-Angle Synthesis of the Even-Parity Anisotropic Boltzmann Equation

A space angle synthesis solution approach has already been applied to the air-over-ground problem. Roberds and Bridgman used "specially tailored" angular trial functions to solve the two dimensional steady state anisotropic Boltzmann equation Ref (4:332). Space angle synthesis was applied directly to Eq (1), and not to the second order form of the Boltzmann equation. Miller et al. (Ref 13:12) have applied phase-space finite elements directly to the isotropic second order Boltzmann equation in x - y geometry. Wheaton (Ref 5) has applied phase-space finite elements to the air-over-ground problem. However a space angle synthesis finite element approach using a flux expansion in bicubic splines and spherical harmonics has not been done.

Because of the complexity of the air-over-ground problem a space-angle synthesis finite element approach seems to be justified. This solution technique might have several advantages, some of which are:

- 1. The elimination of ray effects:
- 2. The numerical advantages of finite elements in combination with a space-angle synthesis approach, may be able to better handle the four-dimensionality of the problem;
- 3. Anisotropic scattering can be easily handled by a "wise and convenient" choice of angular trial functions;

- 4. The computational effort might be reduced, without a loss in accuracy. It is expected that bicubic splines will not require a fine spatial problem grid (mesh size); and
- 5. The boundary conditions and a first scatter source formulation can be easily handled.

The finite element space-angle synthesis technique is merely a spatial and angular expansion of the even parity flux by a tensor product of polynomial functions and spherical harmonics. In this work a tensor product of bicubic polynomial splines is used. This expansion is the trial solution which will be used in the finite element method. The piecewise bicubic spline expansion becomes a local basis in the spatial (p,z) variable. However, the spherical (surface) harmonics which are defined throughout the angular problem domain form a global basis. Therefore, this trial function expansion has a dual basis -- a local basis in space and a global basis in angle. This is the finite element space angle synthesis method.

The starting point of this development is the weak form of the second order even parity Boltzmann equation and essential boundary conditions. The air-over-ground problem is described by the weak form of Eq (48) and the symmetric condition Eq (20). An essential boundary condition, at the air ground interface, is that  $\Psi(\hat{\mathbf{r}},\hat{\boldsymbol{x}})$  is continuous at z=0 and  $z\neq 0$  (see Fig. 1). However, the derivatives of  $\Psi(\hat{\mathbf{r}},\hat{\boldsymbol{x}})$  are discontinuous.

In the remainder of this section the spatial and angular trial functions will be examined. A system of coupled equations for the numerical solution of the air-over-ground problem will be developed. A first scatter (collision) source will be used in this development.

# The Trial Functions

Because this is a four dimensional problem with anisotropic scattering a numerical solution technique is required. The finite element formulation of this problem lends itself directly to such a solution approach. However an application of four dimensional phase-space finite elements to Eq (48) will be very costly (computer costs) and inefficient (Ref 5:33). This is due to the added complexity of anisotropic scattering. Anisotropic scattering increases the bookkeeping and computational difficulties. A local elemental basis in angle requires that the scattering contribution to each element must be computed on an element by element basis for all space and angle elements within the problem domain. Therefore a four-dimensional phase-space finite element formulation of this problem is not a very attractive or realistic approach.

A close examination of Eq (48) shows that the problem operator is self-adjoint positive definite symmetric. This allows the use of standard matrix iterative solution techniques such as Gauss-Seidel, Jacobi or Successive over-relaxation (Ref 14:183). Therefore, a numerical method that includes a finite element solution might be feasible

if the anisotropic scattering contributions can be effectively dealt with. A space-angle synthesis finite element development with a local spline basis in space and a global spherical harmonic basis in angle appears to meet this requirement.

A phase-space finite element problem formulation which eliminated the characteristic lines of the hyperbolic discrete ordinate  $S_n$  equations has been effective in mitigating ray effects (Ref 3:205). The well known  $P_N$  and double- $P_N$  equations of nuclear reactor physics have inherent elliptic features which eliminate ray effects. Therefore space-angle synthesis using spherical harmonics seems to represent an approach which will eliminate ray effects.

The trial function expansion which will be used in this development is  $\frac{1}{2}$ 

$$T(z,\rho,\mu,\chi) = \sum_{iz=1}^{LZ} \sum_{ir=1}^{1R} \sum_{k=0}^{+L} \sum_{m=-k}^{+k} A_{iz,ir} B_{iz}(z) \cdot B_{ir}(\rho) Y_{km}(\mu,\chi)$$
(49)

where  $\rho$ ,z is the spacial coordinate dependence in cylindrical geometry and  $\mu$ ,  $\chi$  represents the  $\hat{\Omega}$  angular variable in Fig. 2.

 $\Psi(z,\rho,\mu,\chi)$  = even parity fluence at position r,z and in direction  $\mu,\chi$ ,

 $^{A}$ iz,ir = flux expansion or mixing coefficients,  $^{2}$ ,m

B<sub>iz</sub>(z) cubic polynomial spline in the z-variable (z-spline),

 $B_{i,p}(\sigma) = \text{cubic polynomial } \rho\text{-spline},$ 

 $Y_{\ell,m}(u_{\ell,\lambda})$  = spherical harmonic function, Eq (51).

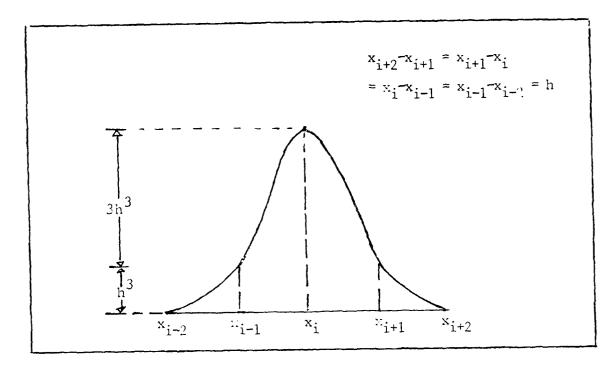


Figure 4. Cubic Spline With Evenly Spaced Nodes

iz, ir, £ and m are the trial function summation indices and

IZ = total number of z-splines,

IR = total number of  $\rho$ -splines,

L = degree of the spherical harmonic expansion.

The definition of a cubic polynomial splines (Ref 15:89) with evenly spaced nodes (knots) is

$$\begin{cases} (x_{i+2}^{2}-x)^{3}, & x_{i+1} \leq x \leq x_{i+2} \\ (x_{i+1}^{2}-x)^{3}-4(x_{i+1}^{2}-x)^{3}, & x_{i} \leq x \leq x_{i+1} \\ (x_{i+1}^{2}-x)^{3}-4(x_{i+1}^{2}-x)^{3}+6(x_{i}^{2}-x)^{3}, & x_{i+1}^{2} \leq x \leq x \end{cases}$$

$$(50)$$

$$\begin{cases} (x_{i+2}^{2}-x)^{3}-4(x_{i+1}^{2}-x)^{3}+6(x_{i}^{2}-x)^{3}, & x_{i+1}^{2} \leq x \leq x \end{cases}$$

$$(50)$$

The traph of this spire function is shown in Fig. 4. The spherical harmonic function (Net 6:600) will be defined as

$$Y_{\ell m}(\mu,\chi) = C_{\ell m} P_{\ell m}(\mu) e^{im\chi} = C_{\ell m} P_{\ell m}(\mu) \{\cos(m\chi) + i\sin(m\chi)\}$$
 (51)

where

$$C_{\ell m} = \sqrt{\frac{2\ell+1}{4\pi} \cdot \frac{(\ell-m)!}{(\ell+m)!}}$$
 (52)

The trial function expansion, Eq (49) can be made to satisfy the symmetry of Eq (20). This is a symmetric condition in  $\chi$  which directly implies that the solution must also be an even function in the variable  $\chi$ . Therefore, the angular trial function expansion of Eq (49) must also be even in  $\chi$ . Dropping the isin(m $\chi$ ) term from Eq (51) and substituting into Eq (49) gives

$$\Psi(z, \rho, \mu, \chi) = \sum_{i, z=1}^{L} \sum_{i, r=1}^{L} \sum_{\ell=0}^{L} \sum_{m=0}^{\infty} \Lambda_{iz, ir} B_{iz}(z) \cdot B_{ir}(\rho) \cdot Q_{\ell m}$$
 (53)

where

$$Q_{\ell m} = C_{\ell m} P_{\ell m}(\mu) \cos(m\chi)$$
 (54)

and by using the orthonormal properties of spherical harmonics the m index begins at zero instead of -% (see Appendix D).

The essential boundary condition at the air ground interface must also be applied to Eq (53). The fluence continuity requirements can be satisfied by this expansion in bicuoic polynomial splines and spherical harmonics. Both of thes functions are continuous throughout the problem domain. The cubic splines are also twice continuously

differentiable but the spatial z-derivative of the solution fluence is not continuous at the ground interface. However, the z-splines can be modified to have discontinuous derivatives at this interface (Ref 16). A Double-P<sub>N</sub> or Yvon's method (Ref 6:163) can also be used to accomodate the fluence discontinuity at  $\mu=0$ . In this development the air-ground interface will not be included in the problem domain, and therefore, this interface boundary condition will not be enforced.

Since a Galerkin method is being used the test or weight functions are

$$\gamma(z,\rho,\mu,\chi) = B_{jz}(z) B_{jr}(\rho)Q_{kn}$$
 (55)

Substituting Eqs (55) and (53) into Eq (48) produces Eq (58) where for simplicity the  $\rho,z,\mu,\chi$  dependencies are omitted and

$$B_{i} = B_{iz}(z)B_{ir}(\rho) \tag{56}$$

and

$$\beta_{\parallel} = \beta_{\parallel 2}(z)\beta_{\parallel r}(z) \tag{57}$$

$$\sum_{iz=1}^{IZ} \sum_{ir=1}^{IR} \sum_{\ell=0}^{L} \sum_{m=0}^{\ell} A_{ij} \int_{\ell m} \{\langle \widehat{\Omega} \cdot \nabla (E_{j} \cdot O_{kn}), K^{u}(\widehat{\Omega} \cdot \nabla (B_{i}O_{km})) \rangle$$

$$+ \langle B_{j}Q_{lm}, G^{g}(B_{j}Q_{lm}) \rangle d\hat{r} + \int_{S} \int_{S} \langle \hat{D} \cdot \hat{n} | B_{j}Q_{kn}B_{j}Q_{lm}d\hat{d}\hat{s} \rangle$$

$$= \int_{R} \langle \langle \hat{D} \cdot \nabla \rangle \langle B_{j}Q_{lm} \rangle, \langle u^{g}S^{u} \rangle - \langle B_{j}Q_{lm}, S^{g}\rangle d\hat{r}$$
(50)

This is a system of coupled algebraic equations where the unknown quantities are the  $A_{ij}$  mixing coefficients. Eq (58) can also be written as

$$\begin{bmatrix} \vec{K} \end{bmatrix} \begin{bmatrix} \vec{A} \end{bmatrix} = \begin{bmatrix} \vec{S} \end{bmatrix}$$

$$N \times N \quad N \times 1 \quad N \times 1$$

$$(59)$$

where

$$N = IZ \cdot IR \cdot (L+1) \cdot (L+2)/2 \tag{60}$$

 $\bar{K}$  = Coefficient or stiffness matrix, where each element is a summation of terms in the square brackets of Eq (58)

Ā = A<sub>iilm</sub> mixing coefficient vector

 $\ddot{S}$  = Source vector which is the right hand side of Eq (58)

The  $A_{ij2m}$  coefficients of Eq (53) will be obtained from a computer solution of Eq (59). These mixing coefficients can then be substituted into Eq (53) to give the even parity fluence,  $\Re(z,\rho,\mu,\chi)$ . In a solution of Eq (59) the elements of the K matrix and S vector must be computed. This computation involves an evaluation and summation of the individual expanded terms of Eq (58). This expansion is carried out in Appendix  $\Xi$ .

The directional gradient operator in cylindrical geometry is defined as (Ref 6:59)

$$\hat{\mathbb{S}} \cdot \hat{\mathbf{v}}_{z} = \sqrt{1 - u^{2} \cos \chi} \, \, \frac{\Im(\phi \phi)}{2\pi} \, - \, \frac{1}{2} \, \, \frac{\Im\{\phi \sqrt{1 - u^{2} \sin \chi}\} + u \frac{\partial \phi}{\partial z}}{2\pi}$$
 (61)

this is the conservative form of the directional derivative in two dimensional  $(\rho,z)$  geometry with azimuthal symmetry.

The  $K^{u}$  and  $G^{g}$  operators have been defined in Eqs (24) and (35) of Chapter III. The scalar product of the velocity and normal vectors is given in Appendix E as

$$\hat{\Omega} \cdot \hat{\mathbf{n}} = \begin{cases} \hat{\mathbb{C}} \cdot \hat{\mathbf{n}}_{\mathbf{z}} & \text{for the horizontal outer surfaces (top or bottom) of the problem cylinder, and as} \\ \hat{\mathbb{C}} \cdot \hat{\mathbf{n}}_{\rho} & \text{for the vertical surface (side) of the cylinder} \end{cases}$$
(62a)

where

$$\hat{\mathfrak{D}} \cdot \hat{\mathbf{n}}_{\mathbf{Z}} = \begin{cases} &\text{on the top surface, and} \\ -\mu &\text{on the bottom surface} \end{cases}$$
 (62b)

and

$$\hat{\Omega} \cdot \hat{\mathbf{n}}_{\rho} = \sqrt{1 - \mu^2 \cos \chi} \tag{62c}$$

The normal unit vectors  $\hat{n}_0$  and  $\hat{n}_z$  are shown in Fig. 9, Appendix E.

Expanding the expressions in Eq (58) produces an integral—differential equation which has twenty-eight terms (see Appendix E). These terms, except for the source terms, can be easily separated into a product of z,  $\rho$ ,  $\mu$  and  $\chi$  integrals. This is an integral separation of variables which is a direct result of Eq. (53); where, it is assumed that the solution can be expressed in a form where the dependent variables are separable. This separation property simplifies the individual integrals which have to be evaluated. It allows for the evaluation of only single integrals and not the more complicated double, triple or quadruple integrals. By this separation of variables it may be possible to integrate most of these single integrals analytically and thus avoid a numerical integration process.

The Spherical Harmonic Integrals. The use of a spherical harmonic angular trial function expansion was motivated by six basic considerations:

- 1. Because of the global nature of these functions the computational effort will be substantially reduced.
- 2. Spherical harmonics are well-known functions with orthonormal and symmetric (odd, even) properties.
- 3. The scattering cross sections are usually expanded in spherical harmonics (see Appendix A).
- 4. Spherical harmomics will produce a system of equations which are elliptic and invariant under continuous coordinate rotations (Ref 1:362).
- 5. An analytic or closed form evaluation of the angular integrals might be possible.
- 6. The even parity angular fluence, Eq (53), can be easily integrated to give the total particle fluence. This integration is carried out in Appendix I.

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The term-by-term expansion of Eq (58) has been partly carried out in Appendix E. The resulting angle integrals are only dependent on the degree of the spherical harmonic trial function expansion which is used. They are not dependent on the problem parameters and therefore they can be independently evaluated. They can be evaluated once, and thereafter, used as a part of the problem input data.

Three approaches were pursued in an attempt to evaluate the angle integrals which are produced by this expansion. The first approach was to use the orthogonal properties of the associated Legendre functions and the well-known properties of sines and cosines to analytically evaluate these integrals. However, because of their complicated nature (see Appendix F) a closed form analytic integration was not easily obtained for most of them.

The second approach was to use a computer routine that can evaluate these integrals in a symbolic or algebraic sense. Such a reutine will transform the integrals into algebraic expressions. The program Macvama which was written by the Massachusetts Institute of 1 cm, down Math Group was used in this attempt. However,

because of the time constraint on this research project and the need to learn a new programming language this approach was abandoned.

Finally, it was decided to evaluate these integrals by a numerical integration technique. Because it is possible to separate the integration variables, the computational effort can be substantially reduced by using a single (one variable) integration routine. So as to further reduce the computational effort, Eq (58) was completely expanded and twenty distinct angle integrals were identified. These integrals can be found in Appendix F. A ten point Newton-Cotes single integration routine was used to evaluate them. They were evaluated for each combination of the m, %, k and n trial and weight function subscripts.

Bicubic Polynomial Splines. The spatial dependence of the particle fluence in the air-over-ground problem is approximated by a product of cubic splines. The use of cubic polynomial splines in the trial function expansion of Eq (53) requires the formation of a tensor product space. This space is made up of bicubic polynomial splines which are products of p and z-splines on a rectangular grid (Rof 15:131). The exact shape of these bicubic splines are obtained from a variational principle or the equivalent method of weighted residuals (i.e., a solution of Eq (58)).

picuric polynomial splines are being used for the pillowin promons:

1. They are processed continuous and form a final crisis such that the integral continuous are set of the integral.

This reduces the number of integrals which must be evaluated and also produces a sparse and banded coefficient matrix.

- 2. A separation of the  $\rho$  and z integration variables is possible.
- 3. Third degree polynomials such as cubic splines have a faster rate of convergence than those of lower degree. Cubic splines are also twice continuously differentiable and thus they are very smooth functions. For a given problem mesh size splines will produce a coefficient (stiffness) matrix that is smaller but less sparse than hermites or lagrange polynomials.

The expansion of Eq (58) with a trial function of bicubic splines and spherical harmonics is carried out in Appendix E. A further expansion of these equations and a separation of the variables of integration produced seventeen distinct  $\rho$  and z integrals. These integrals which include the space source integrals are listed in Appendix G. The source integrals are derived from an interpolation of the first scatter source over the entire spatial problem domain.

The Source Torms. A numerical solution to the air-overground problem and the second order Boltzmann equation requires that the source terms (right hand side of Eq (58)) must be evaluated. These terms form the individual elements of the problem source vector in Eq (59). The even and odd parity sources  $S^{0}$  and  $S^{0}$  will be defined as the first scatter or collision even and odd parity sources. The first scatter source  $S(\hat{r},\hat{A})$  is the number density of particles which leave the burst point and undergo only one collision before being scattered into direction  $\hat{A}$  at position  $\hat{r}$ . Streaming neutrons which leave the burst point and do not collide before reaching position  $\hat{r}$ ,  $\hat{A}$  are not included in the collision source.

The use of a first scatter source makes the air-overground problem more isotropic. It removes the strongly anisotropic streaming particles from being a part of the problem source. Therefore, the solution fluence of Eq (58) will be the scattered even parity fluence  $\Psi_{\mathbf{s}}(\hat{\mathbf{r}},\hat{\Omega})$  and not the total even parity fluence  $\Psi_{\mathbf{t}}(\hat{\mathbf{r}},\hat{\Omega})$ . The total even parity fluence can be defined as

$$\Psi_{t}(\hat{r}, \hat{\Omega}) = \Psi_{s}(\hat{r}, \hat{\Omega}) + \Psi_{d}(\hat{r}, \hat{\Omega})$$
 (63)

where

$$\Psi_{d}(\hat{r}, \hat{\Omega})$$
 = streaming uncollided particles at position  $(\hat{r}, \hat{\Omega})$ .

A precise mathematical definition of the  $S^g$  and  $S^u$  sources will now be developed. Also a source interpolation procedure will be outlined. This source interpolation is used in order to simplify the source integrals of Appendix E (E-40 to E-46).

The First Scatter Sour . The even and odd parity sources have been defined as

$$S^{u}(\hat{\mathbf{r}},\hat{\boldsymbol{\beta}}) = \{ \{ S(\hat{\mathbf{r}},\hat{\boldsymbol{\beta}}) - S(\hat{\mathbf{r}},-\hat{\boldsymbol{\beta}}) \}$$
(10)

$$S^{g}(\hat{\mathbf{r}}, \hat{\mathbf{r}}) = \frac{1}{2} \{ S(\hat{\mathbf{r}}, \hat{\mathbf{r}}) + S(\hat{\mathbf{r}}, -\hat{\mathbf{r}}) \}$$
(9)

If  $S^u$  and  $S^g$  are first scatter source densities then  $S(\hat{r},\hat{n})$  and  $S(\hat{r},-1)$  must also be defined as first scatter source particles/unit volume. If a position (a,b) in the problem domain is chosen then a unit vector from the burst point (0,zb) can be defined as

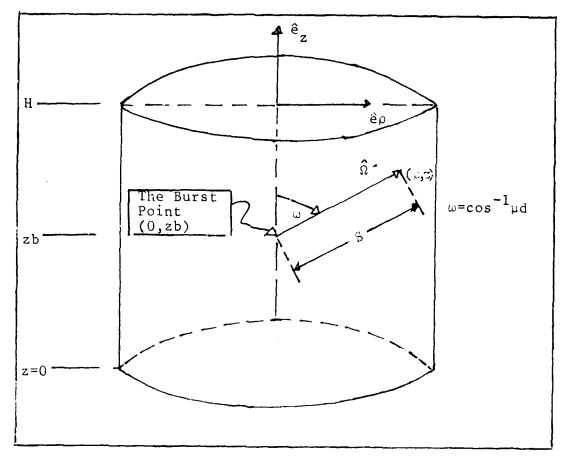


Figure 5. First Scatter (Collision)
Source Direction Vectors

$$\hat{\Omega}'(\rho,z) = \frac{\rho \hat{e}_{\rho} + (z-zb)\hat{e}_{z}}{\{\rho^{2} + (z-zb)^{2}\}^{\frac{1}{2}}}$$
(64)

Fig. 5 shows the direction vectors of this first scatter (collision) source.  $\hat{\Omega}$  is the direction that all streaming (uncollided) particles have at point (p,z).

By definition only particles which are streaming radially outward from the burst point can be included in the direct fluence. Therefore the direct fluence at point ( $\rho$ ,z) is in the  $\hat{\Omega}$  direction and can be written as

$$\phi_{d}(\rho,z,\hat{\Omega}') = \frac{Y}{4\pi S^{2}} \exp\{-\int_{0}^{S} \sigma_{t}(z)ds\}$$
 (65)

where

$$s = \{\rho^2 + (z-zb)^2\}^{\frac{1}{2}}$$
 (66)

and  $\int ds$  means that the integration is carried out along the path s (see Fig. 5). Also

Y = Total particle yield of the nuclear explosion at the burst point.

$$\sigma_{t}(z) = \begin{cases} \sigma_{t}(0)e^{-z/sh} & \text{for } z > 0 \\ \sigma_{t} & \text{(ground) for } z < 0 \end{cases}$$
(67)

sh = atmospheric scale height

The term  $\int_0^S \sigma_t(z) ds$  is the average number of collisions which a particle undergoes in traveling from the burst point (0,zb) to point  $(\rho,z)$ . From Fig. 5 the distance s can also be written as

$$s = (z-zb)/ud$$
 (68)

and therefore by changing variables

$$ds = dz/\mu d \tag{69}$$

where ud is a function of o and z (but constant along a path length S)

$$\mu d(z,z) = \cos \omega = (z-zb)/s = (z-zb)/\{\rho^2 + (z-zb)^2\}^{\frac{1}{2}}$$
 (70)

The integral term of Eq (65) can now be written for z > 0 as

$$\tau(\rho,z) = \int_{c}^{s} \sigma_{t}(z) ds = \frac{\sigma_{t}(o)}{ud} \int_{zb}^{z} e^{-z/sh} dz$$
 (71)

and finally as

$$\tau(\rho,z) = \frac{\sigma_{t}(\rho)}{\mu d} \{e^{-zb/sh} - e^{-z/sh}\} \cdot sh$$

$$= \{\sigma_{t}(zb) - \sigma_{t}(z)\} \cdot sh$$

From the above derivation it follows that

$$\tau(\rho,z) = \begin{cases} \{\sigma_{t}(zb) - \sigma_{t}(z)\} \cdot \sinh/\mu d \text{ for } z > 0 \\ \{\sigma_{t}(zb) - \sigma_{t}(0)\} \cdot \sinh/\mu d - \sigma_{t}z/\mu d \text{ for } z < 0 \end{cases}$$
 (73)

also

$$\phi_{\mathbf{d}}(\rho, \mathbf{z}, \hat{\Omega}') = \frac{Y}{4\pi S^2} \exp(-\tau(\rho, \mathbf{z}))$$
 (74)

Note that  $\phi_d(\rho,z,\hat{\Omega}')$  is only a function of  $\rho$  and z.

The first scatter source at  $(\rho,z)$  and with direction  $\hat{\Omega}$  are those particles which undergo their first collision at  $(\rho,z)$  and are scattered from direction  $\hat{\Omega}'$  to  $\hat{\omega}$ .

Therefore the first scatter source can now be defined as

$$S(\rho,z,\hat{\Omega}) = \sigma^{S}(z,\hat{\Omega}\cdot\hat{\Omega}')\phi_{d}(\rho,z,\hat{\Omega}')$$
 (75)

where  $\sigma^{\bf S}$  is not a function of  $\hat{\Omega}$  but of the scattering angle  $\hat{\Omega} \cdot \hat{\Omega}' = 0$  and z. From Fig. 5 and Fig. 2  $\hat{\Omega}'$  is defined by  $\mu d$  and  $\chi = 0$  i.e.  $\hat{\Omega}' = (\mu', \chi')$  where  $\mu' = \mu d$  and  $\chi' = 0$ .

By use of the addition theorem it is shown in Appendix H that Eq (75) can be written as

$$S(s,z,\hat{z}) = 2I_{d}(r,z,\hat{z})\gamma_{0}^{-z/sh}\sum_{k=0}^{L}\sum_{m\neq k=0}^{\infty} \frac{s}{s}(0)C_{0m}^{2}\gamma_{m}(u)P_{dm}(ud)\cos m\chi$$

$$(76)$$

and that the even and odd parity first scatter sources are

$$S^{g}(\rho,z,\hat{\Omega}) = \psi_{\mathbf{d}}(\rho,z,\hat{\Omega}')e^{-z/sh}\sum_{\ell=0}^{L}\sum_{m^{k}=0}^{\ell}\sigma_{\ell}^{s}(0)C_{\ell m}^{2}\{1+(-1)^{\ell-m}\}P_{\ell m}(\mu \mathbf{d})P_{\ell m}(\mu)ccsm\chi$$
(77)

and

$$S^{tt}(\rho,z,\hat{\Omega}) = \psi_{d}(\rho,z,\hat{\Omega}')e^{-z/sh}\sum_{k=0}^{L}\sum_{m=0}^{k}\sigma_{k}^{s}(0)C_{om}^{2}\{1-(-1)^{k-m}\}P_{km}(\mu d)P_{km}(\mu)cosm\chi$$
(78)

where  $m^*$  means that all terms with a m = 0 subscript must be divided by two, and cosm $\chi$  should be interpreted as cosine (m $\chi$ )

Source Interpolation. Because of the complicated nature of the source expressions, Eqs (77) and (78), and the need to integrate the source terms of Appendix E, a spatial source interpolation will be used. This interpolation which simplifies the source integrals is necessary if a very tedious (double or quadruple) integration is to be avoided. By this interpolation process the source terms of Appendix E can all be separated into a product of single integrals.

It is important to note that  $\mu d$  which is given by Eq (70) is a function of  $\rho$  and z and therefore  $\mathbb{P}_{2m}(\mu d)$  is also a function of  $\rho$  and z. Furthermore  $\phi_{2}(\rho,z,\hat{z}')$  of Eq (74) is a function of  $\rho$  and z. Beginning with Eqs (77) and (78) they can be rewritten as

$$S^{g}(\mu,z,\hat{\Omega}) = \sum_{n=0}^{L} \sum_{m'=0}^{\ell} \sigma_{n}^{g}(0)C_{2m}^{2}(1+(-1)^{\ell-m})\Lambda_{\ell,m}(\mu,z)P_{\ell,m}(\mu)\cos(m\chi)$$
(79)

and

$$S^{u}(c,z,\hat{\Omega}) = \sum_{\ell=0}^{L} \sum_{m^{\ell}=\ell}^{\ell} \sigma_{\ell}^{s}(0) c_{\ell m}^{2} \{1 - (-1)^{\ell-m}\} A_{\ell m}(\rho,z) P_{\ell m}(\mu) cosm\chi$$
 (80)

where

$$N_{\ell m}(\rho, z) = \phi_{\mathbf{d}}(\rho, z, \hat{\Omega}^{-}) P_{\ell m}(\mu d) e^{-z/sh}$$
(81)

A spatial  $(\rho,z)$  interpolation of the even and odd parity sources, Eqs (79) and (80), is therefore an interpolation of  $A_{\ell m}(\rho,z)$ . In this project these first scatter second order sources were interpolated by a combination of piecewise bi-linear Lagrange polynomial functions. Specifically,  $S^g$  was approximated by a tensor product of linear Lagrange polynomials as follows

$$S^{g}(\rho,z,\hat{\rho}) = \sum_{i=1}^{NZ} \sum_{j=1}^{NR} S^{g}(\rho_{j},z_{i},\hat{\Omega})H_{j}(\rho)H_{i}(z)$$
(82)

where

$$S^{g}(\rho_{j},z_{i},\hat{z}) = \text{the even parity source, Eq (79)}$$
 evaluated at the spacial nodes  $(\rho_{j},z_{i})$ 

NZ = total number of z-nodes

NR ≈ total number of R-nodes

 $H(\rho) \approx \rho$ -linear Lagrange polynomial

II(z) = z-linear Lagrange polynomial

These linear polynomials are defined as

$$W_{i}(x) = \begin{cases} \frac{x_{i+1} - x}{x_{i+1} - x_{i}}, & x_{i+1} < x_{i} = 0, & \text{if } x_{i} < x_{i} = 0, \\ \frac{x_{i+1} - x_{i}}{x_{i+1} - x_{i}}, & x_{i} < x_{i+1} < x_{i} \end{cases}$$

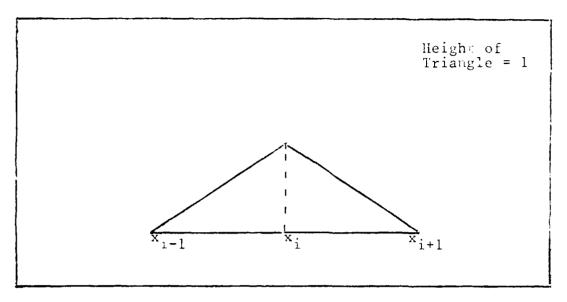


Figure 6. A Linear Lagrange Polynomial Function.

The product  $H_j(z) \cdot H_i(z)$  of Eq (82) forms a tensor product space on a rectangular grid, in the  $\rho$ , z plane (Ref 15:129) Substituting Eq (82) for  $S^g(\rho_j, z_i \hat{})$  in Eq (77) gives

$$S^{g}(\rho,z,\hat{\Omega}) = \sum_{k=0}^{L} \sum_{m^{N}=0}^{2} [z_{\lambda}^{g}(0)C_{km}^{2}\{1 - \sum_{m=0}^{NZ} \sum_{m} \sum_{m=0}^{NZ} \sum_{m} \Delta_{km}(\rho_{j},z_{i})H_{j}(\rho)H_{i}(z)]$$

$$= \sum_{m^{N}=0} \sum_{m^{N}=0}^{2} +(-1)^{k-m} P_{km}(\mu) cosm\chi \sum_{m} \sum_{m} \Delta_{km}(\rho_{j},z_{i})H_{j}(\rho)H_{i}(z)$$

$$= \sum_{m^{N}=0} \sum_{m^{N}=0}^{2} \sum_{m^{N}=0}^{2} (1 - \sum_{m^{N}=0}^{NZ} \sum_{m} \sum_{m} \Delta_{km}(\rho_{j},z_{i})H_{j}(\rho)H_{i}(z)$$

$$= \sum_{m^{N}=0} \sum_{m^{N}=0}^{2} \sum_{m^{N}=0}^{2} (1 - \sum_{m^{N}=0}^{NZ} \sum_{m} \sum_{m} \Delta_{km}(\rho_{j},z_{i})H_{j}(\rho)H_{i}(z)$$

$$= \sum_{m^{N}=0} \sum_{m^{N}=0}^{2} \sum_{m^{N}=0}^{2} (1 - \sum_{m^{N}=0}^{NZ} \sum_{m} \sum_{m^{N}=0}^{NZ} \sum_{m^{N}=0}^{$$

Similarly the odd parity source can also be expressed as

$$S^{u}(s,z,\hat{s}) = \sum_{k=0}^{L} \sum_{m'=0}^{\infty} \left[ z_{s}^{s}(0) c^{2} \int_{\mathbb{R}^{m}} \left\{ 1 - (-1)^{\frac{2}{2} - m} \right\} P_{\xi m}(u) \cos m \sum_{k=1}^{NZ} \sum_{j=1}^{NR} A_{\xi m}(x_{j},z_{i}) H_{j}(s) H_{i}(s) \right]$$
(85)

Dis (30) and (7)? In now be substituted into the source terms of Appendix 7 (5-40 to E-76). This substitution and a separation of his progration variables produced six distinct space integrals. These space integrals are listed in Appendix

G. The source angle integrals have been included in those integrals which are presented in Appendix F.

# V Computer Implementation and Results

A numerical solution of the air-over-ground problem can be obtained from a computer solution of the system of coupled algebraic equations which are represented by Eq (58). These equations can be written in the matrix form of Eq (59) and through a direct inversion or some other matrix solution process the  $A_{i,j,2,m}$  mixing coefficients can be found.

The computer solution to this problem was accomplished by a two step process. Each element of the stiffness matrix and source vector was computed and assembled in the matrix form of Eq (59). Then the mixing coefficients were computed by the iterative matrix solution method of successive over-relaxation. An indirect iterative matrix solution method is possible because Eq (58) and the stiffness matrix of Eq (59) is symmetric positive definite. A computer program which assembles the problem matrices and computes the mixing coefficients and particle fluences was written. This program which is written in FORTRAN V is listed in Appendix J.

Using a ten point Newton-Cotes numerical integration routine, each of the thirty-seven integrals in Appendix F and G were evaluated. The angular integrals were evaluated for each om, kn combination of the trial and weight function subscripts. Selected products of these integrals were then used to generate each of the twenty-eight terms (E-19 to E-46) of Appendix E. Following the prescription of Appendix

E the first twenty-one terms were then added (or subtracted) to produce the elements of the stiffness matrix.

The next seven terms gave the elements of the source vector.

Writing the synthesized Boltzmann equation, Eq (58), in operator notation as

$$L(\forall_{iz,ir}) \eta_{jz,jr} = S \eta_{jz,jr}$$

$$\downarrow_{i,m} \qquad k,n \qquad k,n$$
(86)

where  $\Psi$  and  $\gamma$  are defined by Eqs (53) and (55) and

L(
$$\forall_{iz,ir}$$
) $\eta_{jz,jr}$  = Left hand side of Eq (58)  
 $\ell$ , m k, n  
S $\eta_{jz,jr}$  = Right hand side of Eq (58)  
k,n

the  $K_{p,q}$  element of the stiffness matrix is

$$K_{p,q} = L\{B_{iz}(z)B_{ir}(\rho)Q_{2m}\} \cdot B_{jz}(z)B_{jr}(\rho)Q_{kn}$$
 (37)

where

$$q = (m+1) + \lambda(\frac{\lambda+1}{2}) + (\frac{Lmax + 1}{2})(\frac{Lmax + 2}{2})(ir-1) + IR(iz-1)$$
 (88)

and

$$p = (n+1) + k(\underline{k+1}) + (\underline{\max + 1})(\underline{\max + 2}) \cdot t(jr-1) + JR(jz-1)$$
 (39)

The corresponding source vector element  $\mathbf{S}_{n}$  is given by

$$S_{p} = S \cdot B_{jz}(z)B_{jr}(\rho)Q_{kn}$$
 (90)

where p is given by Eq (89) and

Lmax = degree of the spherical harmonic expansion

JR=IR = total number of  $\rho$ -splines

p = the row index of the problem (K) matrix

q = column index of matrix K

B(z), B( $\beta$ ) and Q<sub>lm</sub> are defined in Chapter IV (Eqs (50) and (54)). iz, ir, jr, jz,  $\lambda$ , m, k, and n are the trial and weight function expansion subscripts where

 $\ell, k = 0$  to Lmax

 $m = 0 \text{ to } \ell$ ,

n = 0 to k

iz, jz = 1 to IZ

ir, jr = 1 to IR

and

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IZ = total number of z-splines

IR = total number of  $\rho$ -splines

Using the notation of Eqs (87), (88) and (89) the K-problem matrix and S-source vector can be easily assembled in the following do loop.

DO 10 jz = 1, IZ

DO 10 jr = 1, IR

DO 10 k = 0, Lmax

DO 10 n = 0, k

 $s_p = S*n_{iz, ir, k, n}$ 

DO 10 iz = 1, 1Z

DO IO ir = 1, IR

DO 10 2 = 0, Lmax

DO 10 m = 0.0

$$K_{p,q} = L(\Psi_{iz,ir}) \eta_{jz,jr}$$

10 continue

The assembled coefficient (K) matrix which results from Eq (58) and a bicubic spline spherical harmonic trial function expansion is sparse and symmetric. Because of this symmetry and sparseness the coefficient matrix can be stored within the computer by special storage schemes (Ref 14:70). These sparse and symmetric schemes will greatly reduce the computer storage requirements. As an example, in a trial function expansion where JR = IR = 8 and Lmax = 2 the problem coefficient matrix is a 384 x 384 square symmetric matrix with many elements which are zero. Therefore if this entire matrix is stored within the computer it will req. 9 147456 separate storage locations. This much core storage is already beyond the capacity of most computers. However, by using a sparse and symmetric storage scheme this matrix can be reduced to one with less than 73728 elements. For large trial function expansions (JR = IR = 50, Lmax = 3), special auxiliary storage and solution techniques will be necessarv.

This entire problem (coefficient and source matrices) was assembled on a CDC 6600 computer at the Air Force Institute of Technology. The 384 x 384 problem matrix is too large to be stored in core memory unless a sparse symmetric storage mode is used. Because some of the  $(\rho)$  integrals in Appendix G are discontinuous (infinite) at

 $\rho$  = 0 it was necessary to use a lower  $\rho$ -integration limit of 1.0 E-8. Also, since the first scatter sources of Eqs (77) and (78) are undefined at the burst point none of the problem nodes can be located there (see Fig. 5).

#### Results

The computer routines which are listed in Appendix J were used to produce a numerical solution to the air-over-ground problem. These routines demonstrate the feasibility of using FESAS to produce a computer solution to the two-dimensional steady state anisotropic Boltzmann equation.

This computer program has not been fully developed, refined or debugged and therefore the accuracy of the results has not been evaluated. These results are presented in an attempt to further show that FESAS is a viable numerical solution technique.

The problem domain is a cylinder which sits on the surface of the earth (see Fig. 5). However, the air-ground interface is not included in the problem domain and therefore all ground effects are ignored. The following problem parameters were used

Weapon yield = 1.0E+23 neutrons

Cylinder height = .4km

Cylinder radius = .4km

Burst height = .12 km

Total cross section  $(\varepsilon_{\rm t}(0))$  = 15.0 km<sup>-1</sup>

Table I

Legendre Expansion Coefficients which were used in a Numerical Solution of the Air-Over-Ground Problem

Expansion subscript &	Legendre expansion coefficients					
	σç	σu				
0	10.0	.0				
1	0.0	2.5				

The cross-sections in Table I were arbitrarily chosen and they do not represent the actual values for air at sea level. A relative convergence criteria (.001) which is accurate to three significant figures was used.

The program execution times varied with the degree of the spherical harmonic trial function expansion and the problem mesh (grid) size. The entire problem matrices were stored within core memory and by trial and error it was determined that a relaxation parameter of 1.7 gave the fastest convergence rate. However, as more trial functions were used and the system of equations and matrices grow larger the rate of convergence decreased. In Table II the program execution times and the number of iterations to convergence are listed for various problem mesh sizes. These execution times and convergence rates can be substantially reduced by rewriting or meditying the program routines of appendix J.

Execution Times and Convergence Rates for the FESAS Computer Code Table II

The second secon

Number of Iterations to Convergence	1395	2001	785	801	265	492	168	
Execution Times (secs)	47.8	351.754	60.315	361.3	79.62	816.7	236.83	
Degree (Lmax) of the Spherical Harmonic Trial Function Expansion	0	1	0	Н	0	1	0	
Spatial Grid	1×1	1×1	2×2	2×2	4×4	4×4	8×8	

# PARTICLE (NEUTRON) FLUENCES.

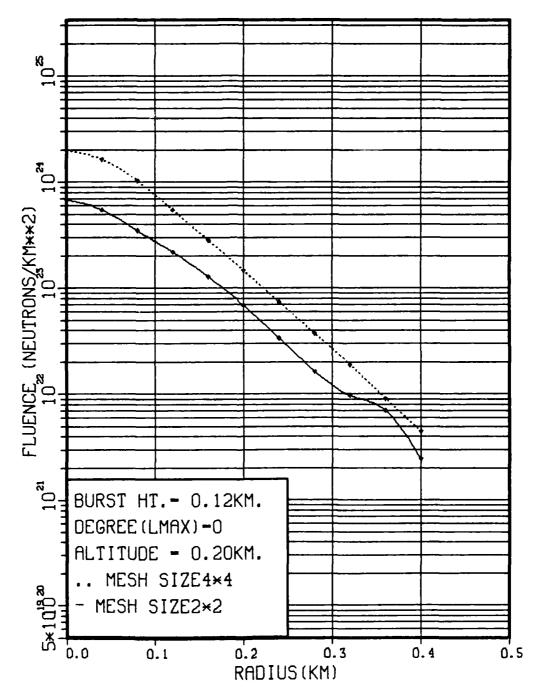


Figure 7. Computed Fluences as a Function of Radius and showing a variation with the Problem Spatial Mesh Size.

# PARTICLE (NEUTRON) FLUENCES.

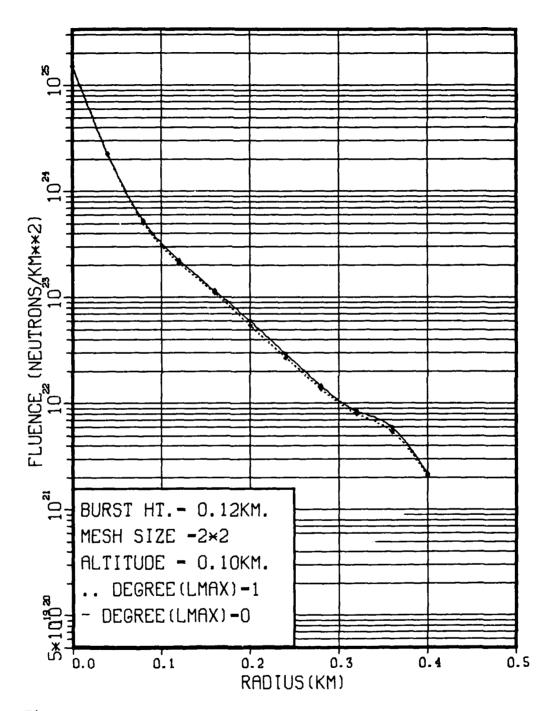


Figure 8. Computed Fluences as a Function of Radius and showing a variation with the Degree of the Angular Spherical Harmonic Trial Function Expansion.

In Figures 7 and 8 fluence values as a function of radius are presented. These values are representative of an altitude of 200 metres and the aforementioned problem parameters. Figure 7 shows a variation of the solution with spatial grid (mesh) sizes whereas Figure 8 shows a variation with the degree of the spherical harmonic trial function expansion. More detailed numerical results can be found in Appendix K.

# VI Conclusions and Recommendations

# Conclusions

A finite element space-angle synthesis (FESAS) solution of the steady state anisotropic Boltzmann equation in two-dimensional cylindrical geometry has been presented. In this presentation a weak form of the even parity steady state Boltzmann equation was developed. It was shown that because the problem equations were positive definite and self-adjoint the Rayleigh-Ritz variational principle and the Bubnov-Galerkin method of weighted residuals are equivalent. The problem solution was formulated by using a trial function expansion in bicubic polynomial splines and spherical harmonics. This trial function expansion has a dual basis — a local basis in space and a global basis in angle.

This development was specialized to the air-over-ground neutral particle transport problem. It was shown that a finite element space-angle synthesis solution is possible and that a first scatter interpolation source can be used. It was also shown that a numerical solution can be achieved and that this solution technique may eliminate ray effects and reduce computational costs.

# Recommendations

The preliminary results of this study have shown that the FESAS method can produce a numerical solution to the steady state Boltzmann equation and the air-over-ground problem. However, because of the time constraints on this

research project a complete development and evaluation of the FESAS method was not accomplished. Therefore, there are a number of recommendations for the further analysis and evaluation of the FESAS method, which must be made. Some of these recommendations are:

- 1. To enforce the boundary conditions at the air-over-ground interface. This can be accomplished by a coalescing or stacking of the nodes (knots) of the bicubic splines and by using a Double-P<sub>M</sub> approximation at this interface.
- 2. Develop or refine the computer algorithm so that a comparative study can be made. This study should include a comparison of the computational costs and accuracy of FESAS to those of Monte Carlo and discrete ordinates. Also, a determination should be made as to whether ray effects have been eliminated.
- 3. Obtain, if possible, a closed form solution to the angle integrals in Appendix F.

The second secon

- 4. Explore other ways of handling the discontinuity (at  $\rho = 0$ ) of the space integrals.
- 5. Use other spatial trial functions. Lower degree bi-quadratic splines, hermites and Lagrange polynomials are possible candidates. A comparison of the results which are obtained from the use of various trial functions can then be made.
- 6. Examine the effects of an improved source interpolation on the solution accuracy and rate of
  convergence. An improved source interpolation
  can be achieved by the use of a smaller source
  (space) grid or a higher degree Lagrange or
  Hermite polynomial interpolation function
- 7. Extend the use of finite element space-angle synthesis to the solution of energy dependent multi-group problems.

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### Appendix A

Derivation of the Scattering Kernel, Inverse Collision Operators and the Even and Odd Parity Collision Cross Sections

# Collision Cross-Sections

The even and odd parity legendre expansion scattering cross-sections  $\sigma_{\ell}^g$  and  $\sigma_{\ell}^u$  originate in the derivation of the second order forms of the Boltzman equation (Chapter III). A definition of these quantities has been given elsewhere (Ref 5:29). This definition is repeated below.

Beginning with the even and odd parity scattering cross sections

$$G_{S}^{g}(\hat{F}_{S}\hat{\Lambda}.\hat{\Lambda}') = \frac{1}{2} \{G^{S}(\hat{F}_{S}\hat{\Lambda}.\hat{\Lambda}') + G^{S}(\hat{F}_{S}-\hat{\Lambda}.\hat{\Lambda}')\}$$
 (A-1)

$$6_{s}^{u}(\hat{r}_{5}\hat{\Lambda}.\hat{\Lambda}') = \frac{1}{2} \left\{ 6(\hat{r}_{5}\hat{\Lambda}.\hat{\Lambda}') - 6^{s}(\hat{r}_{5}-\hat{\Lambda}.\hat{\Lambda}') \right\}$$
 (A-2)

the usual computational practice (Ref 6:131) is to expand these scattering cross sections in Legendre polynomials as

$$G^{S}(\hat{f},\hat{\Omega}.\hat{\Omega}) = \int_{0}^{L} G(\hat{f}) \cdot 2\ell + 1 \cdot P_{\ell}(\hat{\Omega}.\hat{\Omega}')$$
(A-3)

or

$$6^{5}(\hat{r}_{3}-\hat{\Lambda}.\hat{\Lambda}) = \sum_{\ell=0}^{2} 6_{\ell}^{5}(\hat{r}) \cdot 2\ell+1 \cdot P_{\ell}(\hat{\Lambda}.\hat{\Lambda}')$$
(A-4)

where

 $\sigma^3(r,\hat{\mathbb{Q}}\cdot\hat{\mathbb{Q}}')$  = macroscopic scattering cross section

σ<sup>S</sup><sub>c</sub>(r) = Legendre macroscopic cross-section expansion coefficient which is a function of position (material)

L = the degree Legendre polynomial expansion which is used  $P_{\ell}(\Omega \cdot \Omega')$  = Legendre polynomial of degree  $\ell$ .  $\hat{\Omega} \cdot \hat{\Omega}'$  = scattering angle ( $\mu_{O}$  = Cos $\theta$ )

Inserting Eqs (A-3) and (A-4) into (A-1) and (A-2)

gives

$$G_{s}^{3}(\hat{r},\hat{\Lambda}.\hat{\Lambda}') = \sum_{\ell=0}^{L} G_{s}^{2}(\hat{r}) \frac{2\ell+1}{4\pi} \left\{ P_{s}(\hat{\Lambda}.\hat{\Lambda}') + P_{s}(\hat{\Lambda}.\hat{\Lambda}') \right\}$$
(A-5)

$$6_{s}^{"}(\tilde{F},\hat{\Lambda}.\hat{\Lambda}) = \sum_{\ell=0}^{L} 6_{\ell}^{s}(\hat{F}) \frac{2\ell+1}{4\pi} \left\{ P_{\ell}(\hat{\Lambda}.\hat{\Lambda}) - P_{\ell}(\hat{\Lambda}.\hat{\Lambda}) \right\}$$
(A-6)

From the even and odd properties of Legendre polynomials (Ref 17:223)

$$P(\tilde{x}.\tilde{x}) = \begin{cases} \text{even function if } \ell \text{ is even} \\ \text{odd function if } \ell \text{ is odd} \end{cases}$$
 (A-7)

$$\sum_{\ell=0}^{L} \left\{ P_{\ell}(\hat{\lambda}.\hat{\lambda}') - P_{\ell}(\hat{\lambda}.\hat{\lambda}') \right\} = \begin{cases} 2P_{\ell}(\hat{\lambda}.\hat{\lambda}') & \text{if } \hat{\lambda} \text{ is odd} \\ 0 & \text{if } \hat{\lambda} \text{ is even } (A-9) \end{cases}$$

Eqs (A-5) and (A-6) can now be rewritten as

$$G_{S}^{g}(f,\hat{n}.\hat{n}') = G_{G}^{S}(f) \frac{2l+1}{4\pi} \cdot P_{G}(\hat{n}.\hat{n}')$$

$$= G_{S}^{g}(f,\hat{n}.\hat{n}') = G_{G}^{S}(f) \frac{2l+1}{4\pi} \cdot P_{G}(\hat{n}.\hat{n}')$$
(A-10)

and

$$G_{3}^{\prime\prime}(\hat{F},\hat{\Lambda}.\hat{\Lambda}') = \int_{-2\pi}^{L} G_{3}^{\prime\prime}(\hat{F},\hat{\Lambda}.\hat{\Lambda}') \frac{2\pi}{4\pi} P_{3}(\hat{\Lambda}.\hat{\Lambda}')$$
(A-11)

The even and odd parity cross-sections can also be expanded in Legendre polynomials as

$$G_{2}^{3}(\hat{r},\hat{n}.\hat{n}) = \int_{-\infty}^{\infty} G_{2}^{3}(\hat{r}) \cdot \frac{2\ell+1}{4\pi} \cdot P_{2}(\hat{n}.\hat{n}')$$
(A-12)

anu

$$G_{3}^{\prime}(\hat{F},\hat{n}.\hat{n}') = G_{2}^{\prime}(\hat{F}). \frac{2\ell+1}{4\pi} P_{2}(\hat{n}.\hat{n}')$$
(A-13)

Comparing Eqs (A=13), (A=12), (A=11) and (A=10) it is apparent that

$$\sigma_{\ell}^{g}(\hat{\mathbf{r}}) = \begin{cases} \sigma_{\ell}^{s}(\hat{\mathbf{r}}) & \text{for } \ell \text{ even} \\ 0 & \text{for } \ell \text{ odd} \end{cases}$$
 (A-14)

and

$$\sigma_{\ell}^{\mathbf{u}}(\hat{\mathbf{r}}) = \begin{cases} 0 & \text{for } \ell \text{ even} \\ \sigma_{\ell}^{\mathbf{s}}(\hat{\mathbf{r}}) & \text{for } \ell \text{ odd} \end{cases}$$
 (A-15)

Therefore  $\sigma_{\ell}^g(\hat{r})$  and  $\sigma_{s}^g(\hat{r}, \hat{\Omega} \cdot \hat{\Omega}')$  are even functions. Also  $\sigma_{\ell}^u(\hat{r})$  and  $\sigma_{s}^u(\hat{r}, \hat{\Omega} \cdot \hat{\Omega}')$  are odd functions.

# Scattering Kernel

In the development of the even and odd parity forms of the anisotropic steady state Boltzmann equation (Chapter III), it was necessary to express the scattering kernels in terms of the even and odd parity flux (fluence) components. Following the derivation of Wheaton (Ref 5:11) the scattering terms can be written as

$$\int_{4\pi}^{3} (\hat{r}_{j}\hat{\Lambda}.\hat{\Lambda}) \phi(\hat{r}_{j}\hat{\Lambda}') d\hat{\Lambda}' = i \int_{4\pi}^{3} (\hat{r}_{j}\hat{\Lambda}.\hat{\Lambda}') \phi(\hat{r}_{j}\hat{\Lambda}') d\hat{\Lambda}'$$

$$+ i \int_{4\pi}^{3} G_{j}(\hat{r}_{j}\hat{\chi}.\hat{\Lambda}') \phi(\hat{r}_{j}\hat{\Lambda}') d\hat{\Lambda}'$$

$$+ i \int_{4\pi}^{3} G_{j}(\hat{r}_{j}\hat{\chi}.\hat{\Lambda}') \phi(\hat{r}_{j}\hat{\Lambda}') d\hat{\Lambda}' \qquad (A-10)$$

where the integrations are carried out over all directions.

Because  $\sigma_s^g$  is an even function, meaning f(x) = f(-x), it follows that

$$G_{s}^{g}(\hat{r},\hat{n}-\hat{n}') = G_{s}^{g}(\hat{r}_{s}\hat{n}\cdot\hat{n}') \tag{A-17}$$

and therefore

$$\int_{\mathfrak{M}} \mathcal{G}_{\mathfrak{K}}^{\mathfrak{I}}(\hat{r},\hat{n},\hat{n}') \, \phi(\hat{r},\hat{n}') \, d\hat{n}' = \int_{\mathfrak{M}} \mathcal{G}_{\mathfrak{K}}(\hat{r},\hat{n},\hat{n}') \, \left\{ \phi(\hat{r},\hat{n}) + \phi(\hat{r},\hat{n}') \right\} \, d\hat{n}' \qquad (A-18)$$

With the even parity flux defined as

$$\psi(\hat{r},\hat{\Lambda}') = \frac{1}{2} \left\{ \phi(\hat{r},\hat{\Lambda}') + \phi(\hat{r},-\hat{\Lambda}) \right\} \tag{A-19}$$

Eq (A-18) becomes

$$\int_{an} G_s^3(\hat{r},\hat{\Lambda}.\hat{\Lambda}') \phi(\hat{r},\hat{\Lambda}') d\hat{\Lambda}' = \int_{an} G_s^3(\hat{r},\hat{\Lambda}.\hat{\Lambda}') \psi(\hat{r},\hat{\Lambda}') d\hat{\Lambda}'$$
(A-20)

and by a similar derivation

$$\int_{4\pi} G_s^{\mu}(\hat{F}, \hat{\Lambda}.\hat{\Lambda}') \, \phi(\hat{F}, \hat{\Lambda}) d\hat{\Lambda}' = \int_{4\pi} G_s^{\mu}(\hat{F}, \hat{\Lambda}.\hat{\Lambda}') \, \chi(\hat{F}, \hat{\Lambda}') d\hat{\Lambda}' \qquad (A-21)$$

## Inverse Collision Operators

In deriving the second order forms of the Boltzmann equation (Chapter III) the  $G^g$  and  $G^u$  operators were defined as

$$G^{3}f(\hat{\Lambda}) = G_{\varepsilon}f(\hat{\Lambda}) - \int_{0}^{2} G_{\varepsilon}^{3}(\hat{\Lambda}.\hat{\Lambda}')f(\hat{\Lambda}')d\hat{\Lambda}' \qquad (A-22)$$

and

$$G^{\alpha}f(\hat{n}) = \epsilon_{\alpha}f(\hat{n}) - \epsilon_{\alpha}f(\hat{n}.\hat{n})f(\hat{n})d\hat{n}'$$
(A-23)

where the  $\hat{r}$  dependence has been omitted in an attempt to simplify the notation.

Using the addition theorem (Ref 6:609)

$$P_{\ell}(\hat{\lambda}.\hat{\lambda}') = \frac{4\pi}{2\ell+1} \sum_{m=-\ell}^{+\ell} \frac{1}{\ell m} \hat{\lambda}' \hat{$$

in (A-12) and (A-13) gives

$$G_{s}^{3}(\hat{\Lambda}.\hat{\Lambda}') = \sum_{k=0}^{L} \int_{m_{s}-k}^{2k} \left(\frac{1}{2}\right) \left\langle \frac{1}{2}\right\rangle \left(\frac{1}{2}\right)$$
(A-25)

and

$$G_{s}^{u}(\hat{n}.\hat{\Lambda}') = \sum_{\ell=0}^{L} \int_{m_{\ell}-\ell}^{u} G_{\ell}^{u} V(\hat{\Lambda}') V(\hat{\Lambda})$$

$$(A-26)$$

Eqs (A-22) and (A-23) can now be written as

$$G^{3}f(\hat{\Lambda}) = Gf(\hat{\Lambda}) - \sum_{e=0}^{L} G_{e}^{3} V(\hat{\Lambda}) \left(\hat{\Lambda}\right) \left(\hat{\Lambda}\right) d\hat{\Lambda}'$$

$$(A-27)$$

and

$$G_{f}^{u}(\hat{\Lambda}) = G_{f}^{f}(\hat{\Lambda}) - \sum_{\ell=0}^{L} \frac{f_{\ell}^{u}}{M} Y(\hat{\Lambda}) \begin{cases} Y(\hat{\Lambda}^{\ell}) f(\hat{\Lambda}^{\ell}) d\hat{\Lambda}^{\ell} \\ y(\hat{\Lambda}^{\ell}) = 0 \end{cases}$$

$$(A-28)$$

The inverse operators are defined as

$$\mathcal{K}^{g} = [G^{g}]^{-1} \quad \text{and} \quad \mathcal{K}^{u} = [G^{u}]^{-1} \tag{A-29}$$

where it is meant that if

$$G^{\mathcal{G}}(\hat{\mathcal{L}}) = \mathcal{R}(\hat{\mathcal{L}}) \tag{A-30}$$

then

$$K^{3}[G^{3}f(\hat{\lambda})] = K^{3}R(\hat{\lambda}) = f(\hat{\lambda})$$
(A-3)

multiplying Eq (A-30) by  $\int_{4\pi} Y_{kn}^*(\hat{\Omega}) d\hat{\Omega}$  and expanding the  $C^g$  operator gives

$$6 \int f(\hat{x}) \bigvee_{k} \hat{f}(\hat{x}) d\hat{n} - \int 6 \int \int f(\hat{x}) \bigvee_{k} \hat{f}(\hat{x}) d\hat{n} \int f(\hat{x}') d\hat{n}' d\hat{n$$

using the orthonormal properties of spherical harmonics (Ref 6:609)

$$\int_{AR} \frac{y(\hat{x})y(\hat{x})}{kn} d\hat{x} = \mathcal{E}_{ek}\mathcal{E}_{mn}$$
(A-33)

where  $\delta_{ik}$  is the Kronecker delta which is defined by

$$\mathcal{S}_{\ell k} = \begin{cases} 0 & \ell \neq k \\ 1 & \ell = k \end{cases} \tag{A-34}$$

Therefore Eq (A-32) can now be written as

$$6 \int f(\hat{n}) / \hat{k}(\hat{n}) d\hat{n} - 6 \int \int \frac{1}{4\pi} / \hat{k}(\hat{n}) f(\hat{n}) d\hat{n}'$$

$$= \int \mathcal{R}(\hat{n}) / \hat{k}(\hat{n}) d\hat{n} \qquad (A-35)$$

and

$$\int_{\partial R} f(\hat{n}) y_{kn}^{(n)} d\hat{n} = \frac{1}{6\epsilon - 6\epsilon} R(\hat{n}) y_{kn}^{(n)} d\hat{n}$$
(A-36)

Rewriting Eq (36) with a lm spherical harmonic subscript and substituting into the expanded form of Eq (A-30) gives

$$R(\hat{n}) = 6\hat{f}(\hat{n}) - \left\{ \frac{6\hat{e}}{6\hat{e} - 6\hat{g}} \right\} - V(\hat{n}) \left\{ R(\hat{n}') \right\} V(\hat{n}') d\hat{n}' \quad (\Lambda - 37)$$

and by rearranging Eq (A-37)

$$f(\hat{n}) = \frac{1}{6e} \left[ R(\hat{n}) + \sum_{e,m} \left\{ \frac{G_e^9}{6e - G_e^9} \right\} \frac{V(\hat{n})}{e_m} R(\hat{n}) V_{em}^{*}(\hat{n}) d\hat{n} \right]$$
(A-38)

Comparing Eqs (A-3E) and (A-31) it is obvious that  $K^g$  is defined by

$$K^{S}R(\hat{n}) = \frac{1}{6\epsilon} \left[ R(\hat{n}) + \sum_{e,m} \left\{ \frac{6e^{s}}{\epsilon - 6e^{s}} \right\} \cdot y(\hat{n}) \left\{ R(\hat{n}) y(\hat{n}) d\hat{n} \right\} \right]$$
(A-39)

a similar derivation for  $K^{u}$  would produce

$$\mathcal{K}^{u}\mathcal{R}(\hat{n}) = \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{R}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{R}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{R}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{V}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{V}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{V}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \mathcal{V}(\hat{n}) \right]$$

$$= \frac{1}{6\epsilon} \left[ \mathcal{R}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right] \cdot \mathcal{V}(\hat{n}) \left[ \mathcal{V}(\hat{n}) + \sum_{k,m} \frac{6\epsilon^{u}}{6\epsilon - 6u} \right]$$

## Appendix B

## Weak-Form of the Functional

The functional whose Euler equation is the even parity Boltzmann equation of Chapter III is

$$F(u) = \begin{cases} \langle \hat{x}. \nabla u, K''(\hat{x}. \nabla u) \rangle + \langle u, G^3 u \rangle - 2\langle \hat{x}. \nabla u, K''S'' \rangle \\ -2\langle u, S^3 \rangle \end{cases} d\hat{r} + \begin{cases} \langle \hat{x}. \hat{n}/u^2 d\hat{n} \rangle d\hat{s} \end{cases}$$

$$(B-1)$$

where the inner product <f,g> is defined as

$$\langle f, g \rangle = \int_{a\pi}^{*} f(\hat{x}) g(\hat{x}) d\hat{x}$$
 (B-2)

and \* means the complex conjugate.

The minimizing function of this functional is the function  $\Psi$  which is a solution to the second order even parity Boltzmann equation (Ref 10:169) Therefore, a solution to the even parity Boltzmann equation can be found by minimizing Eq (B-1).

Another more useful formulation of this problem is the weak form. This weak form can be found by imposing the condition that a function which satisfies the natural boundary conditions Eqs (41) and (42) and the even parity Boltzmann euqation must also be a minimum of the functional (B-1).

Let the functional, (Eq (B-1), have a minimum at  $\Psi$ . Then, for all  $\eta$  and  $\epsilon$  where  $\epsilon$  can be arbitrarily small and of either sign

$$F(\Psi) \leqslant F(\Psi + \varepsilon \eta)$$
 (B-3)

where  $\forall$  and  $\eta$  are real functions that satisfy the boundary conditions.

Expanding Eq (B-1) in Ψ+εη gives

$$F(\Psi + \mathcal{E}_{\mathcal{T}}) = \int_{\mathcal{R}} \langle \hat{n} \cdot \nabla \Psi, K''(\hat{n} \cdot \nabla \Psi) \rangle dF \qquad (E-4)$$

$$+E$$
  $\langle \hat{n}. PN, K''(\hat{n}. P\Psi) \rangle d\hat{r}$  (B-6)

$$+\varepsilon^{2}\left\langle \hat{n}\cdot V\eta, \kappa''(\hat{n}\cdot V\eta)\right\rangle dF$$
 (B-7)

$$+ \int_{\mathcal{R}} \langle \Psi, G^{3} \Psi \rangle dF \tag{B-8}$$

$$+\varepsilon \int_{\mathbb{R}} \langle \Psi, G^{g} \eta \rangle dF$$
 (B-9)

$$+\varepsilon \int \langle n, G^3 \psi \rangle df$$
 (B-10)

$$+\mathcal{E}^{2}\langle n, G^{2}n\rangle d\hat{r}$$
 (B-11)

$$-2\left\langle \hat{n}\cdot\nabla\Psi_{2}K^{"}S^{"}\right\rangle df \qquad (B-12)$$

$$-2E \int \langle \hat{n}. Y n, k'' S'' \rangle d\hat{r}$$
 (B-13)

$$-2 \left\langle \Psi, S^{3} \right\rangle dF$$
 (B-14)

$$-2E \int_{\mathcal{R}} \langle \mathcal{H}, S^{8} \rangle dF \tag{B-15}$$

$$+ \oint \int /\hat{n} \cdot \hat{n} / \psi^2 d\hat{n} d\hat{s}$$
 (B-16)

$$+2\varepsilon \oint \int_{S} |\tilde{n}| |\Psi| \eta d\tilde{n} d\tilde{s}$$
 (B-17)

$$+ \varepsilon^{2} \int \left\langle \int \mathcal{E} \left\langle \tilde{\eta} \right\rangle \mathcal{D}^{2} d\tilde{\eta} d\tilde{s} \right\rangle$$
 (B-18)

noting that 
$$(B-4) + (B-8) + (B-12) + (B-14)$$

$$+ (B-16) = F(\Psi)$$
 (B-19)

and that  $K^g$ ,  $K^u$ ,  $G^g$  and  $G^u$  are self adjoint operators (Ref 10:174), where if L is self adjoint then

$$\langle Lf(\hat{n}), g(\hat{n}) \rangle = \langle f(\hat{n}), Lg(\hat{n}) \rangle$$

$$= \langle Lg(\hat{n}), f(\hat{n}) \rangle^*$$
(B-20)

\* means the complex congugate. Since  $\eta$  and  $\Psi$  are both real functions (B-5) = (B-6) and (B-9) = (B-10). Therefore, collecting terms and simplifying

$$F(\Psi + E \pi) = F(\Psi) + 2E \left\{ \left\{ \hat{\Lambda} \cdot \nabla \eta, K''(\hat{\Lambda} \cdot \nabla \Psi) \right\} \right\}$$

$$+ \left\{ \eta, G^{3} \Psi \right\} - \left\{ \hat{\Lambda} \cdot \nabla \eta, K'' S'' \right\} - \left\{ \eta, S^{3} \right\} d\hat{F}$$

$$+ \left\{ \left\{ \left| \hat{\Lambda} \cdot \hat{n} \right| \Psi \eta d\hat{\Lambda} d\hat{S} \right| + E^{2} \left\{ \left\{ \left| \hat{\Lambda} \cdot \nabla \eta, K''(\hat{\Lambda} \cdot \nabla \eta) \right| \right\} \right\}$$

$$+ \left\{ \eta, G^{3} \eta \right\} d\hat{F} + \left\{ \left| \left| \hat{\Lambda} \cdot \hat{n} \right| \eta^{2} d\hat{\Lambda} d\hat{F} \right| \right\}$$

$$+ \left\{ \eta, G^{3} \eta \right\} d\hat{F} + \left\{ \left| \left| \hat{\Lambda} \cdot \hat{n} \right| \eta^{2} d\hat{\Lambda} d\hat{F} \right| \right\}$$

$$(B-21)$$

With both  $K^u$  and  $G^g$  being positive definite operators then the  $\varepsilon^2$  term of (B-21) must also be positive. Note that  $\varepsilon^2$  is always positive. Now in order to ensure that  $F(\Psi + \varepsilon \eta) > F(\Psi)$  for  $\varepsilon \neq 0$ , the  $\varepsilon$  term in equation (B-21) must be positive or zero. But  $\varepsilon$  can be of either sign therefore the coefficient of  $\varepsilon$  must be zero, that is

$$\begin{cases} \langle \hat{n}. \nabla n, \kappa^{\alpha}(\hat{n}. \nabla \psi) \rangle + \langle n, G^{2}\zeta' - \hat{n}. \nabla n, \kappa^{\alpha}\zeta^{\alpha} \rangle \\ -\langle n, S^{3} \rangle \rbrace dF + \oint_{S} \langle |\hat{n}. \hat{n}| \psi \eta d\hat{n} d\hat{s} = 0 \end{cases}$$
(B-22)

Eq (B-22) is the weak or Calerkin form of the second order even parity Boltzmann equation. A detailed derivation of equation (B-22) can be found elsewhere (Rer 5:57).

## Appendix C

#### Derivation of the Weak Form From the Galerkin Method of Weighted Residuals

It can be shown that solutions to the second order forms of the Boltzmann equation, by using a variational principle or the method of weighted residuals are equivalent. A proof of this equivalency for the even parity equation is outlined below. However, a similar proof can also be extended to the odd parity second order equation.

The starting point of this proof is the even parity Boltzmann equation

$$-\hat{\chi} \cdot \nabla \chi''(r) \hat{\chi} \cdot \nabla \psi(\hat{r}, \hat{\chi}) + G^{3}(\hat{r}) \psi(\hat{r}, \hat{\chi}) = S^{3}(\hat{r}, \hat{\chi}) \quad (c-1)$$
$$-\hat{\chi} \cdot \nabla \chi''(\hat{r}) S''(\hat{r}, \hat{\chi})$$

and vacuum boundary conditions

$$\Psi(\hat{\xi},\hat{n}) + \kappa''(\hat{\xi}) \left\{ S''(\xi,\hat{n}) - \hat{n} \cdot \nabla \Psi(\hat{r}_{\xi},\hat{n}) \right\} = 0 \qquad (C-2)$$
for  $\hat{n} \cdot \hat{n} < 0$ 

$$\psi(\hat{s},\hat{\Lambda}) - K(\hat{s}) \left\{ S(\hat{s},\hat{\Lambda}) - \hat{\Lambda} \cdot P \psi(\hat{s},\hat{\Lambda}) \right\} = 0 \qquad (c-3)$$

$$\inf \hat{\Lambda} \cdot \hat{\Lambda} > 0$$

In the following equations the  $\hat{r}$  and  $\hat{\Omega}$  dependences will be omitted.

If a trial solution  $\Psi$  is assumed where  $\Psi$  is a linear combination of functions such that

$$\psi = \sum_{i=1}^{N} A_i \psi_i$$
 (c-4)

then the Galerkin method of weighted residuals requires a weight or test function  $\eta$ ; where

$$\gamma = \sum_{i=1}^{N} \psi_{i} \tag{c-5}$$

The requirement that  $\Psi$  should be an exact solution to the problem is imposed by substituting Eq (C-4) into (C-1), and then requiring that the Euclidean norm of the right and left hand sides of (C-1), with respect to the weight function  $\eta$ , are equal.

Applying this requirement to Eq (C-1) gives

where the inner product is defined as

$$\langle f, g \rangle = \int_{\pi} f g d\hat{n}$$
 (C-7)

and the trial and weight functions, Y and n are real functions.
Using the vector identity (Ref 10:169)

$$\begin{cases} \langle \hat{n}.\nabla n,f \rangle d\hat{r} = -\langle r.\hat{n}\nabla f \rangle d\hat{r} \\ + \phi \langle (\hat{n}.\hat{h})n,f \rangle d\hat{s} \end{cases} (c-s)$$

terms A and B of Eq (C-6) becomes

$$\int_{\mathcal{R}} -\langle \hat{n}. \nabla X'' \hat{n}. \nabla Y', \eta \rangle d\hat{F} = \int_{\mathcal{R}} \langle K'' \hat{n}. \nabla Y', \hat{n}. \nabla \eta \rangle d\hat{F}$$

$$- \oint_{\mathcal{S}} \langle (\hat{n}.\hat{n}) K'' \hat{n}. \nabla Y', \eta \rangle d\hat{S} \quad (C-9)$$

and

$$\int_{R} -\langle \hat{n} \cdot \nabla k \, {}^{\mu}S, n \rangle dF = \int_{R} \langle k \, {}^{\mu}S, \hat{n} \cdot \nabla n \rangle dF$$

$$- \oint_{R} \langle (\hat{n} \cdot \hat{n}) k \, {}^{\mu}S, n \rangle dS \qquad (C-10)$$

Substituting Eq (C-9) and (C-10) into (C-6) gives

$$\begin{cases}
\left\langle \hat{n}.\nabla\eta, \kappa^{\alpha}.\nabla\psi\right\rangle + \left\langle G^{3}\psi, \eta\right\rangle \right\} dF \\
+ \left\langle \left\langle \left(\hat{n}.\hat{n}\right)\kappa^{\alpha}S, \eta\right\rangle - \left\langle \left(\hat{n}.\hat{n}\right)\kappa^{\alpha}\hat{n}.\nabla\psi, \eta\right\rangle \right\} dS \\
= \left\langle \left\langle \hat{n}.\nabla\eta\right\rangle, \kappa^{\alpha}S, \psi\right\rangle + \left\langle \gamma, S^{S}\right\rangle dF \qquad (C-11)$$

term A of Eq (C-11) can be rearranged into

$$\widehat{A} = \oint \int (\widehat{n} \cdot \widehat{n}) \, K'' \left\{ S = \widehat{n} \cdot \nabla \, \Psi \right\} \, \widehat{\mathcal{I}} \, d\widehat{n} \, d\widehat{s} \qquad (C-12)$$

Using the boundary condition of Eq (C-2) and (C-3), Eq (C-12) can be written as

Eq (C-12) can also be written as

$$(c-14)$$

where || means the absolute value. Substituting Eq (C-14) into (C-11) gives

$$\begin{cases} \langle \hat{x}, \nabla \eta, K^{\mu} \hat{x}, \nabla \psi \rangle + \langle G^{3} \psi, \eta \rangle d\hat{r} + \delta \langle \hat{p} \hat{x} \hat{n} \rangle \psi \eta d\hat{x} d\hat{s} \\ = \int_{\mathcal{R}} \langle \hat{x}, \nabla \eta, K^{\mu} \hat{s}^{\mu} \rangle + \langle \eta, S^{3} \rangle d\hat{r} \qquad (c-15)$$

Eq (C-15) is the weak or Galerkin form of the even parity Boltzmann equation. It is identical to Eq (B-22) of Appendix B.

## Appendix D

## Expansion Properties of Spherical Harmonics

In Chapters III and IV the angular dependence of the trial functions and cross sections was expanded in spherical harmonics. Because of the two dimensional angular dependence in  $\mu$  and  $\chi$ , and the requirement that the expansion functions should form a complete set, the expansion is presented with m and  $\ell$  subscripts as follows (Ref 6:608)

$$f(\hat{n}) = \sum_{e=0}^{L} \sum_{m=-e}^{+e} f_{e,m} /_{e,m} \hat{n}$$
(D-1)

or

$$f(u, \chi) = \sum_{e=0}^{L} \int_{m=0}^{L} f_{e,m} / f(u, \chi)$$
(D-2)

where

$$f_{e,m} = \int_{-1}^{2\pi} f(u, \chi) \int_{e_{m}}^{*} f(u, \chi) du d\chi$$
 (D-3)

and

$$V(u,\chi) = C_{em} P_{em}(u) e^{im\chi}$$

$$(D-4)$$

and

$$Y(a, \chi) = (-1)^m y_{e,m}^*(a, \chi)$$
 (D-5)

and

$$C_{pm} = \sqrt{\frac{2\ell+1 \cdot (\ell-m)!}{4\pi \cdot (\ell+m)!}}$$
 (D-6)

If  $f(\mu,\chi)$  is even in the angle  $\chi$  then the expansion must also be even in  $\chi$ , and therefore, (D-4) can be rewritten as

$$Re\left[Y_{l,m}\right] = Q_{l,m} = C_{l,m} P_{l,m}(u) Cos(m\chi)_{(D-7)}$$

where the odd iSinx term is omitted. Also from Eq (D-5)

$$Q_{\ell,-m} = (-1)^m C_{\ell,m} P_{\ell,m}(u) cos(m + ) = (-1)^m Q_{\ell,m}$$
 (D-8)

Therefore for an even expansion Eq (D-3) can be written as

$$f_{e,m} = \int_{0}^{2\pi} \int_{-\infty}^{\infty} f(u,\chi) G_{e,m}(u,\chi) du d\chi \qquad (D-9)$$

and

$$+_{1-m} = \int_{0}^{2\pi(r+1)} (-1)^{m} f(u,\chi) Q_{\ell,m}(u,\chi) du d\chi = (-1)^{m} f_{\ell,m}(D-10)$$

For an even  $\chi$  expansion of  $f(u,\chi)$  Eq (D-2) can therefore be written as

$$f(u,\chi) = \sum_{\ell=0}^{L} \left[ f_{\ell 0} Q_{\ell 0} + 2 \sum_{m=1}^{\ell} f_{\ell,m} Q_{\ell,m} \right]$$
 (D-11)

or

$$f(u\chi) = \sum_{l=0}^{L} f_{l,m} Q_{l,m}$$
(D-12)

omero the factor of two has been included in the  $t_{\rm c,m}$  expansion coefficients and Eq. (0-12) forms a complete set.

A similar approach can be used for the cross-section expansion

$$6^{s}(\hat{n}\cdot\hat{n}') = \sum_{e=0}^{l} \frac{2e+i}{4\pi} \cdot 6^{s} P_{e}(\hat{x}\cdot\hat{n}')$$

$$= \sum_{e=0}^{l} \frac{1}{4\pi} \cdot 6^{s} P_{e}(\hat{x}\cdot\hat{n}') V(\hat{n})$$

$$= \sum_{e=0}^{l} \frac{1}{4\pi} \cdot 6^{s} P_{e}(\hat{x}\cdot\hat{n}') V(\hat{n}) \qquad (D-13)$$

where  $P_{\ell}(\hat{\Omega} \cdot \hat{\Omega}')$  is given by Eq (A-24).

Eq (D-13) can be expanded to give

$$b(\hat{x}, \hat{x}) = \sum_{e=0}^{L} \left( \frac{1}{e} y(\hat{x}) y(\hat{x}) + \sum_{e=0}^{e} \frac{1}{e} y(\hat{x}) y(\hat{x}) \right) + \sum_{e=0}^{e} \frac{1}{e} y(\hat{x}) y(\hat{x}) + \sum_{e=0}^$$

and from Eq (D-5)

$$V_{s,m}^{*}(\hat{n}) V_{s,m}(\hat{n}) = V_{s-m}^{*}(\hat{n}) V_{s,m}(\hat{n})$$
(D-15)

Therefore (D-14) becomes

$$b(\hat{A} \cdot \hat{A}') = \sum_{\ell=0}^{2} \left[ c_{5}^{5} \sqrt{(\hat{A})} / (\hat{A}) + 2 \sum_{\ell=0}^{2} c_{5}^{5} \sqrt{(\hat{A})} / (\hat{A}) \right]$$

$$= 2 \sum_{\ell=0}^{2} \left[ c_{5}^{5} \sqrt{(\hat{A})} / (\hat{A}) \right]$$

$$= 2 \sum_{\ell=0}^{2} \left[ c_{5}^{5} \sqrt{(\hat{A})} / (\hat{A}) \right]$$
(D-16)

where m means that all terms with a m = 0 subscript must be divided by two.

Since the scattering cross-sections are real the expansion of Eq (D-16) must also be real and with

Eq (D-16) can be written as

$$\widetilde{O(11)} = 2 \sum_{n=0}^{\infty} \sum_{n=0}^{\infty} \frac{s-2}{n!} \widetilde{O(12)} \widetilde{O(12)} = 2 \sum_{n=0}^{\infty} \sum_{n=0}^{\infty} \frac{s-2}{n!} \widetilde{O(12)} \widetilde{O(12)} = 2 \sum_{n=0}^{\infty} \frac{s-2}{n!} \widetilde{O(12)} = 2 \sum_{n=0}^{$$

Note that

$$los m(\chi-\chi') = Cos m\chi \cdot los m\chi' + Sin m\chi \cdot sin m\chi'$$
(D-19)

The angles  $\mu$  and  $\chi$  are shown in Fig. 2.

Similar derivations would give

$$\delta_s^g(\hat{\mathbf{a}}.\hat{\mathbf{a}}') = 2 \sum_{k=0}^{L} \int_{m'=0}^{g} \mathcal{C}_{km}^g P_{km}(\mathbf{u}) P_{km}(\mathbf{u}) \cos m(\chi - \chi')$$
(D-20)

and
$$\mathcal{C}_{s}^{u}(\hat{\Lambda}:\hat{\Lambda}') = 2 \sum_{\ell=0}^{L} \mathcal{C}_{\ell}^{u} \mathcal{C}_{\ell}^{2} \mathcal{P}_{\ell m}(u') \mathcal{P}_{\ell m}(u) \cos m(\gamma - \gamma')$$
(D-21)

By inserting Eq (D-20) into Eq (A-22) the even parity collision operator can now be written as

and from Eqs (D-15) and (D-17) the inverse odd parity collison operator Eq (A-40), can be written as

$$K^{\mu} = G_{2}^{\mu} \begin{bmatrix} 1 \\ \frac{1}{2} & \frac{1}{2} \\ $

Similar empression for all and to early be written.

Come unclude properties of advanted formenies and the associated Lemmondre functions are (Met. 18: 5:509)

$$P(u) = (-1)^{m} \frac{(\ell-m)!}{(\ell+m)!} P(u)$$

$$(D-25)$$

$$\begin{cases} P(\omega)P(\omega)d\omega = \frac{2}{2e+1} \cdot \frac{(e+m)!}{(e-m)!} & \text{Se} \end{cases}$$
 (D-26)

where

$$\int_{ee'} = \begin{cases}
0 & \text{for } \ell \neq 0 \\
1 & \text{for } \ell = 0
\end{cases}$$
(D-27)

$$\int_{0}^{2\pi} \int_{-1}^{+1} \langle u, y \rangle \chi_{em}^{*}(u, y) du dx = \int_{0}^{2\pi} \int_{0}^{+1} \int_{0}^{1} \int_{0}$$

#### Appendix E

## The Synthesized Boltzmann Equation

In order to formulate a numerical solution to the air-over-ground problem it is necessary to expand the weak form of the even parity Boltzmann equation in a set of trial and test functions. The finite element space-angle synthesis trial and weight functions of Chapter IV will be used in this expansion. Because this is a very tedious and lengthy derivation, the more obvious albegraic steps will be omitted. The starting point of this formulation is the synthesized even parity Boltzmann equation, Eq (58), of Chapter III.

$$\frac{12}{|\mathcal{I}|} \frac{1R}{|\mathcal{I}|} \frac{1}{|\mathcal{I}|} $

where  $B_{i}$  and  $B_{j}$  are tensor products of cubic polynomial splines, and they are given by

$$B_{p} = B_{z}(z) B_{z}(P)$$
 (E-2)

$$B_{j} = B_{j}(z)B_{j}(P) \tag{E-3}$$

also

$$Q = C P(u) \cos m\chi$$
(E-4)

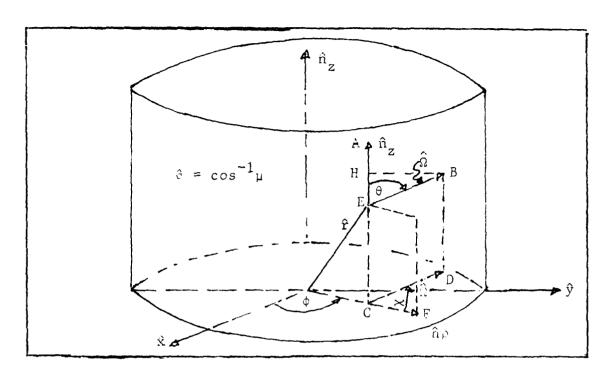


Figure 9. Surface Normal And Particle Velocity Direction Vectors.

and

$$C_{2m} = \sqrt{\frac{2\ell+1}{4\pi}} \frac{(\ell-m)!}{(\ell+m)!}$$
 (E-5)

 $P_{lm}(\mu)$  are the associated Legendre functions.

The directional derivative is given as

$$\hat{\Lambda} \cdot \nabla \phi = \frac{\partial}{\partial \rho} (\rho \phi) - \frac{1}{\rho} \partial \frac{\{\phi B(u, \chi)\}}{\partial \chi} + u \partial \phi \qquad (E-6)$$

where

$$\mathcal{L} = \alpha(\mathcal{U}, \mathcal{V}) = \sqrt{1 - \mathcal{U}^2} \cos \mathcal{V}$$
 (E-7)

and

$$\beta = \beta(u, \chi) = \sqrt{1 - u^2} \sin \chi \qquad (E-8)$$

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$$\hat{\mathcal{N}} \cdot \hat{\mathbf{n}} = \begin{cases} \hat{\mathcal{N}} \cdot \hat{\mathbf{n}}_{\pm} & \text{for the horizontal top and bottom surfaces, and} \\ \hat{\mathcal{N}} \cdot \hat{\mathbf{n}}_{p} & \text{for the vertical (side) of the problem cylinder} \end{cases}$$
(E-9)

n aning are shown in Fig. 7.

If  $\boldsymbol{\hat{\Omega}}$ ,  $\boldsymbol{\hat{n}}_{r}$  and  $\boldsymbol{\hat{n}}_{z}$  are considered to be unit vectors then from Fig. 7

$$AE = EB = CF = 1$$
(E-10)

and therefore for the top surface of the cylinder

$$\hat{\mathcal{L}} \cdot \hat{n}_{Z} = AE \times EB \times Cos\theta = \mathcal{L}$$
(E-11)

and for the bottom surface

$$\lambda \cdot n_z = -u$$
 (E-12)

also CD = HB = projection of  $\hat{\Omega}$  unto the x-y plane and therefore,

CD = HB = EB Sin0

$$= \sqrt{1 - \cos^2 \theta} = \sqrt{1 - \mu^2}$$
 (E-13)

and

$$\hat{\lambda} \cdot \hat{n}_{\rho} = C \square \times CF \times Cos \chi$$

$$= \sqrt{1 - \mu^2 \cos \chi}$$
(E-14)

In Appendix D the even and odd operators were given as

$$K^{\mu}f(\hat{x}) = 6_{t}^{-1} \left[ f(\hat{x}) - 2 \sum_{\ell=0}^{L} \sum_{m=0}^{\ell} \frac{\{\omega_{\ell} u\}}{(\omega_{\ell} - \omega_{\ell} u)} \right]_{4\pi} T(\hat{x}') f(\hat{x}') d\hat{x}' \right]$$
(E-15)

and

$$G^{g}f(x) = 6 f(x) - 2 \sum_{\ell=0}^{\infty} \int_{m^{\ell}=0}^{g} \int_{A_{\ell}} T(x) f(x) dx.$$
(E-16)

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where  $m^*$  means that all terms with a m = 0 index <u>must be</u> divided by two, and

$$T_{em}(\hat{n}) = C_{em} P(u) P_{em}(u) \cos m(x-x')$$
 (E-17)

Also the inner product is defined as

$$\langle f, g \rangle = \int_{4\pi} f g d\hat{n}$$
  
=  $\int_{4\pi} f g d\hat{n}$  for real f (E-18)

Using Eqs (E-16), (E-15), (E-14), (E-11), (E-12) and (E-6) and noting that  $T_{\xi m}$ ,  $Q_{\xi m}$ ,  $B_i$  and  $B_j$  are all real functions, Eq (E-1) can be expanded to give

$$\begin{array}{c|c}
IZ IR L & P \\
\hline
 & P$$

$$-\left(\frac{1}{2}\frac{\partial}{\partial P},P\right)\frac{B_{i}}{\partial P}dV\int_{A}\frac{\partial}{\partial V}\left(\frac{BQ}{Q}\right)Q_{k}d\hat{\Lambda} \qquad (E-20)$$

$$-\int_{\mathcal{C}} \frac{\partial (PB_i) \cdot B_i}{\partial P} \frac{\partial (PQ) \cdot Q}{\partial X} \frac{\partial (PQ) \cdot Q}{\partial X} \frac{\partial (PQ)}{\partial X} \frac{\partial (PQ)}{\partial X}$$
(E-22)

$$+ \int_{\mathcal{R}} 6^{-1/2} \mathcal{B} \cdot \mathcal{B} \cdot d\nu \int_{\mathcal{A}\mathcal{R}} \frac{\partial (\mathcal{B}Q_{\mathbf{A}})}{\partial \chi} \cdot \frac{\partial (\mathcal{B}Q_{\mathbf{A}})}{\partial \chi} d\hat{n} \qquad (E-23)$$

$$-\int_{\mathcal{R}} \frac{\partial}{\partial z} \frac{\partial}$$

+ 
$$\left(6\frac{1}{2}(PB_i).\frac{3}{32}(B_i).\frac{1}{1}dv\right)$$
  $uQ.aQd\hat{n}$  (E-25)

$$-\int_{\mathcal{R}}^{6^{-1}} \frac{\partial (B_j) \cdot B_i}{\partial Z} \frac{\partial v}{\partial z} \int_{\mathcal{A}\pi} \frac{\partial (B_j) \cdot \partial \tilde{x}}{\partial z} (E^{-26})$$

$$+2\sum_{r=0}^{R}\sum_{s=0}^{r}\left\{\begin{pmatrix}g_{i}^{u}\frac{\partial}{\partial r}(\rho B_{i})\frac{\partial}{\partial r}($$

$$-\int_{\mathcal{E}} \frac{1}{\rho^2} \frac{B}{\partial \rho} \frac{\partial}{\partial \rho} \left( \rho B_j \right) dv \left( \Delta Q_j \left\{ \int_{\Delta \pi} \frac{\partial}{\partial x} \left( \beta Q_j \right) T_j d\hat{\pi}^i \right\} d\hat{\pi} \right) (E-29)$$

+ 
$$\left(6\frac{3}{6}\frac{\partial(B_i)}{\partial z}\frac{\partial(PB_i)}{\partial p}\right)\int_{A}^{B} dv \left(2\left(\frac{1}{2}\right)\frac{\partial v}{\partial x}\right)\int_{A}^{B} d\tilde{x}^{2}d\tilde{x}^{2}d\tilde{x}^{3}\right)d\tilde{x}$$
 (E-30)

$$-\int_{\mathcal{R}^{\mu}} \int_{\partial \mathcal{D}} B_{i} \frac{\partial}{\partial \rho} (\rho B_{i}) d\nu \int_{\partial \mathcal{D}} \frac{\partial}{\partial x} g_{\mu\nu} \left\{ \int_{\partial \mathcal{D}} \mathcal{D}_{FS} d\hat{n}^{i} \right\} d\hat{n} \qquad (E-31)$$

$$-\int_{\mathcal{R}} \frac{\partial^{n} \partial(\mathcal{B}_{i})}{\partial \mathcal{I}} \frac{\partial}{\partial \mathcal{I}} d\nu \int_{\mathcal{A}} \frac{\partial^{n} \partial\mathcal{I}}{\partial \mathcal{I}} \frac{$$

$$+ \left( \underbrace{6^{"} \underline{\partial (B_{i})}}_{E_{i}} \underbrace{\partial (B_{i})}_{\partial Z} \underbrace{\partial V \left( \underline{u} \underbrace{Q}_{i} \underbrace{S} \underbrace{\partial V_{i}}_{A_{i}} \underbrace{\partial A_{i}}_{A_{i}} \underbrace{\partial A$$

$$+ \begin{cases} 6 \frac{\delta(B_i)}{\delta(B_i)} \frac{\delta(B_i)}{\delta(B_i)} dv \begin{cases} n & \text{if } dn \\ n & \text{if } dn \end{cases} \begin{cases} n & \text{if } dn \\ n & \text{if } dn \end{cases}$$

$$(E-36)$$

$$+ \begin{cases} 6B_{i}B_{i}dv \end{cases} Q_{i}Q_{i}dx$$
(E-37)

$$-2\sum_{t=0}^{R}\sum_{s=0}^{r}\left\{6, B_{j} B_{i} dv\right\} \mathcal{Q}_{en}\left\{3 \mathcal{Q}_{en} \mathcal{Q}_$$

$$= \int_{\mathcal{R}} \frac{\partial (PB_j)}{\partial P} d\nu \left( \mathcal{A} \mathcal{G} \mathcal{S} \mathcal{A} \mathcal{A} \right)$$
(E-40)

$$-\int_{\mathcal{R}} \frac{1}{2\pi} \frac{1}{2\pi} \frac{1}{2\pi} \int_{\mathcal{A}} \frac{1}{2\pi} \frac{1}{2\pi} \frac{1}{2\pi} \int_{\mathcal{A}} \frac{1}{2\pi} \frac{1}{2\pi} \frac{1}{2\pi} \frac{1}{2\pi} \int_{\mathcal{A}} \frac{1}{2\pi} $

$$+ \int_{\mathcal{R}} \frac{\delta'}{\partial z} \frac{\partial}{\partial z} (B_j) dv \left( u \mathcal{G}_{kn} S'' \mathcal{G}_{n} \right)$$

$$(E^{-12})$$

$$+2\sum_{r=0}^{R}\sum_{s=0}^{r}\left(\frac{6u}{6\pi}\frac{\partial(rB)}{\partial r}\frac{\partial v}{\partial r}\right)dv\int_{a\pi}\left(\frac{Su}{6\pi}\frac{\partial v}{\partial r}\right)dx$$

$$(E-4u)$$

$$-\left(\frac{6}{2}\frac{B_{i}}{P}dv\right)\left(\frac{\partial}{\partial x}\left(BQ_{in}\right)\right)\left(\frac{\partial}{\partial x}\left(BQ_{in}\right)\right)\left(\frac$$

$$+ \int_{\mathcal{R}} \frac{\partial}{\partial z} \frac{\partial}{\partial z} |dv| \int_{\partial T} u \int_{\partial T} \left\{ \int_{\partial T} \frac{\partial}{\partial z} dz \right\} dz$$
(E-45)

$$+\int_{\mathcal{R}} B_{i} d\nu \int_{\mathcal{A}\mathcal{T}} S^{3} d\hat{x}$$
 (E-46)

where

$$G_{Er}^{u} = G_{e}^{-1} \cdot \frac{G_{r}^{u}}{G_{e}^{-G_{r}^{u}}}$$
 (E-47)

and

$$\int_{4\pi} d\hat{n} = \int_{0}^{2\pi} \int_{0}^{1} du dz \tag{E-48}$$

Eqs (E-19) to (E-36) is an expansion of the first term of Eq (E-1). Eqs (E-37) and (E-38) is an expansion of the second term. Eq (E-39) is an expansion of the third term. Eqs (E-40) to (E-46) is an expansion of the right hand side of Eq (E-1).  $\sigma_{\rm r}^{\rm g}$ ,  $\sigma_{\rm t}$  and  $\sigma_{\rm tr}^{\rm u}$  are functions of z and they must be included in the spatial or dv integrals. In cylindrical geometry with azimuthal symmetry

$$dV = 2\pi r dr dz \tag{E-49}$$

and fds means an integration over the surface of the problem cylinder, Fig. 9.

For the air-over-ground problem with an exponentially varying air density

$$G(z) = G(0)e^{-2/sh}$$
 (E-50)

$$G_{e}^{u}(z) = G(0)e^{+2/sh}$$
 (E-51)

$$6_{1}^{3}(2) = 6_{1}^{3}(0)e^{-2/5h}$$
 (E-52)

where  $\sigma_t(o)$ ,  $\sigma_r^g(o)$ ,  $\sigma_{tr}^u(o)$  are cross-sections of air at sea level.

Z = the height above sea-level

sh = atmospheric scale height ~ 7km

In Eqs (E-19) to (E-46) the integrals are separated in the space and angle variables. These are double integrals in space and angle. However, they can be separated into single integrals of the  $\mu$ ,  $\chi$ ,  $\rho$  and z variables.

## Appendix F

#### Angle Integrals of the Synthesized Second Order Boltzmann Equation

An expansion of the even parity Boltzmann equation has produced twenty-eight integral terms (Appendix E). By a further expansion and separation of the integration variables twenty distinct single angle integrals are formed. These angle integrals are dependent on the degree of the spherical harmonic trial function expansion and independent of the problem parameters. They can be evaluated once and thereafter used as a part of the problem input data. In this research project these integrals were numerically integrated for each combination of the £, m, k and n expansion subscripts. They were then stored as a matrix, and selected products were used to produce each of the twenty-eight angle integrals of Eqs (E-19) to (E-46) in Appendix E.

These twenty integrals are

$$\int_{0}^{2\pi} \cos^{2}(x)\cos(mx)\cos(nx) dx \qquad (F-1)$$

$$\begin{cases}
\cos(x) & \frac{\partial}{\partial x} \left\{ \sin(x) \cos(mx) \right\} \cos(nx) dx \\
& \frac{\partial}{\partial x} \left\{ \sin(x) \cos(mx) \right\} \cos(nx) dx
\end{cases}$$
(F-2)

$$\int_{0}^{2\pi} \frac{\cos(\chi)\cos(m\chi)\cos(n\chi)d\chi}{(F-3)}$$

$$\int_{C}^{2\pi} \cos(nx) dx \left\{ \sin(x) \cdot \cos(nx) \right\} dx$$

```
[ 2π
Cos(mx) cos(hx)dx
                                                                              (F-6)
    \begin{cases} 2\pi \\ (\cos(\chi)\cos(n\chi)\sin(m\chi)d\chi \end{cases}
                                                                             (F-7)
    (2T
Cos (nx)·Sin (mx)dx
                                                                             (F-8)
    \int_{0}^{2\pi} \frac{1}{dx} \left\{ \cos(n\chi) \sin(\chi) \right\} \sin(m\chi) d\chi
                                                                             (F-9)
   ("2
Cos(x)cos(mx)cos(nx)dx
                                                                             (F-10)
  \begin{cases} 3\pi/2 \\ \pi_{L}Cos(\chi) cos(m\chi) cos(n\chi) d\chi \end{cases}
                                                                             (F-11)
   \int_{3\pi/2}^{2\pi} \cos(\pi)\cos(m\chi)\cos(n\chi)d\chi
                                                                             (F-12)
   I u Pem (u) Pro (u) du
                                                                             (F-13)
  Ju Pen (u) Pyn (u) du
                                                                             (F-14)
 \begin{cases} (1-u^2)P_{kn}(u)P(u)du\\ \ell m \end{cases}
                                                                            (F-15)
J. VI-u2. M. Pem (M) Pr (M) du
                                                                            (F-16)
S-1 VI-u2. Pen(u) Pkn (u) du
                                                                            (F-17)
\int_{-\infty}^{\infty} u^2 P_{\mu}(u) P_{\mu}(u) du
                                                                            (F-1S)
```

$$\begin{cases}
u P_{f}(u) P_{k}(u) du \\
P_{f}(u) P_{k}(u) du
\end{cases}$$
(F-19)

A numerical evaluation of all twenty integrals was carried out for a third degree spherical harmonic expansion. This evaluation showed that these integrals are equal to zero for many combinations of the lm and kn subscripts.

Integrals (F-10) to (F-14) are a part of the surface integral term which has been partitioned into the outward  $(\hat{\Omega} \cdot \hat{n} > 0)$  and inward  $(\hat{\Omega} \cdot \hat{n} < 0)$  directions. This partitioning was incorporated into the weak form derivation of Appendix C.

## Appendix G

Space Integrals (Bicubic Splines) of the Synthesized Second Order Even Parity Boltzmann Equation

A trial function expansion of the spatial flux dependence, in the weak form of the even parity Boltzmann equation, has been carried out in Appendix E. Bicubic polynomial splines in the p and z variables were used to form a tensor product space. These splines are twice continuously differentiable and have non-zero integrals  $\int_R B_i(x)B_i(x)dx$  for all  $|i-j| \ge 4$ . After a separation of the p and z variables of integration seventeen distinct integral forms are produced. These integrals, which include the source integrals, must be evaluated over the entire problem domain. The space integrals of Appendix E are selected products of the following seventeen single integrals.

$$\int_{A}^{A} \frac{\partial \left\{ PB(P) \right\}}{\partial P} \cdot \frac{\partial \left\{ PB(P) \right\}}{\partial P} dP \tag{G-1}$$

$$\int_{0}^{R} \frac{\partial \{PB(P)\}}{\partial P} \cdot \underbrace{B_{i,r}(P)}_{P} dP \qquad (G-2)$$

$$\int_{P}^{R} B_{jr}(P)B_{jr}(P) dP \qquad (G-3)$$

$$\begin{cases} \frac{\partial}{\partial P} \left\{ P B_{3}(P) \right\} \cdot B_{j, k}(P) dP \qquad (G-4) \end{cases}$$

$$\begin{cases}
\mathcal{B}_{i}(P) \cdot \mathcal{B}_{i}(P) d\rho \\
\end{cases} (G-5)$$

$$\begin{pmatrix} \mathcal{B}_{\mathcal{A}}(P), \mathcal{B}_{\mathcal{A}}(P) dP & (G-6) \end{pmatrix}$$

$$\begin{cases}
R \\
B_{j}(P) \cdot B_{j}(P) dP
\end{cases} (G-6)$$

$$\begin{cases}
C + C \\
C + C
\end{cases} (G-7)$$

$$\int_{0}^{z} e^{\frac{z}{2}} \frac{\partial}{\partial z} \left\{ B_{iz}^{(z)} \right\} \cdot B_{iz}^{(z)} dz \qquad (G-8)$$

$$\begin{cases} \frac{\partial}{\partial z} \left\{ B_{i,2}^{(2)} \right\} \cdot \frac{\partial}{\partial z} \left\{ B_{i,2}^{(2)} \right\} e^{\frac{2}{3} 5 h} dz \qquad (G-9)$$

$$\int_{0}^{H} B_{i,z}(z) B_{j,z}(z) e^{-z/sh} dz$$
 (G-10)

$$\int_{a}^{H} B_{j}(z) \cdot B_{j}(z) dz \tag{G-11}$$

## The Source Integrals

$$\int_{0}^{R} \left\{ PB_{jr}(A) \right\} H_{j}(P) dP \tag{G-12}$$

$$\begin{pmatrix} B_{jr}(P) & H_{j}(P) & dP \\ \end{pmatrix} \tag{G-13}$$

$$\int_{0}^{R} B_{s}(P) H_{s}(P) P dP \qquad (G-14)$$

$$\int_{0}^{H} \frac{B(z)}{J^{2}} H(z) e^{\frac{z}{J}sh} dz \qquad (G-15)$$

$$\begin{cases} \frac{\partial}{\partial z} \left\{ B_{jz}^{(z)} \right\} H_{j}(z) e^{-\frac{z}{3}} dz \qquad (G-16) \end{cases}$$

$$\int_{0}^{H} \frac{B(z) \cdot H_{i}(z) dz}{(G-17)}$$

where

B(z) = cubic polynomial z-spline

 $B(x) = cubic polynomial \rho-spline$ 

sh = atmospheric scale height

R = outer radius of the problem cylinder

H = problem cylinder height

H(z) and  $H(\rho)$  are the source interpolating functions (linear Lagrange polynomials).

## Appendix H

# An Expansion of the First Scatter Source in Legendre Polynomials

In Chapter IV the first scatter source was defined as

$$S(P,Z,\hat{n}) = \delta(Z,\hat{n}.\hat{n}') \phi(P,Z,\hat{n}')$$
(H-1)

where

$$\phi_{d}(\hat{r}, z, \hat{\Omega}') = \text{direct fluence of Chapter IV, Eq (74)}$$

$$\sigma^{S}(z, \hat{\Omega} \cdot \hat{\Omega}') = \text{scattering cross-section}$$

The usual Legendre polynomial cross-section expansion will now be carried out. Also the even and odd parity first scatter source expressions of Chapter IV will be derived. Expanding  $\sigma^{\rm S}$  in Legendre polynomials and using the addition theorem (see Appendix D)

$$6^{s}(z,\hat{n}.\hat{n}')=2\sum_{\ell=0}^{L}\int_{m=0}^{\ell}6^{s}(z)P(u)P(u)\cos(m(x-x'))$$
 (H-2)

where m\* means that all terms with a m = 0 subscript must be divided by two, and

$$G_{\ell}^{s}(z) = G_{\ell}^{s}(0) e^{-\frac{z}{s}h}$$
 (11.12)

From Fig. 5 and Fig. 2 it is apparent that  $\mu$  =  $\mu$ d and  $\chi$  = 0

Therefore

$$S(P_3Z,\hat{\Lambda}) = 2 \phi_d(P_3Z,\hat{\Lambda}') e^{-\frac{Z}{5}h}$$

$$\times \sum_{\ell=0}^{L} \frac{\ell}{m^{\ell}} e^{s}(0) C_{\ell m}^2 P_{\ell m}(ud) P(u) \cos m \chi \quad (H-4)$$

and using the identity (Ref 18:96)

$$P(-u) = (-1)^{e-m}P(u) \tag{H-5}$$

and a little algebra, the even and odd parity first scatter sources can be written as

$$S''(\beta,Z,\hat{\Lambda}) = \frac{1}{2} \left[ S(\beta,Z,\hat{\Lambda}) - S(\beta,Z,-\hat{\Lambda}) \right]$$

$$= \phi_{\beta}(\beta,Z,\hat{\Lambda}') e^{-\frac{2}{3}S} \int_{e^{-2}}^{e^{-2}} \left[ \frac{1}{2} \left( \frac{1}{H^{-6}} \right) \right]$$

$$= (-1)^{e-m} P(x) P(x) Cos(m x)$$

and

$$S(P,Z,\hat{n}) = \frac{1}{2} \left[ S(P,Z,\hat{n}) + S(P,Z,-\hat{n}) \right]$$

$$= \Phi_{d}(P,Z,\hat{n}) e^{\frac{24sh}{2}} \sum_{e=0}^{4} \sum_{m'=0}^{6} \delta(e) C_{em}^{2} \left[ 1 + (-1)^{e-m} P(u) P(ud) \cos(mx) \right]$$
(H-7)

## Appendix I

## A Derivation of the Total Particle Fluence

In Chapter IV a trial function expansion of the even parity angular particle fluence was given as

$$\Psi(P,Z,u,y) = \sum_{|Z=1|} \frac{IZ}{ir} \frac{IR}{l=0} \frac{l}{m=0} \underbrace{A. B_i(z)B(P)Q}_{iy,\ell,m}$$
(I-1)

where

$$\mathcal{G} = \mathcal{C} \mathcal{P}(\mathcal{U}) \mathcal{C}os \, m \, \chi \qquad (I-2)$$

and

$$C_{em} = \sqrt{\frac{2\ell+1.(\ell-m)!}{4\pi(\ell+m)!}}$$
 (I-3)

The  $A_{i,j,\ell,m}$  mixing coefficients are obtained from a numerical solution of the second order synthesized Boltzmann equation of Chapter IV.

The angular even parity fluence is also defined as

$$Y(RZ, u, x) = \frac{1}{2} \left[ \phi(RZ, u, x) + \phi(RZ, -u, -x) \right]$$
 (1-4)

An integration of Eq (I-4) over all directions gives the total even parity particle fluence  $\Psi(\rho,z)$ , and also the total particle fluence  $\Phi(\rho,z)$ . This is because

$$\int \mathcal{Y}(P_3 Z, u, \chi) d\hat{n} = \mathcal{Y}(P_3 Z) \tag{1-5}$$

$$= \frac{1}{2} \int_{a\pi} \Phi(\beta z, u, x) d\hat{n} + \frac{1}{2} \int_{a\pi} \Phi(\beta z, -u, -x) d\hat{n}$$

$$= \int_{a\pi} \Phi(\beta z, u, x) d\hat{n} = \Phi(\beta, z)$$

Therefore

$$\psi(P_3Z) = \phi(P_3Z) \tag{I-6}$$

Eq (I-1) will now be integrated to give the total particle fluence at position  $(\rho,z)$ . Using the orthogonal properties of Legendre polynomials this integration is carried out as follows.

The zero order associated Legendre function is defined

$$P_{0,0}(u) = 1 \tag{1-7}$$

Multiplying Eq (I-1) by Eq (I-7) and integrating over all  $\rho$  and  $\chi$  directions gives

$$\begin{cases} 2\pi (|y(P_{2}z,u,x)|P_{00}(u) du dx \\ = |y(P_{2}z)| = \\ |iz| |it| = 0 \text{ m=0} \end{cases}$$

$$= |y(P_{2}z)| = |z| |it| = 0 \text{ m=0}$$

$$\times \begin{cases} 2\pi (|x|) & P_{00}(u) du dx \\ P_{00}(u) du dx \end{cases}$$

$$(1-8)$$

Substituting Eq (I-2) for  $Q_{\ell m}$ , the integral of Eq (I-8)

becomes

$$\int_{0}^{2\pi} \cos(m x) dx \cdot \int_{-1}^{\infty} C_{nm} P(u) P(u) du \qquad (1-9)$$

$$\int_{0}^{2\pi} \cos(m\chi) d\chi = \begin{cases}
0 & \text{for m } \neq 0 \\
2\pi & \text{for m} = 0
\end{cases} (I-10)$$

and (see Appendix D)

$$\int_{C_{e,o}}^{C_{e,o}} P_{e,o}(u) P_{e,o}(u) du = \frac{2}{2\ell+1} \cdot \frac{(\ell+o)!}{(\ell-o)!} \cdot \frac{S_{e,o}}{2\ell-1} C_{e,o}$$

$$= 2C_{o,o} = \frac{2}{4\pi} (I-11)$$

therefore Eq (I-9) becomes

$$2\pi \cdot 2/\sqrt{4\pi} = \sqrt{4\pi}$$
 (I-12)

and the total particle fluence is

$$\phi(\mathcal{C}_{S}Z) = \psi(\mathcal{C}_{S}Z) = \sqrt{4\pi} \sum_{\substack{i \geq 1 \ i \neq j \leq m \ i \geq i}} P_{i} B_{j}(z) B(P) \quad (1-13)$$
where  $f_{i} = m = 0$ 

The angular particle fluence is given by

$$\phi(P,Z,\hat{n}) = \psi(P,Z,\hat{n}) + \chi(P,Z,\hat{n}) \qquad (I-14)$$

where  $\chi(\rho,z,\hat{\Omega})$  is the odd parity fluence, which, is defined in terms of the even parity fluence and source by the following expression

$$\chi(\beta z, \hat{\lambda}) = k (\beta z) \left[ S(\beta z, \hat{\lambda}) - \hat{\lambda} \cdot \nabla \psi(\beta z, \hat{h}) \right]$$
(1-15)

Therefore once  $\Psi(\rho,z,\hat{\Omega})$  has been found the odd parity fluence and the angular particle fluence  $\varphi(\rho,z,\hat{\Omega})$  can be computed from Eqs (I-15) and (I-14).

#### Appendix J

#### Computer Subroutines

A computer program has been written to solve the synthesized Boltzmann equation of Chapter IV. This program is designed to use a trial function expansion in bicubic splines and spherical harmonics, and to perform a first scatter source interpolation using linear Lagrange polynomials. The program in an assemblage of several subroutines which collectively perform the following tasks.

- 1. Computes all single space and angle integrals.
- 2. Combines the single  $\mu$  and  $\chi$  angle integrals for all combinations of the spherical harmonic expansion subscripts.
- 3. Combines the single o and z integrals for all combinations of the bicubic spline expansion subscripts.
- 4. Assembles the coefficient problem matrix.
- 5. Computes and interpolates the first scatter source.
- 6. Assembles the source vector.

- 7. Checks for symmetry and diagonal dominance.
- 8. Solves for the Ai.j.; m expansion (mixing) coefficients by the method of successive over-relaxation.
- 9. Solves for the total particle fluence.

A ten point Newton-Cotes single integration routine was used to numerically integrate the thirty-seven integrals of Appendices F and G. This integration routine is an in-hou subroutine of the Air Force Aeronautical Systems Division, Wright-Patterson Air Force Base. The overall program logic

has been written in a manner whereby this integration routine could be used. The program is written in Fortran V.

#### Listing of Problem Subroutines

```
PROGRAM MAIN
DIMENSION CL (0:3,0:3) .SH : 0:2 .0:3 .DUSH : 0:3 .0:3) .PL (3,3 .8) .DUPL :
13.30.0A(3.3.300.0A(5) . (5) . (6)
DIMENSION PMZ(-::7), PMP(-2:7), PMX(-2:7), SPZ(7:7:5), SPP(7:7:6), SPX(
17,7,6),XSEVL(0:4),SVZ(7,5,3),SVX(7,5,3),SVR(7,5,3),APD(5,5,3),SB(5
10),A3(50,50),X3DDL(0:4),X3(0:4)
COMMON/A1/PNM,I/A2/ASH//ITYPE,ITS,ITTY
COMMON/BE/ISH/BS/IP
 ◆ THIS ROUTINE SOLVES THE EVEN PARITY ANISOTROPIC BOLTZMANN
 ◆ EQUATION AND THE AIR-OVER GROUND PROBLEM FOR THE REACTION
 ◆ RATE FLUENCE BY CALLING A NUMBER OF SUBPOUTINES AND USING A

    ◆ TEN POINT NEWTON COTES INTEGRATION ROUTINE.

READ(+. +. END=120) DZ.Z2.R2.DR.Z1.R1.ASH.LMAX
READ(*.*.END=120)YD,HB.SIGT,W
PRINT+ * * PROBLEM INPUT DATA:
PRINT+ *
PRINT+. (DZ= <.DZ) (
                       DF = \bullet DF \bullet
                                       尼含=1、尼含、1
                                                     Z2=1•Z2•1
                                                                    H∃H=
                        - R1=4,R1,4 RELAMATION FACTOR =4,W
1.8IH.
           21=1,21,4
PPINT++
PRINT** / LEGENDRE EXPANSION P(LMAX) = / * LMAX
READ: *. *, END=120) (XSEVL(I), I=0.LMAX)
READ: *. *. END=120) (XSDDL(I) · I = 0 · LMAX)
READ(*,*,END=120)(X3(I),I=0,LMAX)
PPINT+, TYD=1, YD, T
                    HE= •HE•
                                    SIGT=/,SIGT
PPINI++
PRINT+. LEGENDRE ODD EMPANSION SCATTERING CROSS SECTION. ALL ODD
IMDMENTS ARE ZERO.
PPINT***
              7.以(EVML) [1..]=A.LMAN()
 PERMIT * * 1
PF[MT+.
           LEGENDRE COD PARITY EMPANSION CONTTERING CROSS-SECTION. 4
PF IHT◆•
ILL (L) EVEN MOMENTO ARE ZERO.
FFINT++
               7.0000L+I).I=0.LMAN0.
PRINT++ (*)
FFINI 99
         LEGENDRE EXPANSION CROSS SECTION (ODD AND EVEN).
 · + THI RR
 PETHT.
 EE TYPE•••
                4.170mm Line [ = 0 + 1259 17 )。
 a⊊INT••
 FRINT.
 도구= (도전터) +1 (★)(토번편(1+2)) 공
 [TT]
 TOLE. COLL
 F BUMPT = 2 U O
 7点にし、可口医療を対しが30 ・バセ
```

```
PRINT * * * THE SPHERICAL HARMONIC COEFFICIENTS (L. M) . *
FRIHT++
DO 5 L=0.LMAX
PRINT+, (
               '∍€E(L∍M)∍M=0•L)
PRINT+, 4 4
PRINT+, 4
\times 1 = 0.
X2=8◆ATAN(1.)
TDL=.00001
KOUNT=200
ITYPE=1
           ANGULAR INTEGRALS OF SPHERICAL HARMONICS FOR THE 4
PRINT+,
PRINT++
            GALERKIN SOLUTION TO THE EVEN-PARITY FORM OF THE 1
PRINT+, 4
           NEUTRON TRANSPORT EQUATION.
PRINT++
           THESE ARE INTEGRALS OVER SOPHI OF THE REAL AND/OF EVEN
PRINT+, 4
           PART OF EXP(-IMX): EXPANDED IN SINES AND COSINES.
FFINT++
DO 15 ISH=1-18
CALL SPHER (LMAX, SH (0, 0, ISH), X1, X2, TBL, KDUNT)
PRINT+./ ANGULAR (HAPMONIC) INTEGRAL=/,ISH
DO 10 M=0.LMAX
PRINT+, C
              /*SH (M*M*ISH) *M=0*LMAX)
PRINT+ * *
FFINT++
CONTINUE
\times 1 = -1.0
X2=1.0
ITYPE=2
        THE ASSOCIATED LENGENDRE POLYNOMIALS INTEGRALS FOR THE
FRINT+ ·
PRINT+. GALERKIN SOLUTION.
PRINT*** THESE ARE THE POLYNOMIALS PL(L*M) (ACCO. LEGENDRE)
FFINT++
          INTEGRATED OVER THE INTERVAL (-1,+1).
FFIHT++
DO 25 IP=1.8
CALL POLY (EMAX.PL (1.1.1P) (X1.X2.TOL, KOUNT)
PRINT+./ ASSOCIATED LEGENDRE POLYNOMIAL INTEGRAL=/.IP
DB 20 IN=1⋅KT
空民共和工 * * *
              *FL - IN(* TT * TF ) * IT = 1 * FT)
丹里[14] **
中国其特工事。
CONTINUE
CALL ANGLE: LMAX: CL: Th: PL: CA: SIGT: XSEVL: XSBDL: KT: XSD
ITTi = 1
MM= (Z2-Z1) /DZ+1.1
DD 30 I=0+HX+2
EM.,(1)=54+1+±65
FMI INSERTION
9000 = 141 + 5
ITTFE=1
          CHATARD VALUES SER CHILHRE LATERPALS
 5 [ to F + +
TO THE .
PB 50 I=1.5
TRUE TRUET TRUET TOUR CUNTY
 SHIPPOS
```

```
ITS=1
 PRINT ◆• < MATRIX VALUES FOR 2-18ACE COMPCE INTEGRALS<
 PPINT++
 DO 58 I=1.3
 CALL SOURCE (SVZ.NZS.TOL.KOUNT)
CONTINUE
MM= (P2-P1) /DF+1.1
 D\Box = \theta \circ I = \theta \circ t t \otimes + \Xi
 PNM(I)=P1+I◆DP
 IF(PMM(I).E0.0.0)PMM(I)=1.0E+3
 PMP(I) = PMX(I)
MRS=MX+2
NT=kT+kT+ ( ) NZS+1> +NZS+ (NPS+1) ).
PPINT++
PRINT++
PPINT++
            THE TOTAL NUMBER OF TRIAL FUNCTIONS = •NT
PRINT++ / THE NUMBER OF LEGENDRE FUNCTIONS = + + KT
PPINT++
PPINT++
PPINT++
          PROBLEM GEOMETRY INPUT DATA.
            MODE (Z.P)
FFINT++
                              -MESH COORDINATE(Z,R):
FFINT++
DD 65 MR=0.NZ0-3
DB 65 MC=0:NPC-3
FRINT ◆ + 4 + 4 MP + + + ME + + 4 + 4
                                       - PHE (ME) - PME (MC)
FFINT .
FFINT++
ITYFE=3
ITI = 0
PRINT++
            MATRIM VALUES FOR RESPACE INTEGRALS.
PFINT++
DD 30 I≈1+6
CALL TELINE CEPRINES, TOLINETY
CONTINUE
ITI=8
PRINT ◆ · · · MATRIX VALUES FOR R-SPACE SOURCE INTEGRALS. ·
FFINI+.
TO 100 T=1.7
CALL IDLAGE SHINE PARE FROM HONOTO
CONTINUE
TO COMPUTE THE INTERPOLATION POINT VALUES
CALL DIRECT LIMANAGEDARMS, AMPS, MESASIGIA HEAVIDA ASHA
TO COMPUTE THE PROPUEM COURCE VECTOR
CALL BVEC: APD: CVC: CVP: CE: HCS: NPS: LMAD: CA: CIGT: FT: HT)
TO COMPUTE THE CTIFFNESS ATRIM.
CALL CTIFF: CA: SPP: CPC: NZC: HPC: CIGT: AS: PNF: PNZ: LMAC: NT:
 TO THE THE STIFFESS MATES
 THE THEFT HATHAIT
TO JOUNE THE PROBLEM MATRIX.
 PRODUCE WASEART FOR THE FOREST AT THE POSES.
o ⊊ THE••
FRINT** THE TOTAL LEAGUAL MEATTERED HOD DIRECT FLUM.
CETHT+.
           TRANS COINT TITLE
                                        TOTAL TOATTREED ELLO
IL DIFFIT FUN.
                        ាស្ត្រ ១១១ ១០១៩ ១៩៩៩៣៩+១៩៨៣៩៩៩៦
```

```
MT= (MTC-3) ◆2
MP= (MPC-3) ◆2
DZ=+Z2-Z1) / MZ
IF=(P2-P1)/MR
 DO 110 I=0.MZ
 DO 110 J=0.MR
ਲ=ਲ1+ਹ≠∄ਲੇ
 IF \cdot P \cdot LT \cdot PNR \cdot (0) \cdot P = PNR \cdot (0)
 C=Z1+I+DZ
CALL TELUM(ANR.ENC.XB.R.Z.KT.NZS.NRS.TAHI.DAHI.HB.SIGT.AIH.YI)
 TEHI=DPHI+TPHI
PRINT+ .
PRINT++
                                                                                                                     · TPHI,
                                                                                                                                                                                 · DPHI ·
                      TFHI
CONTINUE
GO TO 130
PRINT ** "PREMATURE PROBLEM DATA END"
 STOP
END
 SUBPOUTINE SPHER (LMAX, DUSH, X1, X2, TOL, KOUNT)
DIMENSION DUSH (0:3.0:3)
COMMON/B1/M.H/B2/ISH//ITYPE,ITS.ITTY
 ◆ THIS POUTINE NUMERICALLY INTEGRATES THE SINES AND COSINES SINGLE

    INTEGRALS OF THE ANISOTROPIC EVEN PARITY BOLTZMANN EQUATION BY

    USE OF A TEM POINT MEMTON-COTES INTEGRATION POUTINE.

 IF (IIH.LE.9)68 TO 80
IF (ISH.E0.10) THEN
 1=0.
ME=E.◆ATAM(1.)
50 TO 20
ELIE IF (ISH.E0.11) THEN
12=2. *ATAN:1.)
111=5◆戶下戶付(1.)
GD TD 20
ELSE IF (ISH.E0.12) THEN
M1=5◆ATAM(1.)
 West of the Partie !.
FIDME
 CONTINUE
DO BO MED.LMAN
DO 30 N=0.LMAX
CALL OBAD (21.02.TOL.DUSH (M.N).KOUNT)
 IF DUTH M.H.LT.1.0E-10 DUTH M.H. = 0.
CONTINUE
FETTEN
21:1
 CHERROLITINE FOLVILING ADDRESS, 1.02. FOLL BOOKT
DIMENTION DUTY 1.3
   CMMRRE EL MAREEL LA SALLAR LA TRABEATTATTA
• FOLIZIONE COMENTO: THE CHEGRAL OF THE ATTOCIPTED LEGENCE
• FOLIZIONE COMENTO: THE CHEGRAL OF THE CHARACTER EVEN PRESTY
• FOLIZIONE COMENTO: THE CHEGRAL OF THE ATTOCIPTED LEGENCE
• FOLIZIONE COMENTO: THE CHEGRAL OF THE ATTOCIPTED LEGENCE
• FOLIZIONE COMENTO: THE CHEGRAL OF THE ATTOCIPTED LEGENCE
• FOLIZIONE COMENTO: THE CHEGRAL OF THE ATTOCIPTED LEGENCE
• FOLIZIONE COMENTO: THE CHEGRAL OF THE ATTOCIPTED LEGENCE
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IF (IP.LE.6)60 TO 30
 IF (IP.E0.7) THEN
 1/1 = 0.
 ⊠2=1.0
 50 TO 30
 ELSE IF (IP.E0.8) THEN
 5(1=0.0)
 M2 = -1.0
 EMDIF
 CONTINUE
 I d = 0
 DO 40 F=0.LMAX
 DC 40 N=0.8
 IT = 0
 Ibi = Ibi + 1
 DO 40 L=0.LMAX
 IDD 40 M=0.L
 IT=IT+1
 CALL OUAD: X1. 2. TDL. DUPL: IN. IT) . KOUNT)
 IF : DUPL : IM : IT : . LT. 1.0E-10 DUPL : IM : IT ) = 0.
 CONTINUE
 FETURN
 END
 FUNCTION YOU
 DIMENTION FAM: -3:7:
 COMMON B1 M.H/B2/ICH/B3/IP B4/L.k//ITYPE,ITC.ITTY
COMMON A1 PMM, I HA2 ACH AB JP. JT /A4 / FS. IH. LP

    THIS FUNCTION IS USED BY THE TEN POINT NEWTON COTES INTEGRATION◆

    ◆ ROUTINE (AUGI) TO SELECT THE INTEGRATION PUNCTIONS.IT USES THE ◆
    ◆ INFORMATION IN THE ABOVE COMMON PLOCKS TO DETERMINE THE FUNCTION

    WHICH IS PEING INTEGRATED.

 IF (ITTW.EC.1)GB TD 85
 IF: ITYPE.EG.2:60 TO 20
 IF (IIH.ED.1) THEM
 写宝色图页(11)◆◆三◆色图页(10◆11)◆色图页(10◆2)。
 SETTING.
RETURN
ELIE IF FIGH.E0.3.78.17H.E0.10.08.13H.E0.11.08.18H.E0.120 THEM
 V=501 (1) •001 (N•M) •001 (N•1)
RETURN
 ELIE IF (IIH.ED.4) THEN
                          100
分配であれ
可以には、1月 (100年) 表でいた。 100年年
日本の日本・100年 (100年) 100年
                                                                         The state of the s
្លាប់ប្រាស់
សាធ្យាស្រុសស្រុសស្រុសស្រុស
ឯស្រាស់
                         1,000 - 100
  LIL IS ATTHE ME. TO THERE
```

```
Y=008(X) + 008(N+X) + 8IN(M+X)
RETURN
ELGE IF (ICH.E0.8) THEN
Y=005 (N+X) +5 IN (M+X)
RETURN
ELSE IF (ISH.EQ.9) THEN
Y = (C \sqcup S (X) + C \sqcup S (N + X) + N + S \exists N (X) + S \exists N (N + X)) + S \exists N (M + X)
ENDIF
RETURN
CONTINUE
Ĥ≈1-⊠◆◆2
IF (A.LE.1.0E-10) A=0.
IF (IP.EQ.1) THEN
Y≈A+PLF (X+K+N) +PLF (X+L+M)
RETURN
ELSE IF (IP.E0.2) THEN
Y=SORT (A) ★X◆PLF (X,K,N) ◆PLF (X,L,M)
RETURN
ELGE IF (IP.E0.3) THEM
Y=SORT(A) ◆PLF(X+K+N) ◆PLF(X+L+M)
PETURN
ELSE IF (IP.ED.4) THEN
Y=X++2+PLF(X+k+N)+PLF(X+L+M)
RETURN
ELDE IF (IP.E0.5.0F.IP.E0.7.0F.IP.E0.8) THEM
Y=M→PLF(M+k+h) →PLF(M+L+M)
PETURN
ELSE IF (IP.E0.6) THEN
Y=PLF (X+F+M) ◆PLF (X+L+M)
ENDIF
RETURN
CONTINUE
IF (ITYPE.E0.2) GD TD 30
IF: ITT.E0.1060 TD 180
IF (I.EC.1) THEM
f = M
11=1
GE TO THE
FLIE IF HILDOLG FHEN
11=3
60 TO 70
ELIE IF (I.EO.3) THEN
11=2
11=3
GD 70 70
ELIE IF FI.CO.46 THEN
##1
-50 TT 75
TLUE IS -1.50.50 THEN
T = 1
GO TO TO
THRIF
CHILIME
```

```
IF (ITS.EQ.2)60 TO 130
IF (I.E0.1) THEM
M=3
14=3
60 TO 70
ELSE IF (I.EO.E) THEN
11=1
60 TO 70
ELSE IF (I.EO.3) THEN
M=1
11=1
60 TO 70
ELSE IF (I.EQ.4) THEN
M=3
1=1
60 TO 70
ELSE IF (I.EQ.5) THEN
M=1
14=1
60 TO 70
ELDE IF (I.E0.6) THEM
M = 1
M=1
ENDIF
SP1=SPF (M*k*PNM(UP+3)*PNM(UR+2)*PNM(UP+1)*PNM(UP)*M)
IF:ITYPE.E0.2:60 TO 100
IF (I.LE.3) THEM
Y=EMP(M×BSH) ◆SP1◆SP2
FETURN
ELIE IF (I.EO.4) THEN
Y=EMP+-Y>60H++661+600
PETURN
ELDE IF (I.EO.5) THEN
Y=IP1◆FP3
FETHER
EHDIF
CONTINUE
IF (I.LE.3) THEM
IF:01.E0.0.:00=1.0E-30
V=IF1+IF2 I
EFTURN
ELGE IF (I.E0.4.8F.I.E0.5) THEM
E 5 € 63
면도 및 IP (I.B).++ IHEN
· 电对应性电流操作率
i = Tilip Ni
----
1974 F 1990 Z
TE (I.EC.1) THEN
* • = :
45 75 74 8
EL 9 19 10 1 7 .
                -----
```

```
M=2
60 TO 140
ELGE IF (I.E0.3) THEN
14 = 1
50 TO 140
ENDIF
CONTINUE
IF (I.EQ.1) THEM
M=3
60 TO 140
ELSE IF (I.EO.2.OR.I.EQ.3) THEM
14=1
EMDIF
SPH1=3PF(N,KS,PNX(LR-3),PNX(LP-2),PNX(LR-1),PNX(LR),X)
SPH2=HIF(IH.PMX(L).PMX(L+1).X)
IF (ITS.E0.2) 60 TO 160
IF (I.E0.1.OP.I.E0.2) THEN
Y=SPH1+SPH2+EXP(X/ASH)
RETURN
ELSE IF (I.EO.3) THEN
Y=SPH1+SPH2
PETURN
EMDIF
CONTINUE
IF (I.EO.1.DR.I.EO.2) THEN
Y=IPH1+SPH2
RETURN
ELSE IF (I.EO.3) THEN
Y=3PH1+3PH2+X
EMDIF
RETURN
EMD
FUNCTION PLF (X, I, U)
◆ THIS POUTINE SELECTS AND COMPUTES THE ASSOCIATED LEGENDRE
◆ FUNCTION WHICH IS BEING INTEGRATED.
9=1.- ••2
TF-A.LE.1.0E-100A=0.
IF (I.EO.O.AND.J.EO.O) THEN
FLF=1.0
PETURN
ELGE IF (I.EO.1.AND.J.EO.0) THEN
FLF=
PETHEN
ELIE (F. T.EK.1.AMD.1.EG.1) THEM
71.5 = 1057.50
113:17 3 A
SLIE 15 -1.55.2.346...546.00
ig this to
ELIE IF (I.En.E.Ahb.J.E0.1) THEN
ಶ್ವಾಕ್ಕಳ್ತಿ • • ಂದ್ರಾಗ∗ಕ
दहेरा दश
```

ELIE IF -1.80.2.800.3.20.2 1469

```
PLF=@+A
RETURN
ELSE IF (I.EQ. 3. AND. J. EQ. 0) THEN
PLF=(48.*(X**3)+72.*X*B)/48.
RETURN
ELSE IF (I.EO.3.AND.J.EO.1) THEN
PLF=~(SOPT(A) * (288*X**2+72*B))/48
RETURN
ELGE IF (I.EO.G.AND.J.EO.E) THEN
PLF=15.◆>+0
FETUEN
ELSE IF (I.EO.S.AND.J.EO.S) THEN
PLF=-15. + (A++1.5)
ENDIF
RETURN
EMD
 CUPPOUTINE ANGLE(LMAX.CL.SH.PL.CA.SIGT.XCEVL.XSODL.KT.XS)
DIMENCIEN CL/0:3:0:3).3H/0:3:0:3:12/.PL/3:3:80.CA/3:3:30).MCEVL 0:
14) • MEDDL/0:4) • MC (0:4)
 ◆ THIS POUTINE ASSEMBLES THE TOTAL ANGLE INTEGRALS FOR ALL
 ◆ COMBINATIONS OF THE SPHERICAL HARMONIC TRIAL AND WEIGHT FUNCTION

    ■ SUBSCRIPTS.

           TOTAL INTEGRAL VALUES IN MATRIX FORM, OF THE SPHERICAL
PRINT+.
           HARMONIC TRIAL AND WEIGHT FUNCTIONS USED IN THE
PRINT++
          GALERKIN SOLUTION OF THE EVEN-PARITY BOLTOMANN EQUATION
PRINT.
           THESE VALUES ARE SELECTED PRODUCTS OF THE ACCOMMANDED
PRINT++
FFIHI++
           LEGENDRE POLYNOMIALS AND ANALYAR INTEGRALS WHICH WERE
           COMPUTED EARLIER. THEY REPRESENT THE SCATTERING AND
PP[HT++
          NON-TOATTERING ANGULAR INTEGRAL FERNELT WHICH ARE FORMET
PRINT+.
           WHEN TRHERICAL HARMONICT ARE USED AT THE AMBULAR TRIAL AND WEIGHT FUNCTIONS IN THE GALERKIN SOLUTION.
PRINT ..
PFINT .
PFINT+.
PRINT .
DB 220 I=1:29
 IF ~ 1.50.15 THEN
 -=1
 184=1
60 FD 10
ELIE IF (I.ED.2) THEN
 IIA=2
 IF'A≃1
50 10 10
ELIE IF VI.55. 46 THEM
 : 4=
 ្ជដ្
 GC TO 10
 항기도 3톤 HILT .44 7HEH
 -
-
-
-
-
60 TO 11
ELIE (F. 1,50.5) THEM
 15-45-1
```

```
6D TO 10
ELSE IF (I.EG.6) THEM
ISA=5
IPA=2
GO TO 10
ELSE IF (I.EO.7) THEM
ISA=3
IFA=2
50 TO 10
ELSE IF (I.EO.8) THEN
ISA=5
IPA=2
60 TO 32
ELSE IF (I.EO.9) THEM
ISA=6
IPA=4
60 TO 10
ELSE IF (I.EO.19) THEN
ISA≈6
IPA≈6
GO TO 10
ENDIF
GO TO 36
DB 30 K=0.LMAX
DO 30 M=0.1
IT = 0
I \sqcup = I \sqcup + 1
DD 30 L=0.LMAX
DB 30 M=0.L
IT = IT + 1
CA:IM.IT.I) = CL:(k.M) + CL:(L.M) + CH:(M.N.ICA) + PL:IM.IT.IFA:
CONTINUE
60 TO 38
DO 35 F=0.LMA
DO 35 H=0+1
TT = ii
I10 = I10 + 1
한번 기트 ( =10.0MA)
한번 기트 ( =10.0MA)
[T=[T+1
设备(IM・IT・I:=2.2.1 +10.+12.+12.+13.+14.+14.M・IUA(★F2.1M・IT・IFA)
BURITHDD
IF (I.LE.9.DR.I.ED.19)50 TO 200
IF (1.50.21.86.1.50.29) THEN
DD 40 -=0.LMH
DO 46 H=6+2
: T=:
11:11 ( + )
DB 410 是电机+运用序
*3 48 Man
TEMP1=[H:M:0:10+10+10+12:-- H:0::0:12:-
THOUGHT TO BE LOCATED BY SELECTION OF THE TABLE THE TABLE
Big fo 40
FUTE OF 00.20.230 FH≥10
```

```
TEMP1=PL(IW,IT,7)+PL(IW,IT,8)
 CA(IW,IT,I) = CL(K,N) + CL(L,M) + TEMP1 + SH(M,N,6)
ENDIF
 CONTINUE
GD TD 200
ELSE IF (I.EQ.22) THEN
 ISA=3
 IPA=3
 60 TO 45
 ELSE IF (I.E0.23) THEN
  ISA=5
  IPA=3
 60 TO 45
 ELSE IF (I.EO.24) THEN
 ISA=6
 IP8=5
 60 TO 45
 EMDIF
 GD TO 55
 DO 50 K=0.LMAX
 DO 50 N=0.k
  IT = 0
  I I d = I I d + 1
DO 50 L=0.LMAX
DO 50 M=0.L
 IT=IT+1
MI=L-M
H=(1.-(-1.) ◆◆MC)
 IF (M.EQ.0)A=A/2.
 \mathsf{CA} \circ \mathsf{IM} \bullet \mathsf{IT} \bullet \mathsf{ID} = \mathsf{CL} \circ \mathsf{L} \bullet \mathsf{MD} \bullet \bullet \mathsf{2} \bullet \mathsf{CL} \circ \mathsf{K} \bullet \mathsf{ND} \bullet \mathsf{SH} \circ \mathsf{M} \bullet \mathsf{H} \bullet \mathsf{ISA}) \bullet \mathsf{FL} \circ \mathsf{IM} \bullet \mathsf{IT} \bullet \mathsf{IFA}) \bullet \mathsf{A} \bullet \mathsf{MS} \circ \mathsf{L} \circ \mathsf{MS}  IF (I.GE.22.AND.I.LE.24) GD TD 200
CONTINUE
 IF (I.EO.28) THEN
 DO 60 K=0.LMAX
DO 60 N=0.K
IT = 0
 Thl = Thl + 1
  DC +0 €=0.EMAS
   38 50 M=0.L
 IT = IT + 1
MI=L-M
(百=(1.+(−1.)◆◆問答)
IF(M.EQ.O)A=A/2.
50 TO 200
     THEF
 UMBIA
18 - 1.80.10- 7484
11441
11847
11847
    : IE = .
  [[4]=[
  I \mapsto E = \mathbb{P}
98 75 120
ELIS 18 (1.20.11) 1958
   ::n=
```

```
ISB=5
ISD=7
ISE=9
IPA=3
IPE=3
60 TO 130
ELSE IF (I.ED.12) THEN
ISA=3
I \supseteq E = F
I [ I = 7
 I0E=8
 IFA=5
 IPB=3
 60 TO 130
 ELSE IF (I.EQ.13) THEM
 ISA=5
 I2B=3
 I2I)=9
 ISE=7
  IPA=3
  IEE=3
  60 TO 130
  ELSE IF (I.E0.14) THEM
  I 16=5
  I 5 F = 5
  I \in D = \emptyset
  15E=∄
   IPB=3
   IFF=3
  60 TO 130
  ELIE IF (I.ED.15) THEN
   ITA=5
   I : E==
   I \subseteq I) = \mathbb{P}
   ICE=8
   IPH=5
   [FF=5
   HOTE IN ELIEN THEM
    7.7 H = 7
    I \cap F = \mathbb{F}
    I\subseteq \underline{I}(=\mathbb{S}
    ITE=T
    IPH=?
    IFE=5
    HE TO 10 SERVICE THEM
     1-1-
     [F: F. = =
     70 TO 1
                1.55.14 · THEN
      1 - A = H
```

```
I TE=6
 I \odot D = \odot
 ISE=8
 IPA=5
 IPB=5
 GD TD 130
 ELSE IF (I.E0.20) THEN
 I THER
 ICE=6
 I 5 D=8
 ISE=8
 IPH=6
 IPB=6
 60 TO 130
 ENDIF
 60 TO 150
 DO 145 K=0.LMAX
 DB 145 N=0⋅k
 ITņ
 I M = IM + 1
 DO 145 L=0.LMAN
 IO 145 M=0.L
 I = 0
 IT = IT + 1
 CA(IM+IT+I)=0.
 DO 145 JR=0.LMAX
 TEMF=0.
 MENUEMBOIL (UP) / (IGT-MIDDL (UP))
 IF · I.EO.200 KNU=KEEVL · UR)
 D□ 140 JI=0.UP
IF=IF+1
IA=2
IF:US.ED.ONIA≕IA-2
 TEMP≈IA→6L(M,H)→6L(L,M)→6L(L,M)→6L(UR,JE)→+2→(EH(JE,N,IEA)→8H(JE,M,IEB)+EH
1 (J:,N,I:D) ◆3H(J:,N,I:E)) →FL(IT,IF,IFA,+FL,IM,IF,IFE) +TEMP
CONTINUE
 CA: IN: IT: I) = TEME + XX U+CA: IW: IT: I)
Confide
Confide
 iFkt.35.10.4⊬t.1.15.20055 ⊤5 20
 IF (I.ED.25) THEN
IIA=3
IIE=6
III=7
ITE=
 * = 4 = 4
 15 E = 3
30 TO 15 0
 ELIE TE I.ES.BAR THEN
 10 Feb
10 Teh
10 Teh
 I [E=-
 15----
 ins:
```

Country of the Countr

```
ELCE IF (I.EO.ST) THEM
 I : H = H
 I \cup F = 0
 I \ni I) = \oplus
 I 9E=8
 IPH=5
 IPE=6
 FIDHE
 DB 175 K=0.LMAX
 DO 175 N=0.8
 I T = 0
 IM = IM + 1
 DB 175 L=0.LMAX
 DO 175 M=0.L
 IF = 0
 IT = IT + 1
 CA: [M: IT: I) = 0.
 DD 175 JR=0.LMAX
 TEMP=0.
 M:U=MSBDL:UP://(SIGT-MSDDL(UR))
 DB 170 JD=0.UF
 IF = IF + 1
 M = L - M
 府=2★(1.一(-1.)◆◆門图)
 IF:UI.E0.0.DR.M.E0.0)A≃A/2
 TEMR=(L.(L.M)+(L.K.M)+(L.(JR,J):(++2+(SH)JS,M,ISA)+SH,JS,M,ISB)+SH.
1 • N • I : D • • : H • : D • M • I : E • • ◆ P L • I T • I P • I P A > P L • I M • I P • I P F ) ◆ A + T E M P
CONTINUE
CA([M*IT*I)=CA(IM*IT*I)+TEMP*MkU*XS*L)
 CONTINUE
 PRINT+. MATRIX VALUE: FOR ANGLE INTEGRAL= 1.
 DO 210 IM=1+FT
 FFINT+.
             - < • C 白 · I ld • I T • I ) • I T = 1 • k T )
 PRINT+.
 FFINT .
 BUNITHOD
 RETURN
 FMI
 CHERENTIME COOPERCHARY
 DIMENSION CLAUSSANSON
 ◆ THIS EQUITINE COMPUTES THE SPHERICAL HARMONIC EXPANSION
 • COEFFICIENTS OF MORMALIZATION COMSTANTS.
 FEAL INF . IDF
 าติ หน้อก.พ.ศหก.ย.ธก.พ. 55 (55 (ศก
กษาย (54
 [F ⋅ INT.50.00 GD TO 50
   o kito oj molektiko ken
  · · □ · , · · □ • ·
```

```
IDF=IDF+I
 FT=INF I DF
 50 TO 100
 FT=1.
 CL(L+M)=30PT((2+L+1))(4+(4+ATAH)1, (()+F))
 CONTINUE
 RETURN
 END
  SUBPOUTINE MALINE (SPM + NMS + TOL + ) DUNT (
 DIMENSION PMM(+2:7).CPM(7.7.6)
 COMMON/A1/PMM.I/A3/JF.JT B4/L.F
  ◆ THIS ROUTINE COMPUTES THE BICUBIC SPLINE INTEGRALS BY CALLING ◆◆
  ◆ A TEN POINT NEWTON-COTES INTEGRATION POUTINE.
 28M+1=9U 06 DU
 DB 90 UC=1.NMS
 SPX(JR*J(I)=0.0
  JT=JC-JR
 IF (APS (UT) .GT. 3) 60 TO 90
⊁F=4
 IF (UT.EQ.O) THEN
 F'I = 1
50 TO 5
ELSE IF (UT.E0.1) THEM
60 TO 5
ELIE IF (UT.ED.2) THEN
+I=3
60 TO 5
ELGE IF (UT.E0.3) THEN
1 I = 4
60 TO 5
ENDIF
F I = 1
IF CUT.EO.-10 THEN
F=3
50 10 5
ELIE IF LT.ER. -2: THEN
* F = ?
50 TO 5
ELIE IF (UT.EO.-3) THEN
FF=1
60 TO 5
EHILLE
DE 10 '= I ++ F
 _ = _(2 = = ++
1 T = 4F - F4 (14)
 TERLITIONED TO 10
 17 (1.94) 1 98 78 19
 THE COURT FOR THE SECURITY OF THE STATE OF THE SECURITY OF THE
BIRTINDS
   TINE TO BE
1. 1 to / 4.
                                  THERE IS NOT THE ENDED DESCRIPTING AT
25 1 5 € €
```

```
DO 100 MR=1.8MC
PRINT ◆ • • * 1 • SPM (MR • MC • I) • MC = 1 • MMC •
FFINT++
FFIHT .
RETURN
END
FUNCTION SPECIP. UP. X1. X2. X3. X4. XX

    THIS POUTINE SELECTS AND COMPUTES THE BIGUBIC SPLINE FUNCTIONS ◆

* THAT ARE BEING INTEGRATED.
H= (114-16)
[]=6◆◆3
21=-3+A++2.
IF (JP.ED.4) THEN
IF (IF.ED.1) THEN
IFF=I
RETURN
ELDE IF (IP.ED.E) THEN
IFF=I1
FETUEN
ELSE IF (IP.E0.3) THEN
IPF=Z+X+Z1
PETURN
ENDIF
EHILLE
\mathbf{F}' = \mathbb{N} \cdot \mathbb{R} - \mathbb{N}
7=2-4+E++3
V1=18.+B++8+Z1
IF JP.EO.30 THEN
IF FIR.ED. 19 THEN
IFF=Y
PETURN
ELIE IF (IF.E0.2) THEN
[FF=11
PETURN
ELIS IF (IP.E0.3) THEN
FRE=1 + 1+11
ser sn
នារំព្រះ
ENI IF
₩≡₽+6+6+₹
V1=V1-18.★6★★8
IF FUR. RO. 2 F THEM
IF + IP.EG. 1: THEM
 c :. -;
Supremental Services
TL 15 (18,50.8) THEN
ADDITION OF A
নি কৈ কি কে. ইক. ১৮ জাল্য
1 T = 1 + 1 + 1 + 1
FFTHFH
```

```
D=×1-×
W=V-4+D++3
M1=V1+12.*P**2
IF (UP.EO.1) THEN
IF (IP.E0.1) THEN
SFF=W
PETURN
ELSE IF KIRLEGUEY THEN
îFF=ld1
PETURN
ELSE IF (IP.E0.3) THEN
SPF=6+1/4611
ENDIF
FIDITE
PETURN
END
SUBPOUTINE SOURCE (SWM. NUS. TOL. FOUNT)
DIMENSION SVX:7.5.3:.PMX:-2:7:
COMMONABLARMMAI A4AKSAIHALRAB4ALAK
◆ THIS POUTINE SELECTS AND COMPUTES THE INTERPOLATED SOURCE
◆ INTEGRAL BY CALLING A TEM POINT NEWTON COTE: INTEGRATION POUTING
DO BO LEST. NMI
DB 80 LC≈1⋅N 1-8
DVM(LE*LC*I)=0.0
LY=LC-LF
IF LY. GT. 1. DP. LY. LT. - 3: GD TO 80
IF TLY.ED. 1: THEN
| I = 4
GD TO 10
ELIE IF LY.EO. OF THEM
GO TO 10
EMPIF
IF (LY.EO.-1) THEN
- [ = ]
F = 1
4D TO 18
ELIE IF (LY.E0.-2) THEN
I = 1
1F=2
60 TO 10
ELIE IF HLY.En. -34 THEN
1 7 = 1
1 15 21
表 1115年
T 14 = 1
[F:415.50.=1:1H==
15 15 - 1=11+1F
_ =1 [F = = +
LF=LF-14 [++]
IF WALTACHOOMS IN
17 H F. ST. 13 98 78
FALL MANAGEMENT OF SPIN OF STATE OF STATE
```

```
SVX(LP,LC,I) =SVX(LP,LC,I)+TV
 IH=IH+1
CONTINUE
PRINT+. / MATRIX SOURCE INTEGRAL=/.I
DO 90 MR=1.NXS
PRINT++ (1)
             - (,SVX(MR,MC,I),MC=1,NXS-2)
PPINT++ (
RETURN
THI
FUNCTION HIF (IH - X1 - X2 - X)

    THIS ROUTINE COMPUTES THE LINEAR LAGRANGE POLYNOMIALS WHICH ◆◆◆

 ◆ ARE BEING INTEGRATED AS A PART OF THE SOURCE INTEGRALS.
IF (IH.E0.2) THEN
HIF=(X2-X)/(X2-X1)
RETURN
ELSE IF (IH.EQ.1) THEN
HIF = (X - X1) \times (X2 - X1)
EMDIF
RETURN
END
 SUBPOUTINE DIRECT(LMAX.APD.FNZ.PNP.NPS.NZS.SIGT.HB.YD.ASH)
DIMENSION APD(5.5.3), PMZ(-2:7), PMP(-2:7)
 ◆ THIC ROUTINE COMPUTES THE DIRECT SOURCE POINT VALUES WHICH ARE ◆

    USED IN THE FIRST COLLISION SOURCE INTERPOLATION.

PPINT++ / /
PRINT * * INTERPOLATION POINT VALUES. ELEMENT (I.J.) IS FOR POINT OF
1 F)
PRINT++ 4 4
I d = 0
DO 50 L=0.LMAX
DO 50 M=0.L
Ibl \approx Ibl + 1
DD:40 IT=1:MEC-2
 ID 40 I=1+HFI-E
CHEFNC (ITHE) - HE
PZH=10FT(FMP)1-10 • • 2+2H• • 2.
IF (RZH.E0.0) 68 TO 60
UD≈ZH-RZH
IF (UD. EQ. 0) GO TO 20
TAU=CIGT ◆ (EMP) - HEVACH (HEMP) - PPZ (IT-1) (ACH) (→ACH) UD
TANGETIAT * FI'D : GDNC : IT = 10 GOG + GND : I = 10
INCITEIS HIH!
PENTITUE
          MATERN MALUES FOR INTERFOLATION FORM MALUES FOR NESS (F
ESTITA.
43 • FILMPENSER 2:
##INT*•
DD 45 MR=1.HZI-2
⊒≘[NT••+
               → 無無限の機能を対しますが、まずに対します。資格に当まる。
99 [HT++
全部性工程规据
```

```
60 TO 70
 PRINT • • < IMPROPER INPUT DATA. A MODE CANNOT BE AT THE BURST FOINT • Z
1= • PM2 (IT-1)
 STOR
 RETURN
 END
 SUBPOUTINE BYECKAPD:SVZ:SVP:SB:NZS:NRS:LMAX:CA:SIGT:KT:NT>
 DIMENCION APD(5.5.3).SVZ(7.5.3).SVP(7.5.3).SB(50).CA(3.3.30)

    THIS POUTINE ASSEMBLES THE PROBLEM SOURCE VECTOR.

 PRINT+ - C
 PRINT •• / 38URCE MATRIX COLUMN VECTOR./
 PFINT+
 I \le = 0
 DB 90 UZ=1.NZS
 DO 90 JR=1,NRS
 JH=0
 DO 80 JL≃0.LMAX
 D□ 80 JM=0.JL
 JA=JA+1
 I \mathbb{S} = I \mathbb{S} + 1
 SB \cap ISD = 0.
 DD 75 M=1.7
 30 = 1.0
 TEMPREO.
 IĤ≃Ĥ
 DD 70 IL=0.LMAN
 DO 70 IM=0.IL
 IA=IA+1
 TEMP2=0.
BD 60 IZ=1.MZ3-2
 DO 60 IP=1.MPS-2
 192=12-92
 IDF=IF+JF
 IF (ICC.GT.1.DP.ICZ.LT.-3)6D TO 60
 IF: ICP.6T.1.DP.JCF.LT.-3060 TO 60
 IF 9.20.19 5 89
 · _ = ;
145 = 1
林帝= 30
50 TO 50
 ELIE IF (M.ED.2) THEM
147 = 1
14F = 3
 ·== [
  ELIF IF (N.EG.4) THEN
1000
 ٠, ٠,٠٠٠
```

```
60 TO 50
ELSE IF (N.EQ.5) THEN
MZ=1
MR=2
MA=26
GD TD 50
ELSE IF (M.EQ.6) THEM
MZ=2
MP=3
MA=27
GD TO 50
ELSE IF (M.EQ.7) THEM
MZ=3
MR=3
NA=28
ENDIF
TEMP1=SVZ(JZ:IZ:NZ)+SVR(JR:IR:NR)+APD(IZ:IR:IA)+CA:JA:JA:IA:NA)
TEMP2=TEMP1+TEMP2
CONTINUE
TEMPS=TEMPS+TEMPS
IF(NA.EQ.23.OF.NA.EQ.26)00≃-1.0
IF (N.EO.7) 60 TO 78
TEMPS=TEMPS+S+ATAN(1.)/SIGT+SC
60 TO 75
TEMPS=TEMPS+S+ATAN(1.)+30
SB(IS)=TEMP3+SB(IS)
CONTINUE
            /*RT** ELEMENT(S) TO ELEMENT NUMBER=/*IS
PRINT+, C
           - - - /**SB**KSD**KS=IS-*KT+1D**ISD
CONTINUE
PRINT+, 1
FRINT++
PPINT++
PETURN
EHD
SUBFOUTINE STIFF(CA.SPR.SPZ.NZS.NRS.SIGT.AS.PNR.PNZ.LMAM.NT)
DIMENSION CA(8.9.90).SPP(7.7.6).SPC(7.7.5).AC(50.50).PNR(-2:7).P
1112 (48:7)
*************************************
THIS POUTINE ACCEMBLES THE PROPLEM COEFFICIENT OR STIFFMESS MATRIX.
DB 20 U=1.NT
DO 20 F=1-NT
H_{-}^{*}(J_{1}) = 0.0
1. 1=0
[6] 30 J[≈1•H7]
5D 39 39±1•095
20 50 JESP-6MAN
20 50 Man-96
  .a. ; + !
 A = 29 + 1
1 1=0
TE 50 (1741-02)
THE GO TRAINET
```

```
DO 50 IL=0.LMAX
DO 50 IM=0.IL
IA=IA+1
I \cup = I \cup +1
TEMP1=0.
TEMP2=0.
LD=UZHIZ
MF=JF-IF
IF ABS LD .GT. 30 GD TD 33
(F.AES/ME).61.3/60 TO 93
DO 32 N=1-18
ME=E
IF (N.ED. 1) THEM
MF = 1
HZ=1
GO TO 30
ELIE IF (N.EO.2) THEN
 14F=2
 MZ=1
 NE=1
 68 TO 30
 ELDE IF (N.ED. 3) THEN
 HF=4
 원=일배
 60 TO 30
 ELIE IF M.ED. 41 THEM
 MF=3
 MC=1
 #E=1
 60 TO 35
 ELSE IF M. ED. 50 THEM
 MF = 3
 ME=1
  60 TO 30
  ELSE IF MI.EC. BY THEM
  148=5
  ME=2
  iaĘ".
  FILE IF WH.ST. AN THEM
  MF = 4
  HI=E
  60 TO €1
  ELIE IF (N.ED.8) THEM
  :4F = 5
  p_{i,j} = 0
   15 = 1
   ---
   ELLE IF 14.EA. - THEM
   ,: · -
   ELIE AF 14.6%. 140 Theb
   45⊊)
47⊊1
```

```
ELSE IF (N.EQ. 11) THEM
MP=2
112=1
ME=1
60 TO 30
ELSE IF (N.E0.12) THEM
ME=4
MZ=2
GD TO 30
ELSE IF (N.ED.13) THEM
MF=2
MZ=1
ME=1
60 TO 25
ELSE IF (N.E0.14) THEN
MR=3
MZ=1
6D TD 30
ELSE IF (M.EQ.15) THEM
MR=5
NZ=2
t:E=1
GO TO 30
ELDE IF (N.EQ. 16) THEM
11F=4
 MZ≈2
 GO TO 21
 ELSE IF (M.EO.17) THEM
 MP=5
 ME=2
 ME=1
 68 TO 83
 ELSE IF (M.EQ.18) THEM
 대단=중
 MZ=3
 GD TO 30
 EHDIF
 TEM===FF ( FF . JF . NF) + 3FD ( IZ . JZ . NZ)
 30 to 12
 *EM=[FF:UP.IF:NP)*SFZ(IZ:UZ:NZ)
 GO TO 32
 TEM=SPR(IR.JP.NR) +SPZ(JZ.IZ.HZ)
 GO TO 32
 TEM=(PF) JF . IP . NP) . (FZ (JZ . IZ . NZ)
 TEMP1=TEM+CA: JA: IA: N: +: -1: ++HE+TEMP1
  CHILDINE
  वर्षे (दे.वरे.३.८२.१८.३१.३५ ५<u>० १७</u>.३९
१७/२३१४५६१.६४.१५५<mark>० १८</mark>.३९
  IF A.J.EA.IS THEM
  1_1=4
  IF LD.FO.GO THEN
  1,374
  7<u>0</u> 70
```

```
ELIE IF (LD.EG.-1) THEM
L2=3
60 TO 35
ELIE IF (LD.E0.-2) THEM
L2=£
60 TO 35
FIHIMS
ELDE IF (UD.E0.2) THEM
1 1 = 9
IF (LD.EO.O) THEM
LE=3
 60 TO 35
 ELIE IF (LD.E0.1) THEM
 L3=4
 60 TO 35
 ELCE IF (LD.EQ.-1) THEM
 <u>1</u>2=2
 60 TO 35
 ENDIF
  ELDE IF (UC.EG.3% THEN
  L1=2
   IF (LD.ED. D) THEN
   LE=E
   60 TO 35
  ELIE IF (LD.ED.) THEN
    LE=3
    60 TO 35
   ELSE IF (LD.EG.E) THEN
    L2=4
   EMPIF
   PEDI=JRF(1.L1.PMJ)JZ-8:.PMZ(JZ-8).PMZ(JZ-1).PMZ(JZ-1).PMZ(JZ).PMZ(J)...*SPF(1.
 1L2, PMZ (IZ-3), PMZ (IZ-3), PMZ (IZ-1), PMZ (IZ), PMZ (0))
    60 TO 40
     :E2170.
    FERENCE - JC
ME-MCI-JC
     HERMOLTIC
TERROLET. B. OP. MO. GT. SREED TH 48
TO RESEMBLIST, BREET TO 45
THE COURT BRIDE TO ASSESSED TO 45
THE COURT BRIDE TO BE TO
     F D=HF I - JF
       IF () D. GT. 2. DF. MD. GT. 2068 TO 47
       าลิทธศ (กลิทธิรัฐริสิทธินา 🕹 🛶 "สาก ) .
        and to the set of EMERGENTEMENT
        - MATTHIE
         . <u>=</u> = 11500 .
         E 15
```

SUBPOUTINE BOUND ON IMPRICATEMEN NMS) DIMENSION FMON-8:7% ◆ THIS ROUTINE SELECTS AND COMPUTES THE VALUES OF THE RIGURIC ◆ SPLINE FUNCTIONS AT THE OUTER PADIUS AND TOP SUPPACE OF THE ◆ PROBLEM CYLINDER. THESE VALUES ARE REQUIRED IN THE EVALUATION OF ◆ THE SUPPRICE INTEGRAL TERMS. LI=MII-MI  $\bigcup_{i=1}^{n} I_i = \bigcup_{i=1}^{n} I_i = I_i \subseteq I_i$ IF (LI.ED.O) THEM L1=1IF +LD.ED.O) THEN L2=1 60 TO 40 ELIE IF (LD.EO.1) THEN LEFE 60 TO 40 ELSE IF 'LD.EG.&' THEN L2=3 60 TO 40 EMBIE ELIE IF (LI.EG.1) THEM L1=관 IF (LD.ED.G) THEH LEEE 50 TO 40 ELSE IF (LP.EO. -1) THEN L2=1 GO TO 46 ELIE IF (LD.EG.1) THEN L2≈3 GO TO 40 ENDIF ELIE IF (LI.ED.2) THEN L1=3 IF (LD.EO.O) THEN 13 TP 46 REIGHT OF HER. BOLG. - THER LE≈E GD TB 40 ELIE IF LIV. ED. -2 THEN L2=1 EHULE FINITE The contract of the contract o FETHER

 TOTAL SECTION CHARLEST THE
 HAR IT SECTION SETTIMENTS. TERRITORISHE MATERIC PER CLAMP TO

ouganyothan agai alaar Atasaatta aa toolaa

```
PRINT++
           MATRIX CHECK FOR DIAGONAL DOMINANCE.
PRINT++ 1
PRINT++ C POW
                                          SUM OF POW ELEMENTS.
                   DIAGONAL ELEMENT
PRINT+, 4 4
DO 40 I=1.NT
TEMP=0.0
DO 30 J=1.NT
TEMP=TEMP+ABS (AS (I, U))
TEMP=TEMP+ABS(AS(I.I))
                    /*AS(I,I)*
FFINT+ *
          • I •
                                   • TEMP
PRINT++
CONTINUE
PRINT+ ·
PRINT*** MATRIX CHECK FOR SYMMETRY.
PPINT++
DO 60 I=1.NT
D□ 60 J=1•I
PRINT++C AS(C+I+C+C+J+C) = C+AS(I+J)+C
                                                ADMONDS OF SEC.
15 (J. I)
CONTINUE
RETURN
END
SUBPOUTINE SOLVE (W.SB.AR.HT.ME)
DIMENSION AS(50,50), SB(50), CA(50), ME(50)
◆ THIS POUTINE SOLVES THE PROBLEM MATRICES BY THE ITERATIVE METHOS
◆ OF SUCCESSIVE OVER-RELAMATION.
FRINT+.
PRINT++
DO 20 I=1.HT
MH \cdot I \cdot = 0.
ME \cdot I = 0.
IT = 0
IT = IT + 1
IO 50 F=1⋅NT
TEMP=IB(K)
TO 43 (=1.MT
T취성단=T취성단+AIN + 30 ★2.6 + 31
 图 (1) · 由TEME+WARE(1) + 1 (1 + 2) 1 (1) (1
DB 60 I=1:HT
IF ('P'I).E0.0.068 TD 60
TDL = (MB \cdot I) - MA \cdot I + (MB \cdot I)
IF(881)(TBL).6T..001)60 TB 70
CONTINUE
GD TO GO
 THE THE
 (É-1). ™.1900-58 TO 196
DG 80 I=1.0T
 -----
          THE TOURTED HAS COMMERCED. THE FINAL TOUR MASS FOOL
           THE NUMBER OF ITERATIONS WERE - . IT
FFINT++
           II | 1 1 1 + .
 raftyr.,
```

```
PRINT ◆ · · THE PROPLEM HAS NOT CONVERGED IN ( • IT • · ITERATIONS . )
PPINT++
PRINT+. THE EPPOP = .TOL
PPINT++
              \bullet \text{ME} \cdot \text{I} \circ \bullet \text{I} = \text{I} \bullet \text{MT}
FFINT+++
FETURN
FHI
 JUBPOUTINE TELUX:PMP:PMZ:XB:P:Z:kT:MZS:MPS:TPHI:DPHI:HB:SIGT:ASH:.
DIMENSION PMZ(-2:7) .PMR(-2:7) .XB(50) .ISPF(4.4)

    THIS POUTINE COMPUTES THE SCATTERED AND TOTAL REACTION RATE

    ▼ FLUENCE.

CALL IEMAT (ISPF.PMP.PMZ.R.Z.MPS.MZS)
12=5
TEMP1=0.
DO 40 I=1•4
MC=ICPF(I:1)
IF (M2.61.H23)68 TB 40
+ Z=+ Z-1
トドェニ
DO 30 J=1+4
MR=IIFF+J+2+
IF MR.GT.MRI GO TO 30
##F=1++ T◆++MP-1+MP(◆+M2~1))
F = F - 1
TEMP=(18 · MF) ◆SPF (1 · FC · PNC (MZ-3) · PNC (MZ-2) · PNC · MC-1) · PNC · MC(1) · C · ◆ 3FF ·
11.FF.PMP(MF=3).FMP(MP=3).PMF(MP=1).FMP(MP).P/
TEMP1=TEMP1+TEMP
BUNITHOD
CONTINUE
A=308T:16.◆ATAN:1.>>
TPHI=A+TEMP1
ZH=Z-HB
PCH=00PT (P++2+CH++2)
IF:PZH.E0.0/60 TO 80
4.5=34 F78
1801 1.80. GBD 77 80
60 TO 70
TAU=11GT+ENR+-Z ACHN+R
DPHI=YD*EMP(-TAU)
(16*ATAM)1.0*PZH**2)
60 TO 90
PRINT ◆・ FRRGR. THE FLUX CANNOT BE COMPUTED AT THE BURST POINT.
TUE
SETTEN
21:50
 INDECHION IDEAL STEEL FREE FREE FROM F. S. MEL FROM S.
 艺术的原则,美国的人员设备自身体系统,由1956年(4),10分钟的设计,10分钟
 • THEE TOURTH - CONNECTED THE CHURCH OF THE FLOWER CRUINES AT

    → [ELECTED TRACE PDINT].
```

```
DB 40 I=0:NZS-3
 IF(Z.GE.PNZ(I).AND.J.LT.PNZ(I+1)/GB TB 50
 CONTINUE
 DD 60 J=1.4
 \texttt{ISPF}(J \bullet 1) = \texttt{I} + J
 DD 70 I=0:MRS-3
 \texttt{IF} \, (\texttt{R.GE.FMR} \, (\texttt{I}) \, . \, \texttt{AND.F.LT.FMR} \, (\texttt{I+1}) \, ) \, \texttt{GD} \, \, \texttt{TD} \, \, \texttt{80}
 CONTINUE
 DO 85 J=1.4
 IGPF(J.2)=I+J
 60 TO 100
 PRINT++4 THE SPACIAL POSITION OF4,2,4 OR 4.8,4 IS NOT WITHIN THE
IPROBLEM DIMENSIONS.
 STOP
 PETUPN
 THI
```

### Appendix K

#### Numerical Results

The computer subroutines which are listed in Appendix J were used to produce numerical results for varying problem spatial mesh sizes and degrees of the spherical harmonic trial function expansion. Some of these results are presented in Figures 10 through 44. They are valid for the problem parameters which were presented in Chapter V. However, the cross-sections which were used do not accurately represent the values for air at sea level. Also, the air-ground interface was not included in the problem domain and therefore all ground effects were ignored.

Because of the time constraints on this research project neither an evaluation of the accuracy of these results nor a comparison to a discrete ordinate or Monte Carlo calculation was accomplished. Therefore, the results are presented solely in an attempt to show that finite element spaceangle synthesis is a viable solution technique for solving the two-dimensional steady state anisotropic Boltzmann equation. They are not meant to represent a precise and exact solution to the air-over-ground problem, but rather to demonstrate that FESAS may be a feasible alternate solution technique to Monte Carlo and discrete ordinates.

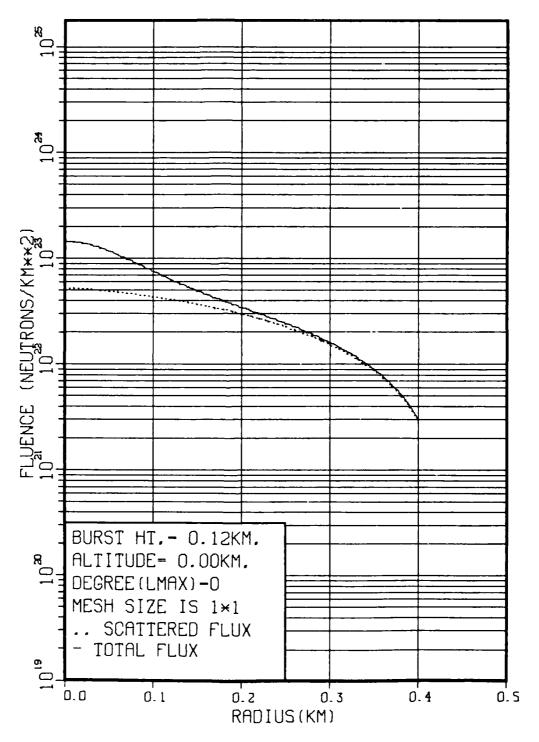


Figure 10

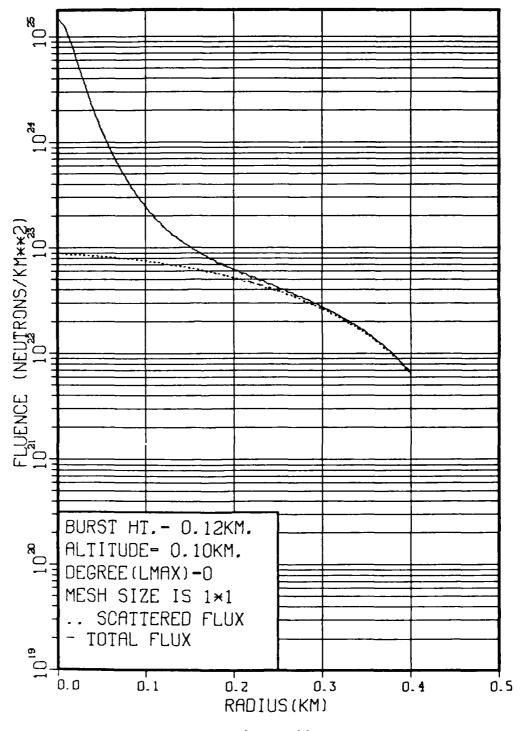


Figure 11

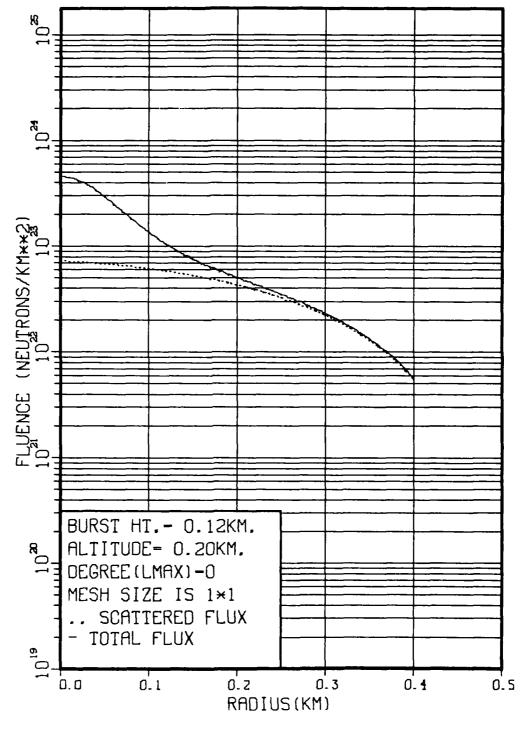


Figure 12

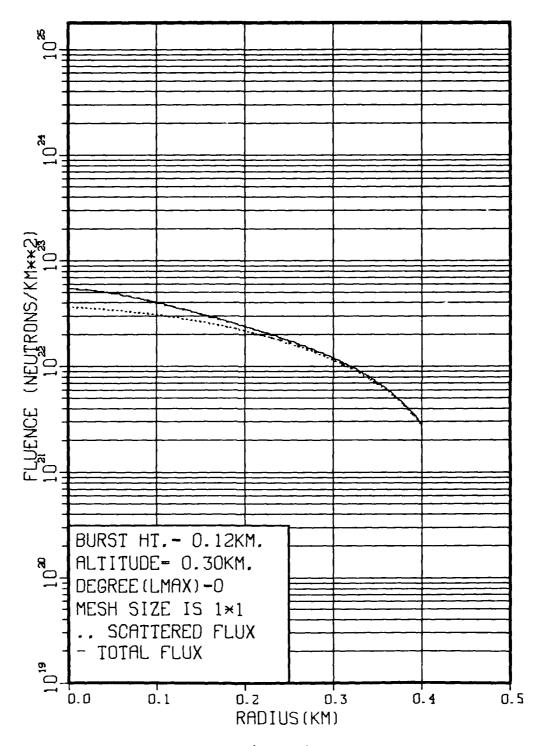


Figure 13

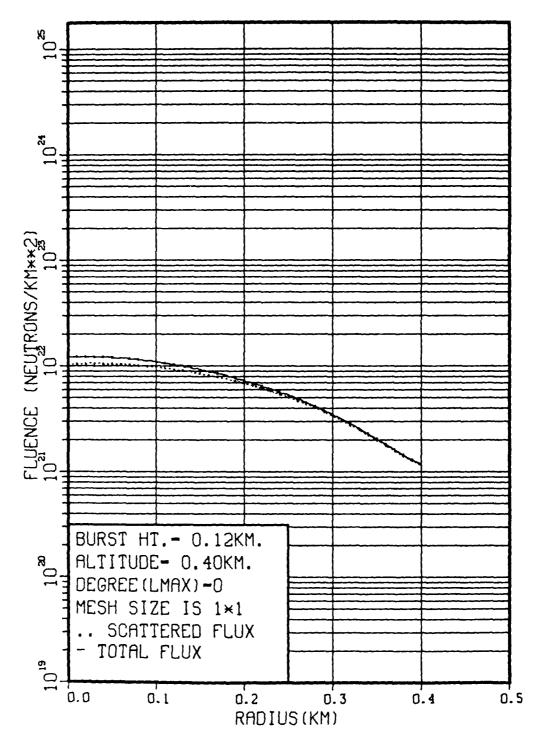


Figure 14

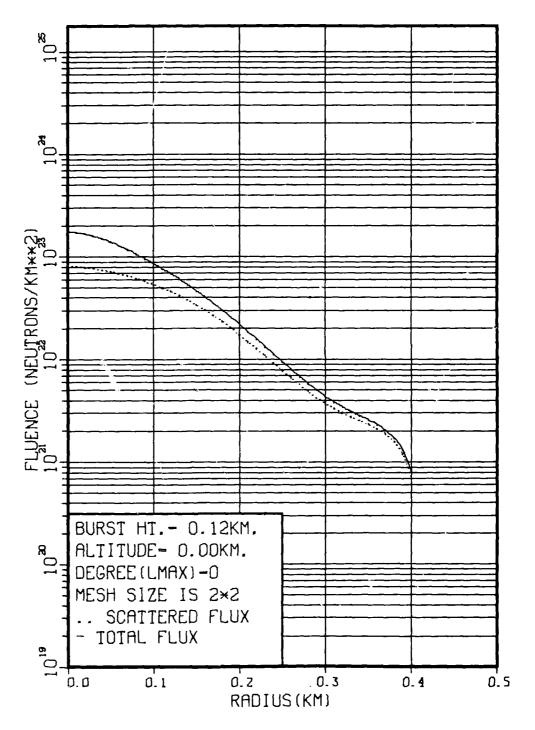


Figure 15

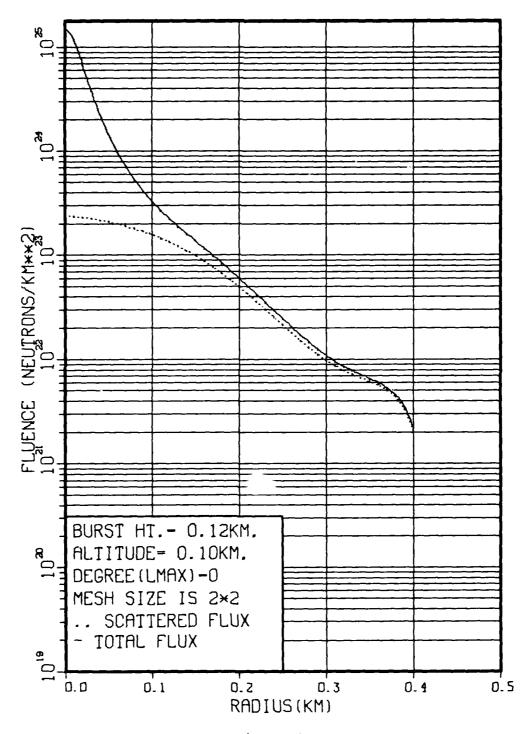


Figure 16

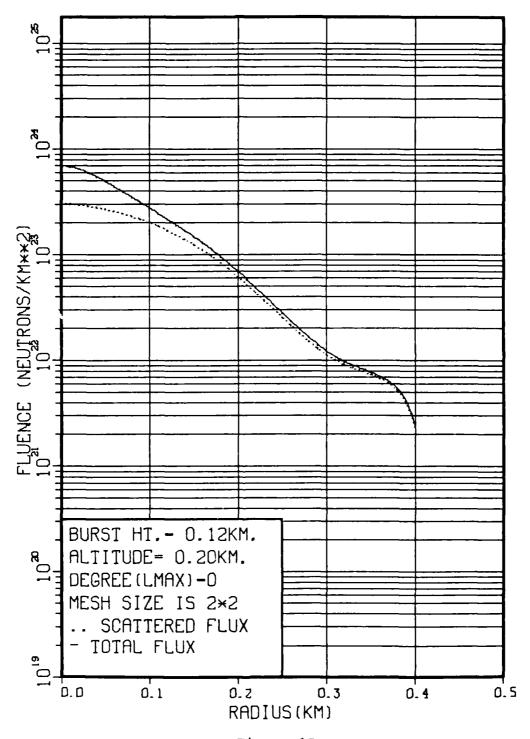


Figure 17

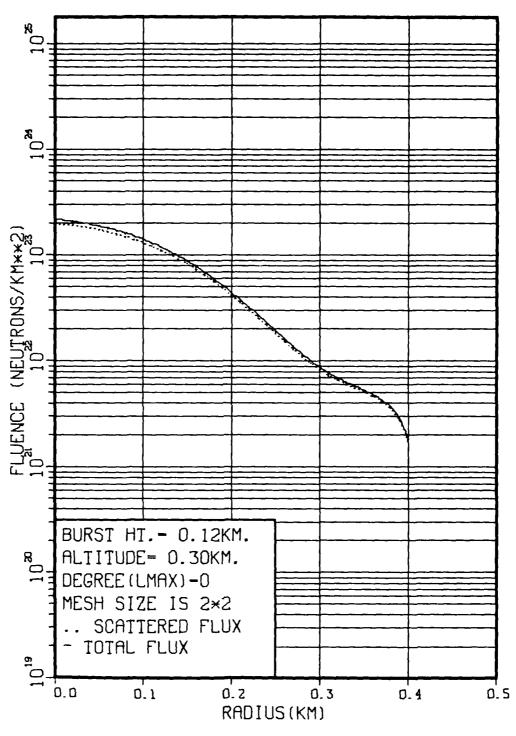


Figure 18

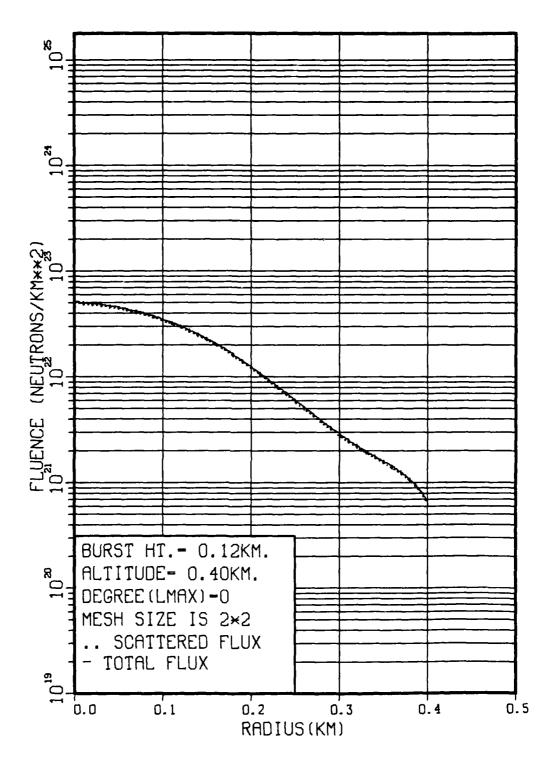
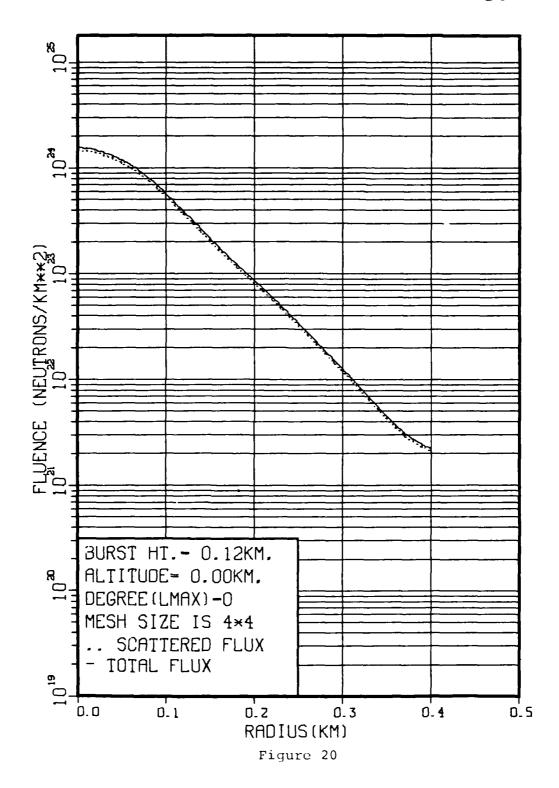


Figure 19



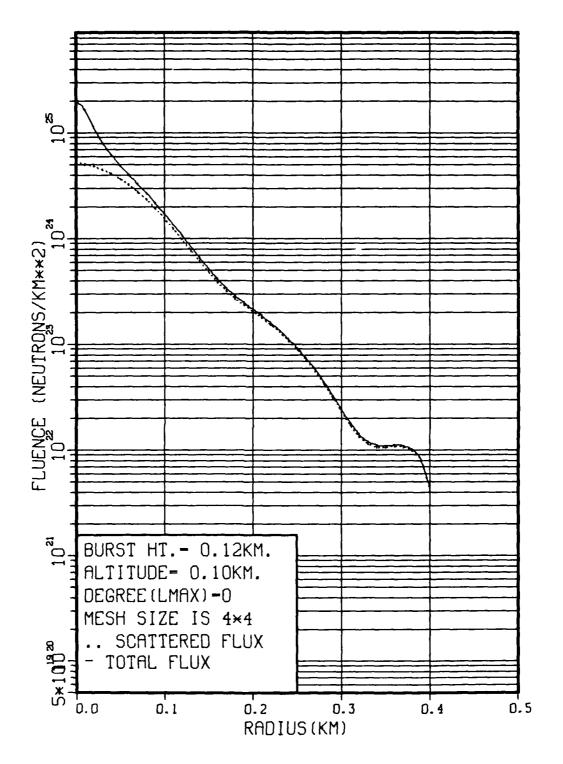


Figure 21

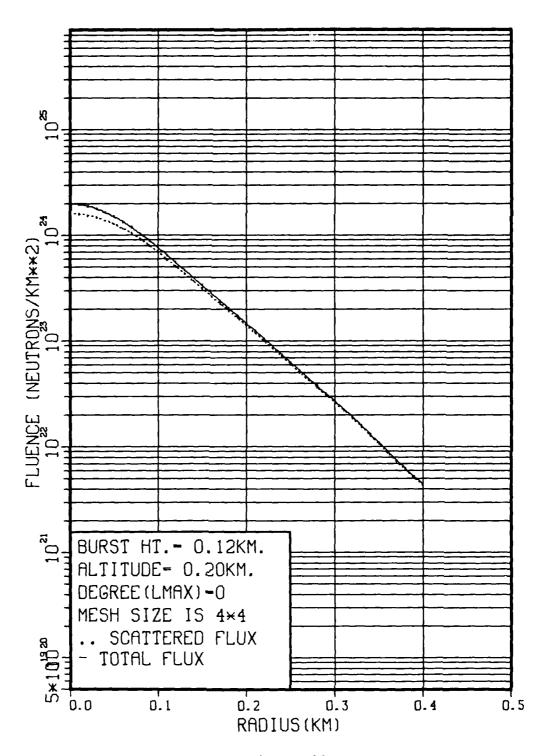


Figure 22

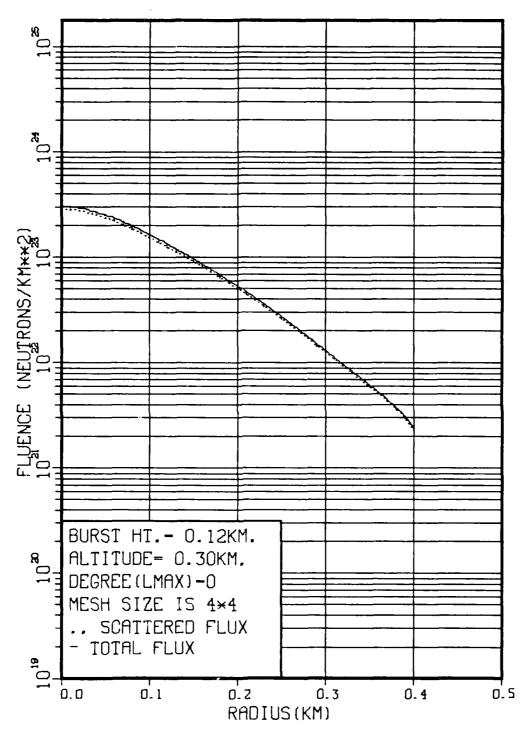
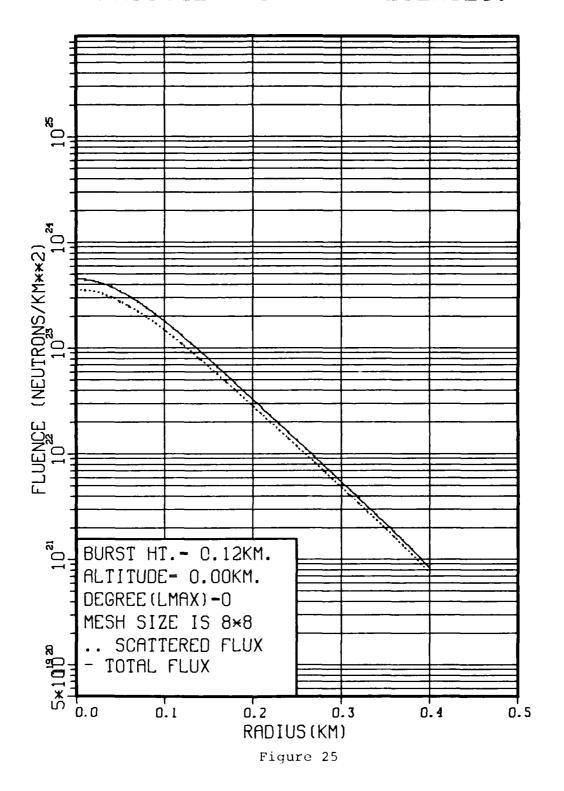


Figure 23

Figure 24



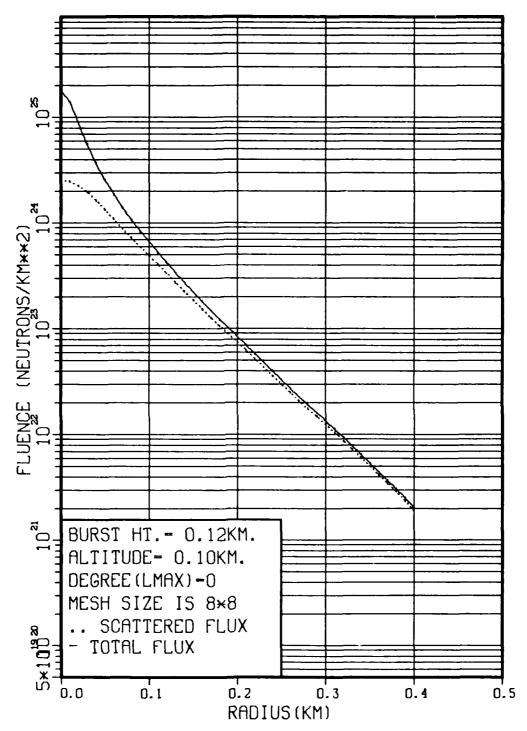


Figure 26

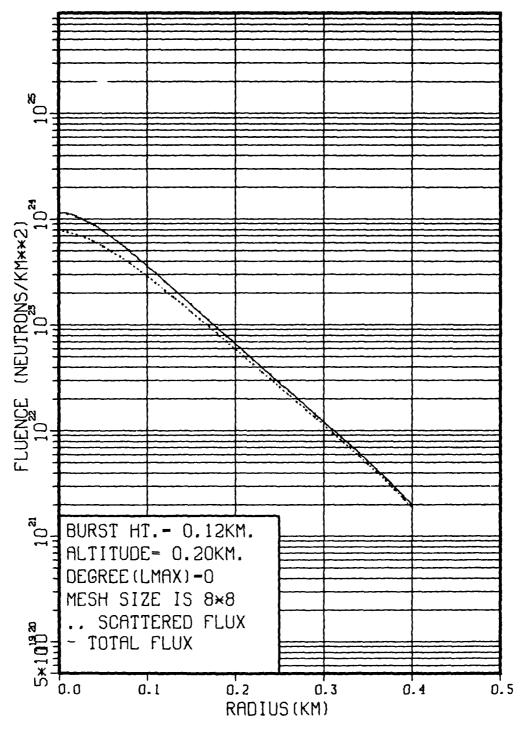


Figure 27

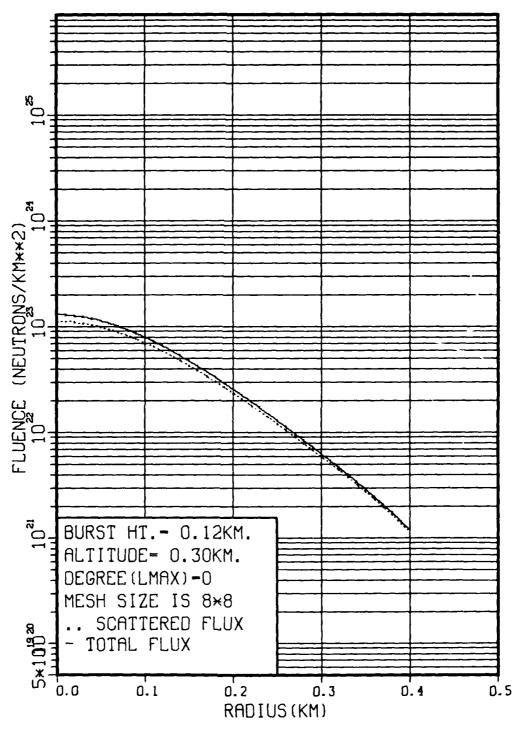


Figure 28

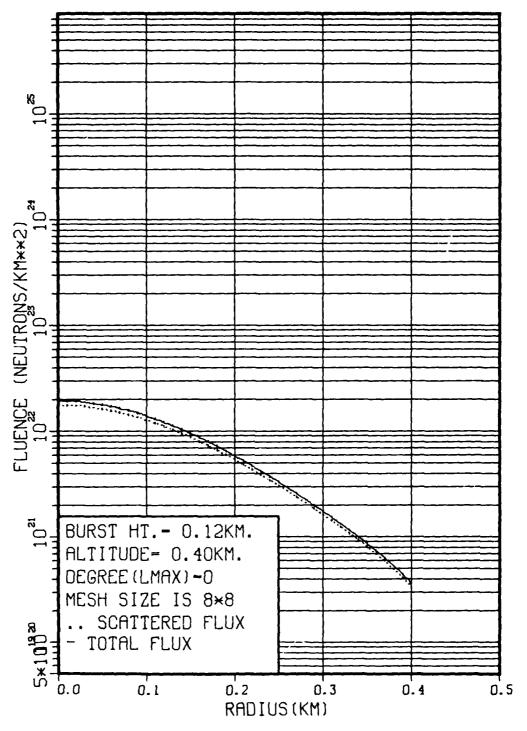
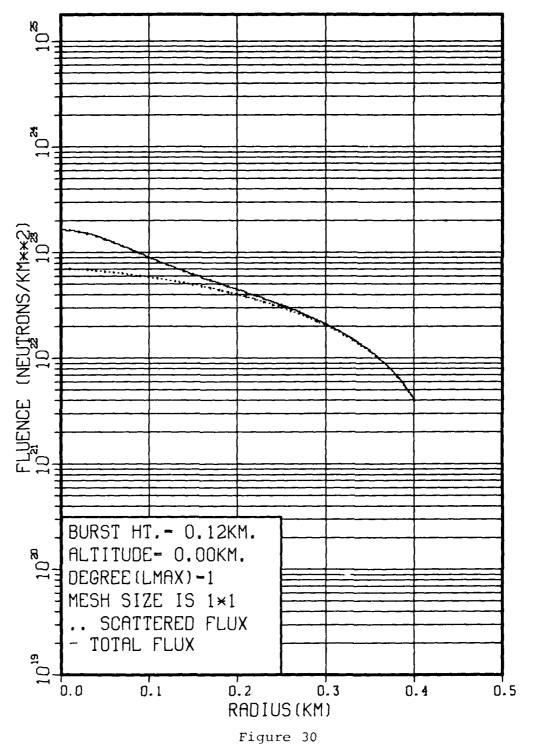
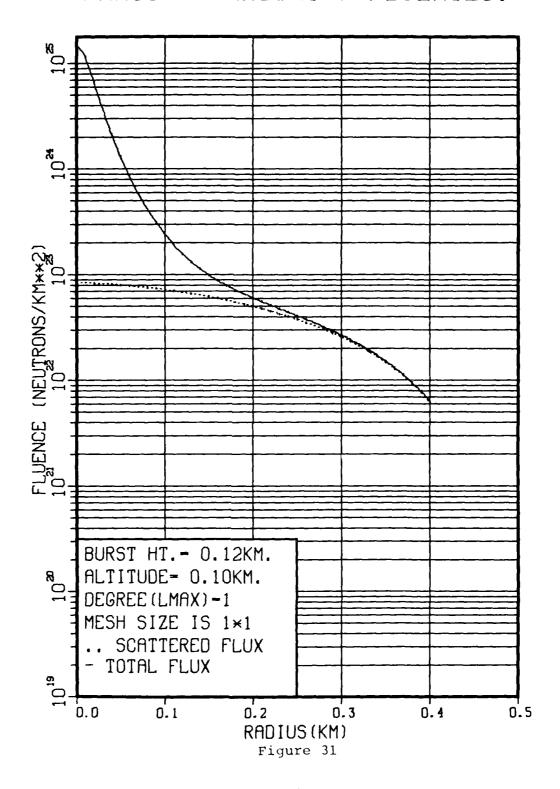
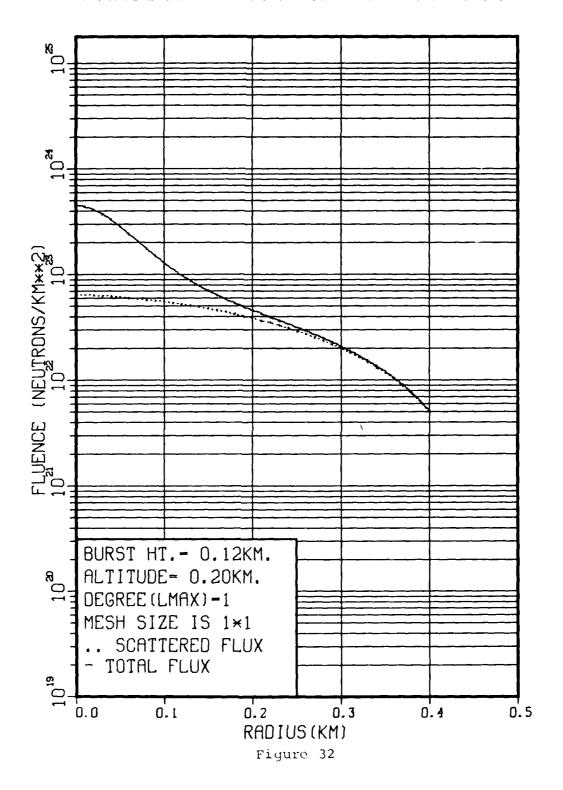


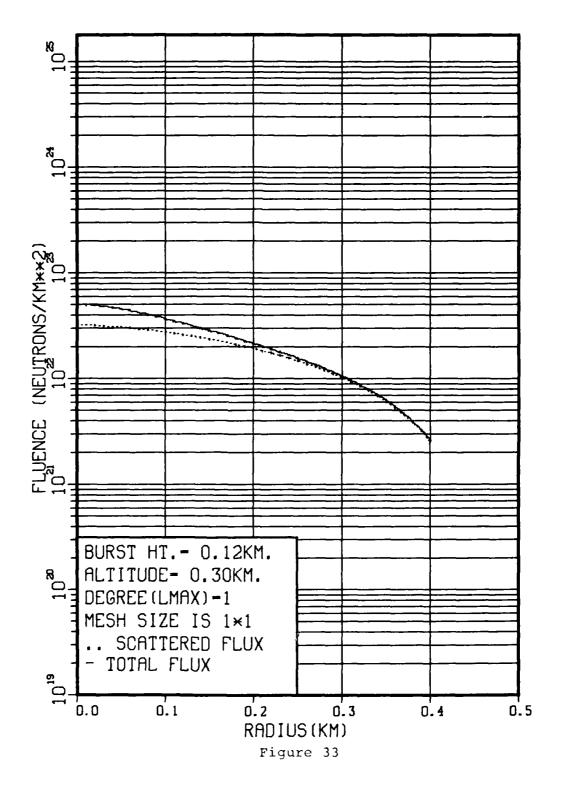
Figure 29

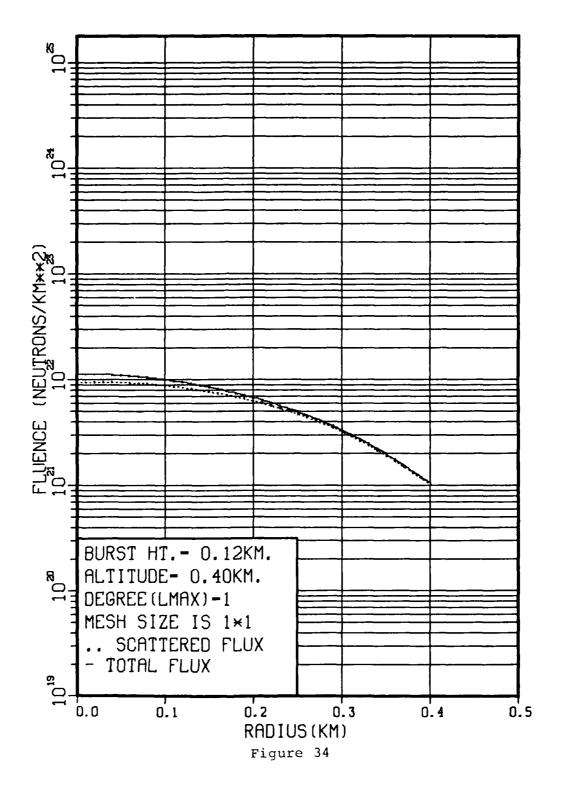


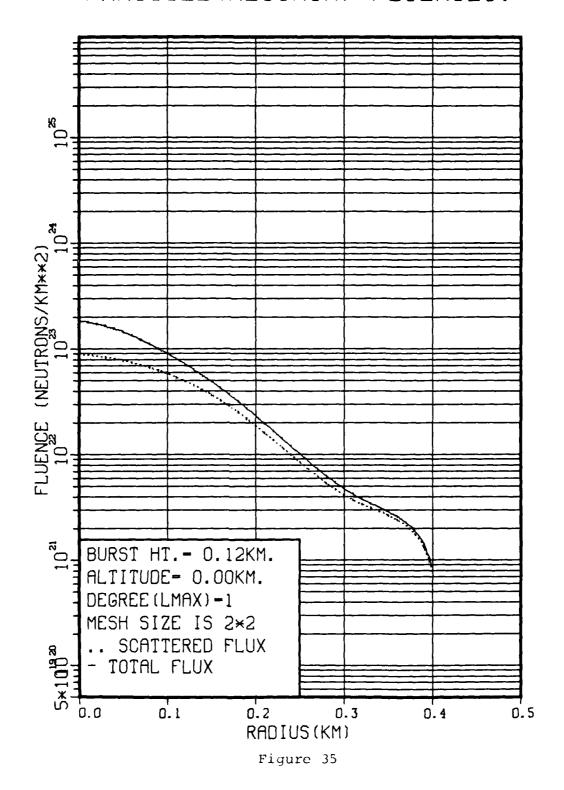
,



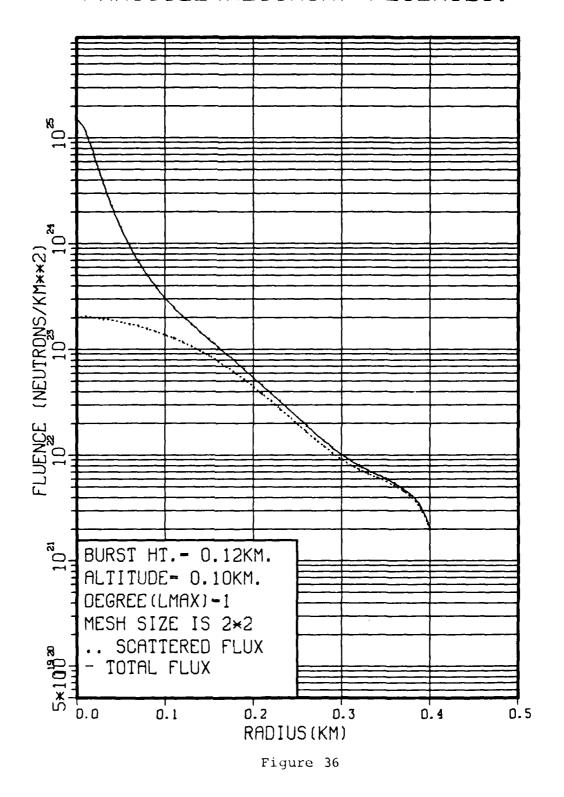








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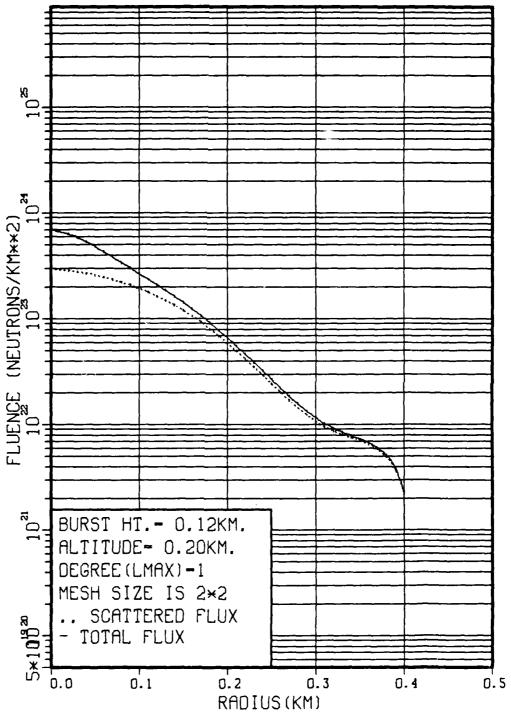
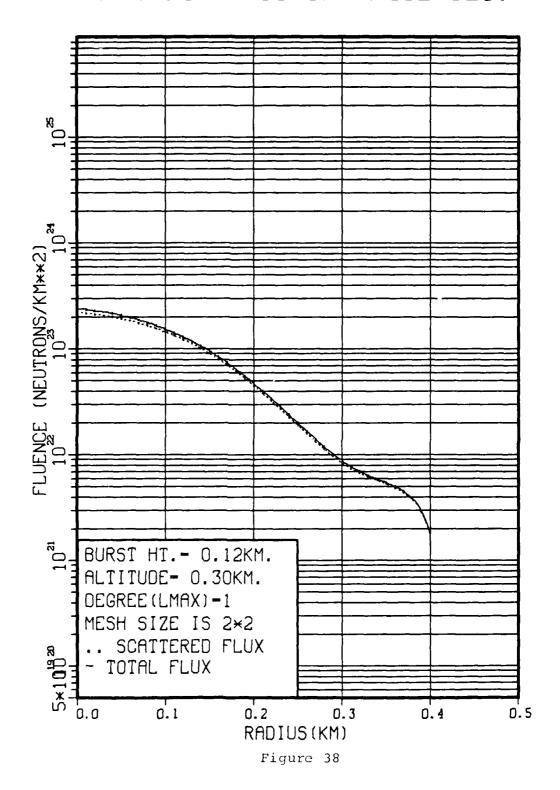


Figure 37



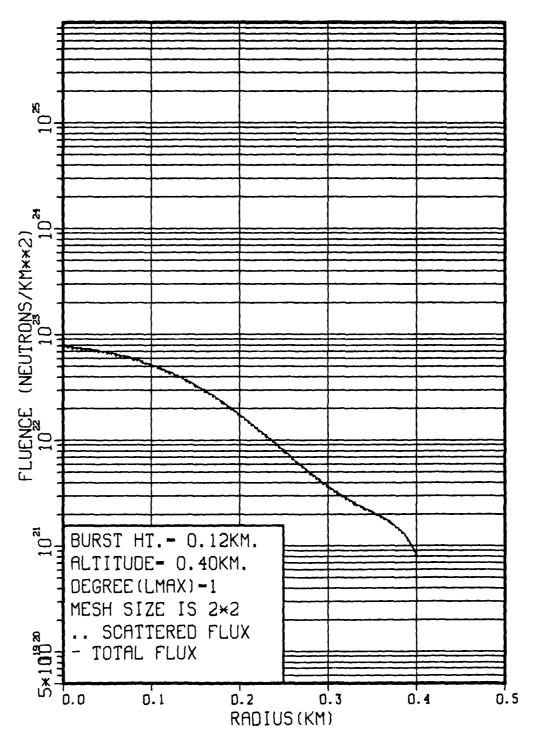
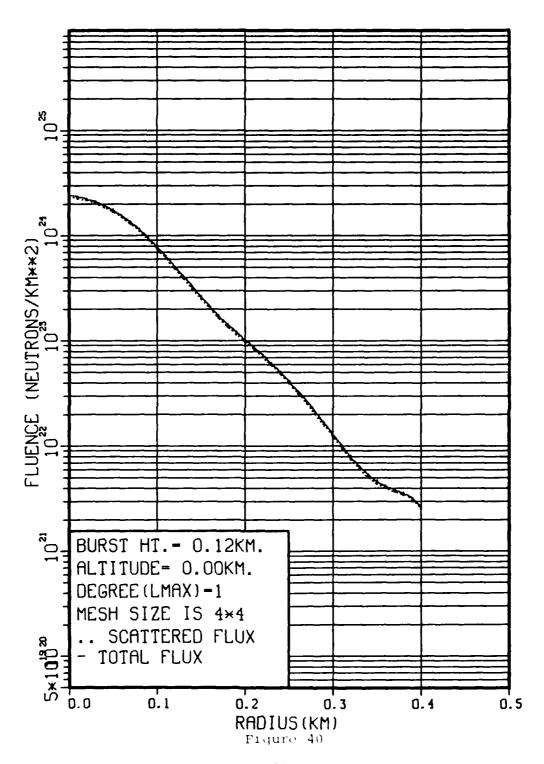
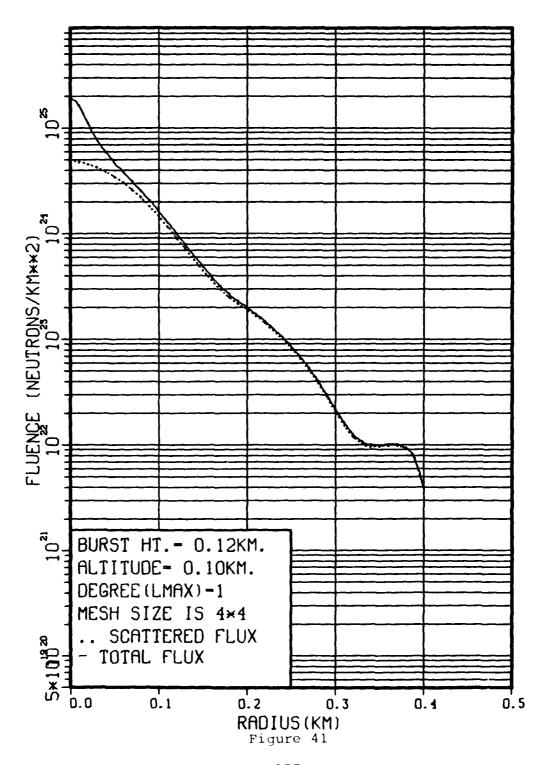
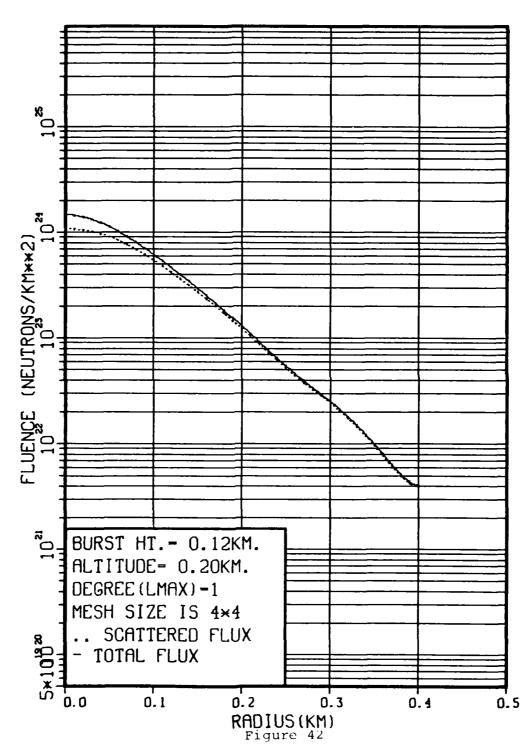
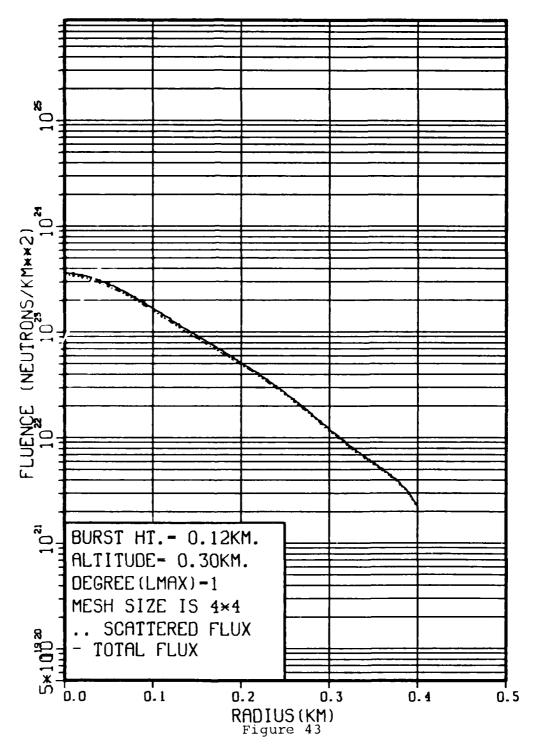


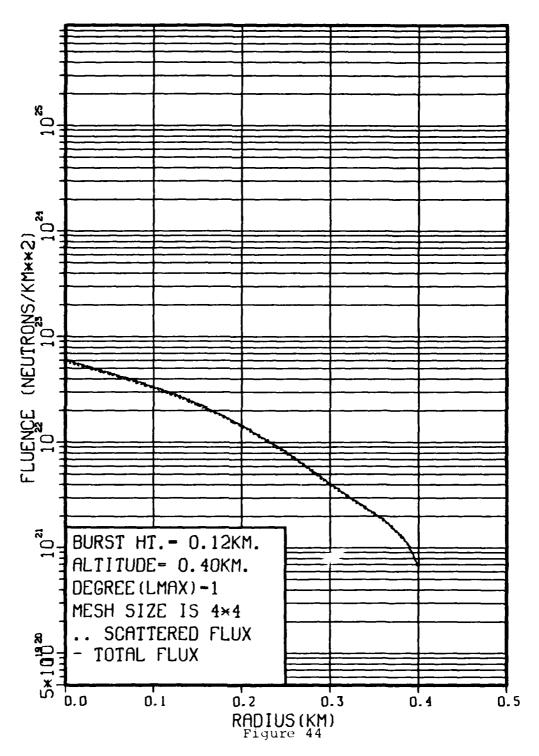
Figure 39











#### <u>Vita</u>

Eze E. Wills was born in Cuyana, South America, in April, 1951. He served as a Food Inspection Specialist in the enlisted ranks of the U.S. Army from 1972 to 1974. In 1974, he was promoted to the rank of Sergeant. Upon completion of his two year military commitment he was discharged from the Army in 1974. He then worked as a Security Guard with the Department of Navy, Navy Headquarters, Washington, D.C. from 1974 to 1976. In 1975 he enrolled at the University of Maryland, from which he graduated in December 1978 with the degree of Bachelor of Science, Nuclear-Chemical Engineering. He also completed ROTC training at the University of Maryland and was commissioned to the USAF in December 1978. He subsequently attended Graduate School at the University of Maryland and worked as a Chemical Engineer with the Department of the Navy, Naval Ordinance Station, Indian Head, Maryland. He entered the School of Engineering Air Force Institute of Technology in August 1979.

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A finite element space-angle synthesis solution of the steady state ani		
tropic Boltzmann (transport) equation in a two-dimensional cylindrical geometric		
has been developed. Starting from a variational principle, the Bubnov-Galer		
solution method was applied to the second order even parity form of the Boltzmann equation. A trial function flux expansion in bicubic splines and		
spherical (surface) harmonics was used. A first scatter (collision) source		
an exponentially varying atmosphere was also incorporated into this development		
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Finite element space-angle synthesis (FESAS) was developed as an alternate solution approach and an improvement in comparison to the methods of Monte Carlo and discrete ordinates. FESAS does not have the inherent characteristics which have produced the ray effect problem in discrete ordinates. Also, FESAS may result in lower computational costs than those of Monte Carlo and discrete ordinates.

The second order even parity form of the Boltzmann equation was derived and shown to be symmetric, positive definite and self-adjoint. The equivalence of a variational minimization principle and the Bubnov-Galerkin method of weighted residuals was established. The finite element space angle synthesi system of equations was expanded and a numerical computer solution approach was implemented. A computer program was written to solve for the trial function expansion (mixing) coefficients, and also to compute the particle flux

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